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# Hand-book on STATISTICAL DISTRIBUTIONS for experimentalists 

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## 1 Introduction

In experimental work e.g. in physics one often encounters problems where a standard statistical probability density function is applicable. It is often of great help to be able to handle these in different ways such as calculating probability contents or generating random numbers.

For these purposes there are excellent text-books in statistics e.g. the classical work of Maurice G. Kendall and Alan Stuart [1,2] or more modern text-books as [3] and others. Some books are particularly aimed at experimental physics or even specifically at particle physics $[4,5,6,7,8]$. Concerning numerical methods a valuable references worth mentioning is [9] which has been surpassed by a new edition [10]. Also hand-books, especially [11], has been of great help throughout.

However, when it comes to actual applications it often turns out to be hard to find detailed explanations in the literature ready for implementation. This work has been collected over many years in parallel with actual experimental work. In this way some material may be "historical" and sometimes be naïve and have somewhat clumsy solutions not always made in the mathematically most stringent may. We apologize for this but still hope that it will be of interest and help for people who is struggling to find methods to solve their statistical problems in making real applications and not only learning statistics as a course. Even if one has the skill and may be able to find solutions it seems worthwhile to have easy and fast access to formulæ ready for application. Similar books and reports exist e.g. $[12,13]$ but we hope the present work may compete in describing more distributions, being more complete, and including more explanations on relations given.

The material could most probably have been divided in a more logical way but we have chosen to present the distributions in alphabetic order. In this way it is more of a hand-book than a proper text-book.

After the first release the report has been modestly changed. Minor changes to correct misprints is made whenever found. In a few cases subsections and tables have been added. These alterations are described on page 182. In October 1998 the first somewhat bigger revision was made where in particular a lot of material on the non-central sampling distributions were added.

### 1.1 Random Number Generation

In modern computing Monte Carlo simulations are of vital importance and we give methods to achieve random numbers from the distributions. An earlier report dealt entirely with these matters [14]. Not all text-books on statistics include information on this subject which we find extremely useful. Large simulations are common in particle physics as well as in other areas but often it is also useful to make small "toy Monte Carlo programs" to investigate and study analysis tools developed on ideal, but statistically sound, random samples.

A related and important field which we will only mention briefly here, is how to get good basic generators for achieving random numbers uniformly distributed between zero and one. Those are the basis for all the methods described in order to get random numbers
from specific distributions in this document. For a review see e.g. [15].
From older methods often using so called multiplicative congruential method or shiftgenerators G. Marsaglia et al [16] introduced in 1989 a new "universal generator" which became the new standard in many fields. We implemented this in our experiments at CERN and also made a package of routines for general use [17].

This method is still a very good choice but later alternatives, claimed to be even better, have turned up. These are based on on the same type of lagged Fibonacci sequences as is used in the universal generator and was originally proposed by the same authors [18]. An implementations of this method was proposed by F. James [15] and this version was further developed by M. Lüscher [19]. A similar package of routine as was prepared for the universal generator has been implemented for this method [20].

## 2 Probability Density Functions

### 2.1 Introduction

Probability density functions in one, discrete or continuous, variable are denoted $p(r)$ and $f(x)$, respectively. They are assumed to be properly normalized such that

$$
\sum_{r} p(r)=1 \quad \text { and } \quad \int_{-\infty}^{\infty} f(x) d x=1
$$

where the sum or the integral are taken over all relevant values for which the probability density function is defined.

Statisticians often use the distribution function or as physicists more often call it the cumulative function which is defined as

$$
P(r)=\sum_{i=-\infty}^{r} p(i) \quad \text { and } \quad F(x)=\int_{-\infty}^{x} f(t) d t
$$

### 2.2 Moments

Algebraic moments of order $r$ are defined as the expectation value

$$
\mu_{r}^{\prime}=E\left(x^{r}\right)=\sum_{k} k^{r} p(k) \quad \text { or } \quad \int_{-\infty}^{\infty} x^{r} f(x) d x
$$

Obviously $\mu_{0}^{\prime}=1$ from the normalization condition and $\mu_{1}^{\prime}$ is equal to the mean, sometimes called the expectation value, of the distribution.

Central moments of order $r$ are defined as

$$
\mu_{r}=E\left((k-E(k))^{r}\right) \quad \text { or } \quad E\left((x-E(x))^{r}\right)
$$

of which the most commonly used is $\mu_{2}$ which is the variance of the distribution.
Instead of using the third and fourth central moments one often defines the coefficients of skewness $\gamma_{1}$ and kurtosis ${ }^{1} \gamma_{2}$ by

$$
\gamma_{1}=\frac{\mu_{3}}{\mu_{2}^{\frac{3}{2}}} \quad \text { and } \quad \gamma_{2}=\frac{\mu_{4}}{\mu_{2}^{2}}-3
$$

where the shift by 3 units in $\gamma_{2}$ assures that both measures are zero for a normal distribution. Distributions with positive kurtosis are called leptokurtic, those with kurtosis around zero mesokurtic and those with negative kurtosis platykurtic. Leptokurtic distributions are normally more peaked than the normal distribution while platykurtic distributions are more flat topped.

[^0]
### 2.2.1 Errors of Moments

For a thorough presentation of how to estimate errors on moments we refer to the classical books by M. G. Kendall and A. Stuart [1] (pp 228-245). Below only a brief description is given. For a sample with $n$ observations $x_{1}, x_{2}, \ldots, x_{n}$ we define the moment-statistics for the algebraic and central moments $m_{r}^{\prime}$ and $m_{r}$ as

$$
m_{r}^{\prime}=\frac{1}{n} \sum_{r=0}^{n} x^{r} \quad \text { and } \quad m_{r}=\frac{1}{n} \sum_{r=0}^{n}\left(x-m_{1}^{\prime}\right)^{r}
$$

The notation $m_{r}^{\prime}$ and $m_{r}$ are thus used for the statistics (sample values) while we denote the true, population, values by $\mu_{r}^{\prime}$ and $\mu_{r}$.

The mean value of the $r$ :th and the sampling covariance between the $q$ :th and $r$ :th moment-statistic are given by.

$$
\begin{aligned}
E\left(m_{r}^{\prime}\right) & =\mu_{r}^{\prime} \\
\operatorname{Cov}\left(m_{q}^{\prime}, m_{r}^{\prime}\right) & =\frac{1}{n}\left(\mu_{q+r}^{\prime}-\mu_{q}^{\prime} \mu_{r}^{\prime}\right)
\end{aligned}
$$

These formula are exact. Formulæ for moments about the mean are not as simple since the mean itself is subject to sampling fluctuations.

$$
\begin{aligned}
E\left(m_{r}\right) & =\mu_{r} \\
\operatorname{Cov}\left(m_{q}, m_{r}\right) & =\frac{1}{n}\left(\mu_{q+r}-\mu_{q} \mu_{r}+r q \mu_{2} \mu_{r-1} \mu_{q-1}-r \mu_{r-1} \mu_{q+1}-q \mu_{r+1} \mu_{q-1}\right)
\end{aligned}
$$

to order $1 / \sqrt{n}$ and $1 / n$, respectively. The covariance between an algebraic and a central moment is given by

$$
\operatorname{Cov}\left(m_{r}, m_{q}^{\prime}\right)=\frac{1}{n}\left(\mu_{q+r}-\mu_{q} \mu_{r}-r \mu_{q+1} \mu_{r-1}\right)
$$

to order $1 / n$. Note especially that

$$
\begin{aligned}
V\left(m_{r}^{\prime}\right) & =\frac{1}{n}\left(\mu_{2 r}^{\prime}-\mu_{r}^{\prime 2}\right) \\
V\left(m_{r}\right) & =\frac{1}{n}\left(\mu_{2 r}-\mu_{r}^{2}+r^{2} \mu_{2} \mu_{r-1}^{2}-2 r \mu_{r-1} \mu_{r+1}\right) \\
\operatorname{Cov}\left(m_{1}^{\prime}, m_{r}\right) & =\frac{1}{n}\left(\mu_{r+1}-r \mu_{2} \mu_{r-1}\right)
\end{aligned}
$$

### 2.3 Characteristic Function

For a distribution in a continuous variable $x$ the Fourier transform of the probability density function

$$
\phi(t)=E\left(e^{\imath x t}\right)=\int_{-\infty}^{\infty} e^{\imath x t} f(x) d x
$$

is called the characteristic function. It has the properties that $\phi(0)=1$ and $|\phi(t)| \leq 1$ for all $t$. If the cumulative, distribution, function $F(x)$ is continuous everywhere and $d F(x)=f(x) d x$ then we reverse the transform such that

$$
f(x)=\frac{1}{2 \pi} \int_{-\infty}^{\infty} \phi(t) e^{-\imath x t} d t
$$

The characteristic function is related to the moments of the distribution by

$$
\phi_{x}(t)=E\left(e^{\imath t x}\right)=\sum_{n=0}^{\infty} \frac{(\imath t)^{n} E\left(x^{n}\right)}{n!}=\sum_{n=0}^{\infty} \frac{(\imath t)^{n} \mu_{n}^{\prime}}{n!}
$$

e.g. algebraic moments may be found by

$$
\mu_{r}^{\prime}=\left.\frac{1}{v^{r}}\left(\frac{d}{d t}\right)^{r} \phi(t)\right|_{t=0}
$$

To find central moments (about the mean $\mu$ ) use

$$
\phi_{x-\mu}(t)=E\left(e^{\imath t(x-\mu)}\right)=e^{-\imath t \mu} \phi_{x}(t)
$$

and thus

$$
\mu_{r}=\left.\frac{1}{\imath^{r}}\left(\frac{d}{d t}\right)^{r} e^{-\imath t \mu} \phi(t)\right|_{t=0}
$$

A very useful property of the characteristic function is that for independent variables $x$ and $y$

$$
\phi_{x+y}(t)=\phi_{x}(t) \cdot \phi_{y}(t)
$$

As an example regard the sum $\sum a_{i} z_{i}$ where the $z_{i}$ 's are distributed according to normal distributions with means $\mu_{i}$ and variances $\sigma_{i}^{2}$. Then the linear combination will also be distributed according to the normal distribution with mean $\sum a_{i} \mu_{i}$ and variance $\sum a_{i}^{2} \sigma_{i}^{2}$.

To show that the characteristic function in two variables factorizes is the best way to show independence between two variables. Remember that a vanishing correlation coefficient does not imply independence while the reversed is true.

### 2.4 Probability Generating Function

In the case of a distribution in a discrete variable $r$ the characteristic function is given by

$$
\phi(t)=E\left(e^{t t r}\right)=\sum p(r) e^{\imath t r}
$$

In this case it is often convenient to write $z=e^{\imath t}$ and define the probability generating function as

$$
G(z)=E\left(z^{r}\right)=\sum p(r) z^{r}
$$

Derivatives of $G(z)$ evaluated at $z=1$ are related to factorial moments of the distribution

$$
\begin{aligned}
& G(1)=1 \text { (normalization) } \\
& G^{1}(1)=\left.\frac{d}{d z} G(z)\right|_{z=1}=E(r) \\
& G^{2}(1)=\left.\frac{d^{2}}{d z^{2}} G(z)\right|_{z=1}=E(r(r-1)) \\
& G^{3}(1)=\left.\frac{d^{3}}{d z^{3}} G(z)\right|_{z=1}=E(r(r-1)(r-2)) \\
& G^{k}(1)=\left.\frac{d^{k}}{d z^{k}} G(z)\right|_{z=1}=E(r(r-1)(r-2) \cdots(r-k+1))
\end{aligned}
$$

Lower order algebraic moments are then given by

$$
\begin{aligned}
& \mu_{1}^{\prime}=G^{1}(1) \\
& \mu_{2}^{\prime}=G^{2}(1)+G^{1}(1) \\
& \mu_{3}^{\prime}=G^{3}(1)+3 G^{2}(1)+G^{1}(1) \\
& \mu_{4}^{\prime}=G^{4}(1)+6 G^{3}(1)+7 G^{2}(1)+G^{1}(1)
\end{aligned}
$$

while expression for central moments become more complicated.
A useful property of the probability generating function is for a branching process in $n$ steps where

$$
G(z)=G_{1}\left(G_{2}\left(\ldots G_{n-1}\left(G_{n}(z)\right) \ldots\right)\right)
$$

with $G_{k}(z)$ the probability generating function for the distribution in the $k$ :th step. As an example see section 29.4.4 on page 105 .

### 2.5 Cumulants

Although not much used in physics the cumulants, $\kappa_{r}$, are of statistical interest. One reason for this is that they have some useful properties such as being invariant for a shift in scale (except the first cumulant which is equal to the mean and is shifted along with the scale). Multiplying the $x$-scale by a constant $a$ has the same effect as for algebraic moments namely to multiply $\kappa_{r}$ by $a^{r}$.

As the algebraic moment $\mu_{n}^{\prime}$ is the coefficient of $(\imath t)^{n} / n$ ! in the expansion of $\phi(t)$ the cumulant $\kappa_{n}$ is the coefficient of $(\imath t)^{n} / n$ ! in the expansion of the logarithm of $\phi(t)$ (sometimes called the cumulant generating function) i.e.

$$
\ln \phi(t)=\sum_{n=1}^{\infty} \frac{(\imath t)^{n}}{n!} \kappa_{n}
$$

and thus

$$
\kappa_{r}=\left.\frac{1}{\imath^{r}}\left(\frac{d}{d t}\right)^{r} \ln \phi(t)\right|_{t=0}
$$

Relations between cumulants and central moments for some lower orders are as follows

$$
\begin{array}{rlrl}
\kappa_{1}=\mu_{1}^{\prime} & & \\
\kappa_{2}=\mu_{2} & & \mu_{2}=\kappa_{2} \\
\kappa_{3}=\mu_{3} & & \mu_{3}=\kappa_{3} \\
\kappa_{4}=\mu_{4}-3 \mu_{2}^{2} & & \mu_{4}=\kappa_{4}+3 \kappa_{2}^{2} \\
\kappa_{5}=\mu_{5}-10 \mu_{3} \mu_{2} & & \mu_{5}=\kappa_{5}+10 \kappa_{3} \kappa_{2} \\
\kappa_{6}=\mu_{6}-15 \mu_{4} \mu_{2}-10 \mu_{3}^{2}+30 \mu_{2}^{3} & \mu_{6}=\kappa_{6}+15 \kappa_{4} \kappa_{2}+10 \kappa_{3}^{2}+15 \kappa_{2}^{3} \\
\kappa_{7}=\mu_{7}-21 \mu_{5} \mu_{2}-35 \mu_{4} \mu_{3}+210 \mu_{3} \mu_{2}^{2} & \mu_{7}=\kappa_{7}+21 \kappa_{5} \kappa_{2}+35 \kappa_{4} \kappa_{3}+105 \kappa_{3} \kappa_{2}^{2} \\
\kappa_{8}=\mu_{8}-28 \mu_{6} \mu_{2}-56 \mu_{5} \mu_{3}-35 \mu_{4}^{2}+ & \mu_{8}=\kappa_{8}+28 \kappa_{6} \kappa_{2}+56 \kappa_{5} \kappa_{3}+35 \kappa_{4}^{2}+ \\
& +420 \mu_{4} \mu_{2}^{2}+560 \mu_{3}^{2} \mu_{2}-630 \mu_{2}^{4} & & +210 \kappa_{4} \kappa_{2}^{2}+280 \kappa_{3}^{2} \kappa_{2}+105 \kappa_{2}^{4}
\end{array}
$$

### 2.6 Random Number Generation

When generating random numbers from different distribution it is assumed that a good generator for uniform pseudorandom numbers between zero and one exist (normally the end-points are excluded).

### 2.6.1 Cumulative Technique

The most direct technique to obtain random numbers from a continuous probability density function $f(x)$ with a limited range from $x_{\min }$ to $x_{\max }$ is to solve for $x$ in the equation

$$
\xi=\frac{F(x)-F\left(x_{\min }\right)}{F\left(x_{\max }\right)-F\left(x_{\min }\right)}
$$

where $\xi$ is uniformly distributed between zero and one and $F(x)$ is the cumulative distribution (or as statisticians say the distribution function). For a properly normalized probability density function thus

$$
x=F^{-1}(\xi)
$$

The technique is sometimes also of use in the discrete case if the cumulative sum may be expressed in analytical form as e.g. for the geometric distribution.

Also for general cases, discrete or continuous, e.g. from an arbitrary histogram the cumulative method is convenient and often faster than more elaborate methods. In this case the task is to construct a cumulative vector and assign a random number according to the value of a uniform random number (interpolating within bins in the continuous case).

### 2.6.2 Accept-Reject technique

A useful technique is the acceptance-rejection, or hit-miss, method where we choose $f_{\max }$ to be greater than or equal to $f(x)$ in the entire interval between $x_{\text {min }}$ and $x_{\text {max }}$ and proceed as follows
i Generate a pair of uniform pseudorandom numbers $\xi_{1}$ and $\xi_{2}$.
ii Determine $x=x_{\text {min }}+\xi_{1} \cdot\left(x_{\text {max }}-x_{\text {min }}\right)$.
iii Determine $y=f_{\max } \cdot \xi_{2}$.
iv If $y-f(x)>0$ reject and go to $\mathbf{i}$ else accept $x$ as a pseudorandom number from $f(x)$.

The efficiency of this method depends on the average value of $f(x) / f_{\max }$ over the interval. If this value is close to one the method is efficient. On the other hand, if this average is close to zero, the method is extremely inefficient. If $\alpha$ is the fraction of the area $f_{\max } \cdot\left(x_{\max }-x_{\min }\right)$ covered by the function the average number of rejects in step iv is $\frac{1}{\alpha}-1$ and $\frac{2}{\alpha}$ uniform pseudorandom numbers are required on average.

The efficiency of this method can be increased if we are able to choose a function $h(x)$, from which random numbers are more easily obtained, such that $f(x) \leq \alpha h(x)=g(x)$ over the entire interval under consideration (where $\alpha$ is a constant). A random sample from $f(x)$ is obtained by
i Generate in $x$ a random number from $h(x)$.
ii Generate a uniform random number $\xi$.
iii If $\xi \geq f(x) / g(x)$ go back to $\mathbf{i}$ else accept $x$ as a pseudorandom number from $f(x)$.

Yet another situation is when a function $g(x)$, from which fast generation may be obtained, can be inscribed in such a way that a big proportion $(f)$ of the area under the function is covered (as an example see the trapezoidal method for the normal distribution). Then proceed as follows:
i Generate a uniform random number $\xi$.
ii If $\xi<f$ then generate a random number from $g(x)$.
iii Else use the acceptance/rejection technique for $h(x)=f(x)-g(x)$ (in subintervals if more efficient).

### 2.6.3 Composition Techniques

If $f(x)$ may be written in the form

$$
f(x)=\int_{-\infty}^{\infty} g_{z}(x) d H(z)
$$

where we know how to sample random numbers from the p.d.f. $g(x)$ and the distribution function $H(z)$. A random number from $f(x)$ is then obtained by
i Generate two uniform random numbers $\xi_{1}$ and $\xi_{2}$.
ii Determine $z=H^{-1}\left(\xi_{1}\right)$.
iii Determine $x=G_{z}^{-1}\left(\xi_{2}\right)$ where $G_{z}$ is the distribution function corresponding to the p.d.f. $g_{z}(x)$.

For more detailed information on the Composition technique see [21] or [22].

A combination of the composition and the rejection method has been proposed by J. C. Butcher [23]. If $f(x)$ can be written

$$
f(x)=\sum_{i=0}^{n} \alpha_{i} f_{i}(x) g_{i}(x)
$$

where $\alpha_{i}$ are positive constants, $f_{i}(x)$ p.d.f.'s for which we know how to sample a random number and $g_{i}(x)$ are functions taking values between zero and one. The method is then as follows:
i Generate uniform random numbers $\xi_{1}$ and $\xi_{2}$.
ii Determine an integer $k$ from the discrete distribution $p_{i}=\alpha_{i} /\left(\alpha_{1}+\alpha_{2}+\ldots+\alpha_{n}\right)$ using $\xi_{1}$.
iii Generate a random number $x$ from $f_{k}(x)$.
iv Determine $g_{k}(x)$ and if $\xi_{2}>g_{k}(x)$ then go to $\mathbf{i}$.
$\mathbf{v}$ Accept $x$ as a random number from $f(x)$.

### 2.7 Multivariate Distributions

Joint probability density functions in several variables are denoted by $f\left(x_{1}, x_{2}, \ldots, x_{n}\right)$ and $p\left(r_{1}, r_{2}, \ldots, r_{n}\right)$ for continuous and discrete variables, respectively. It is assumed that they are properly normalized i.e. integrated (or summed) over all variables the result is unity.

### 2.7.1 Multivariate Moments

The generalization of algebraic and central moments to multivariate distributions is straightforward. As an example we take a bivariate distribution $f(x, y)$ in two continuous variables $x$ and $y$ and define algebraic and central bivariate moments of order $k, \ell$ as

$$
\begin{aligned}
\mu_{k \ell}^{\prime} \equiv E\left(x^{k} y^{\ell}\right) & =\iint x^{k} y^{\ell} f(x, y) d x d y \\
\mu_{k \ell} \equiv E\left(\left(x-\mu_{x}\right)^{k}\left(y-\mu_{y}\right)^{\ell}\right) & =\iint\left(x-\mu_{x}\right)^{k}\left(y-\mu_{y}\right)^{\ell} f(x, y) d x d y
\end{aligned}
$$

where $\mu_{x}$ and $\mu_{y}$ are the mean values of $x$ and $y$. The covariance is a central bivariate moment of order 1,1 i.e. $\operatorname{Cov}(x, y)=\mu_{11}$. Similarly one easily defines multivariate moments for distribution in discrete variables.

### 2.7.2 Errors of Bivariate Moments

Algebraic $\left(m_{r s}^{\prime}\right)$ and central $\left(m_{r s}\right)$ bivariate moments are defined by:

$$
m_{r s}^{\prime}=\frac{1}{n} \sum_{i=1}^{n} x_{i}^{r} y_{i}^{s} \quad \text { and } \quad m_{r s}=\frac{1}{n} \sum_{i=1}^{n}\left(x_{i}-m_{10}^{\prime}\right)^{r}\left(y_{i}-m_{01}^{\prime}\right)^{s}
$$

When there is a risk of ambiguity we write $m_{r, s}$ instead of $m_{r s}$.

The notations $m_{r s}^{\prime}$ and $m_{r s}$ are used for the statistics (sample values) while we write $\mu_{r s}^{\prime}$ and $\mu_{r s}$ for the population values. The errors of bivariate moments are given by

$$
\begin{aligned}
\operatorname{Cov}\left(m_{r s}^{\prime}, m_{u v}^{\prime}\right)= & \frac{1}{n}\left(\mu_{r+u, s+v}^{\prime}-\mu_{r s}^{\prime} \mu_{u v}^{\prime}\right) \\
\operatorname{Cov}\left(m_{r s}, m_{u v}\right)= & \frac{1}{n}\left(\mu_{r+u, s+v}-\mu_{r s} \mu_{u v}+r u \mu_{20} \mu_{r-1, s} \mu_{u-1, v}+s v \mu_{02} \mu_{r, s-1} \mu_{u, v-1}\right. \\
& +r v \mu_{11} \mu_{r-1, s} \mu_{u, v-1}+s u \mu_{11} \mu_{r, s-1} \mu_{u-1, v}-u \mu_{r+1, s} \mu_{u-1, v} \\
& \left.-v \mu_{r, s+1} \mu_{u, v-1}-r \mu_{r-1, s} \mu_{u+1, v}-s \mu_{r, s-1} \mu_{u, v+1}\right)
\end{aligned}
$$

especially

$$
\begin{aligned}
V\left(m_{r s}^{\prime}\right)= & \frac{1}{n}\left(\mu_{2 r, 2 s}^{\prime}-\mu_{r s}^{\prime 2}\right) \\
V\left(m_{r s}\right)= & \frac{1}{n}\left(\mu_{2 r, 2 s}-\mu_{r s}^{2}+r^{2} \mu_{20} \mu_{r-1, s}^{2}+s^{2} \mu_{02} \mu_{r, s-1}^{2}\right. \\
& \left.+2 r s \mu_{11} \mu_{r-1, s} \mu_{r, s-1}-2 r \mu_{r+1, s} \mu_{r-1, s}-2 s \mu_{r, s+1} \mu_{r, s-1}\right)
\end{aligned}
$$

For the covariance ( $m_{11}$ ) we get by error propagation

$$
\begin{aligned}
V\left(m_{11}\right) & =\frac{1}{n}\left(\mu_{22}-\mu_{11}^{2}\right) \\
\operatorname{Cov}\left(m_{11}, m_{10}^{\prime}\right) & =\frac{\mu_{21}}{n} \\
\operatorname{Cov}\left(m_{11}, m_{20}\right) & =\frac{1}{n}\left(\mu_{31}-\mu_{20} \mu_{11}\right)
\end{aligned}
$$

For the correlation coefficient (denoted by $\rho=\mu_{11} / \sqrt{\mu_{20} \mu_{02}}$ for the population value and by $r$ for the sample value) we get

$$
V(r)=\frac{\rho^{2}}{n}\left\{\frac{\mu_{22}}{\mu_{11}^{2}}+\frac{1}{4}\left[\frac{\mu_{40}}{\mu_{20}^{2}}+\frac{\mu_{04}}{\mu_{02}^{2}}+\frac{2 \mu_{22}}{\mu_{20} \mu_{02}}\right]-\frac{1}{\mu_{11}}\left[\frac{\mu_{31}}{\mu_{20}}+\frac{\mu_{13}}{\mu_{02}}\right]\right\}
$$

Beware, however, that the sampling distribution of $r$ tends to normality very slowly.

### 2.7.3 Joint Characteristic Function

The joint characteristic function is defined by

$$
\begin{aligned}
\phi\left(t_{1}, t_{2}, \ldots, t_{n}\right) & =E\left(e^{\imath t_{1} x_{1}+\imath t_{2} x_{2}+\ldots t_{n} x_{n}}\right)= \\
& =\int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \ldots \int_{-\infty}^{\infty} e^{\imath t_{1} x_{1}+\imath t_{2} x_{2}+\ldots+\imath t_{n} x_{n}} f\left(x_{1}, x_{2}, \ldots, x_{n}\right) d x_{1} d x_{2} \ldots d x_{n}
\end{aligned}
$$

From this function multivariate moments may be obtained e.g. for a bivariate distribution algebraic bivariate moments are given by

$$
\mu_{r s}^{\prime}=E\left(x_{1}^{r} x_{2}^{s}\right)=\left.\frac{\partial^{r+s} \phi\left(t_{1}, t_{2}\right)}{\partial\left(\imath t_{1}\right)^{r} \partial\left(\imath t_{2}\right)^{s}}\right|_{t_{1}=t_{2}=0}
$$

### 2.7.4 Random Number Generation

Random sampling from a many dimensional distribution with a joint probability density function $f\left(x_{1}, x_{2}, \ldots, x_{n}\right)$ can be made by the following method:

- Define the marginal distributions

$$
g_{m}\left(x_{1}, x_{2}, \ldots, x_{m}\right)=\int f\left(x_{1}, \ldots, x_{n}\right) d x_{m+1} d x_{m+2} \ldots d x_{n}=\int g_{m+1}\left(x_{1}, \ldots, x_{m+1}\right) d x_{m+1}
$$

- Consider the conditional density function $h_{m}$ given by

$$
h_{m}\left(x_{m} \mid x_{1}, x_{2}, \ldots x_{m-1}\right) \equiv g_{m}\left(x_{1}, x_{2}, \ldots, x_{m}\right) / g_{m-1}\left(x_{1}, x_{2}, \ldots, x_{m-1}\right)
$$

- We see that $g_{n}=f$ and that

$$
\int h_{m}\left(x_{m} \mid x_{1}, x_{2}, \ldots, x_{m-1}\right) d x_{m}=1
$$

from the definitions. Thus $h_{m}$ is the conditional distribution in $x_{m}$ given fixed values for $x_{1}, x_{2}, \ldots, x_{m-1}$.

- We can now factorize $f$ as

$$
f\left(x_{1}, x_{2}, \ldots, x_{n}\right)=h_{1}\left(x_{1}\right) h_{2}\left(x_{2} \mid x_{1}\right) \ldots h_{n}\left(x_{n} \mid x_{1}, x_{2}, \ldots, x_{n-1}\right)
$$

- We sample values for $x_{1}, x_{2}, \ldots, x_{n}$ from the joint probability density function $f$ by:
- Generate a value for $x_{1}$ from $h_{1}\left(x_{1}\right)$.
- Use $x_{1}$ and sample $x_{2}$ from $h_{2}\left(x_{2} \mid x_{1}\right)$.
- Proceed step by step and use previously sampled values for $x_{1}, x_{2}, \ldots, x_{m}$ to obtain a value for $x_{m+1}$ from $h_{m+1}\left(x_{m+1} \mid x_{1}, x_{2}, \ldots, x_{m}\right)$.
- Continue until all $x_{i}$ :s have been sampled.
- If all $x_{i}$ :s are independent the conditional densities will equal the marginal densities and the variables can be sampled in any order.


## 3 Bernoulli Distribution

### 3.1 Introduction

The Bernoulli distribution, named after the swiss mathematician Jacques Bernoulli (16541705), describes a probabilistic experiment where a trial has two possible outcomes, a success or a failure.

The parameter $p$ is the probability for a success in a single trial, the probability for a failure thus being $1-p$ (often denoted by $q$ ). Both $p$ and $q$ is limited to the interval from zero to one. The distribution has the simple form

$$
p(r ; p)=\left\{\begin{array}{llll}
1-p=q & \text { if } & r=0 & \text { (failure) } \\
p & \text { if } & r=1 & \text { (success) }
\end{array}\right.
$$

and zero elsewhere. The work of J. Bernoulli, which constitutes a foundation of probability theory, was published posthumously in Ars Conjectandi (1713) [24].

The probability generating function is $G(z)=q+p z$ and the distribution function given by $P(0)=q$ and $P(1)=1$. A random numbers are easily obtained by using a uniform random number variate $\xi$ and putting $r=1$ (success) if $\xi \leq p$ and $r=0$ else (failure).

### 3.2 Relation to Other Distributions

From the Bernoulli distribution we may deduce several probability density functions described in this document all of which are based on series of independent Bernoulli trials:

- Binomial distribution: expresses the probability for $r$ successes in an experiment with $n$ trials $(0 \leq r \leq n)$.
- Geometric distribution: expresses the probability of having to wait exactly $r$ trials before the first successful event $(r \geq 1)$.
- Negative Binomial distribution: expresses the probability of having to wait exactly $r$ trials until $k$ successes have occurred $(r \geq k)$. This form is sometimes referred to as the Pascal distribution.
Sometimes this distribution is expressed as the number of failures $n$ occurring while waiting for $k$ successes $(n \geq 0)$.


## 4 Beta distribution

### 4.1 Introduction

The Beta distribution is given by

$$
f(x ; p, q)=\frac{1}{B(p, q)} x^{p-1}(1-x)^{q-1}
$$

where the parameters $p$ and $q$ are positive real quantities and the variable $x$ satisfies $0 \leq$ $x \leq 1$. The quantity $B(p, q)$ is the Beta function defined in terms of the more common Gamma function as

$$
B(p, q)=\frac{\Gamma(p) \Gamma(q)}{\Gamma(p+q)}
$$

For $p=q=1$ the Beta distribution simply becomes a uniform distribution between zero and one. For $p=1$ and $q=2$ or vise versa we get triangular shaped distributions, $f(x)=2-2 x$ and $f(x)=2 x$. For $p=q=2$ we obtain a distribution of parabolic shape, $f(x)=6 x(1-x)$. More generally, if $p$ and $q$ both are greater than one the distribution has a unique mode at $x=(p-1) /(p+q-2)$ and is zero at the end-points. If $p$ and/or $q$ is less than one $f(0) \rightarrow \infty$ and/or $f(1) \rightarrow \infty$ and the distribution is said to be J-shaped. In figure 1 below we show the Beta distribution for two cases: $p=q=2$ and $p=6, q=3$.


Figure 1: Examples of Beta distributions

### 4.2 Derivation of the Beta Distribution

If $y_{m}$ and $y_{n}$ are two independent variables distributed according to the chi-squared distribution with $m$ and $n$ degrees of freedom, respectively, then the ratio $y_{m} /\left(y_{m}+y_{n}\right)$ follows a Beta distribution with parameters $p=\frac{m}{2}$ and $q=\frac{n}{2}$.

To show this we make a change of variables to $x=y_{m} /\left(y_{m}+y_{n}\right)$ and $y=y_{m}+y_{n}$ which implies that $y_{m}=x y$ and $y_{n}=y(1-x)$. We obtain

$$
\begin{aligned}
f(x, y) & =\left\lvert\, \begin{array}{cc}
\frac{\partial y_{m}}{\partial x} & \frac{\partial y_{m}}{\partial y} \\
\frac{y_{n}}{\partial x} & \frac{\partial y_{n}}{\partial y}
\end{array}\right. \| f\left(y_{m}, y_{n}\right)= \\
& =\left|\begin{array}{cc}
y & x \\
-y & 1-x
\end{array}\right|\left\{\frac{\left(\frac{y_{m}}{2}\right)^{\frac{m}{2}-1} e^{-\frac{y_{m}}{2}}}{2 \Gamma\left(\frac{m}{2}\right)}\right\}\left\{\frac{\left(\frac{y_{n}}{2}\right)^{\frac{n}{2}-1} e^{-\frac{y_{n}}{2}}}{2 \Gamma\left(\frac{n}{2}\right)}\right\}= \\
& =\left\{\frac{\Gamma\left(\frac{m+n}{2}\right)}{\Gamma\left(\frac{m}{2}\right) \Gamma\left(\frac{n}{2}\right)} x^{\frac{m}{2}-1}(1-x)^{\frac{n}{2}-1}\right\}\left\{\frac{\left(\frac{y}{2}\right)^{\frac{m}{2}+\frac{n}{2}-1} e^{-\frac{y}{2}}}{2 \Gamma\left(\frac{m+n}{2}\right)}\right\}
\end{aligned}
$$

which we recognize as a product of a Beta distribution in the variable $x$ and a chi-squared distribution with $m+n$ degrees of freedom in the variable $y$ (as expected for the sum of two independent chi-square variables).

### 4.3 Characteristic Function

The characteristic function of the Beta distribution may be expressed in terms of the confluent hypergeometric function (see section 43.3) as

$$
\phi(t)=M(p, p+q ; \imath t)
$$

### 4.4 Moments

The expectation value, variance, third and fourth central moment are given by

$$
\begin{aligned}
E(x) & =\frac{p}{p+q} \\
V(x) & =\frac{p q}{(p+q)^{2}(p+q+1)} \\
\mu_{3} & =\frac{2 p q(q-p)}{(p+q)^{3}(p+q+1)(p+q+2)} \\
\mu_{4} & =\frac{3 p q\left(2(p+q)^{2}+p q(p+q-6)\right)}{(p+q)^{4}(p+q+1)(p+q+2)(p+q+3)}
\end{aligned}
$$

More generally algebraic moments are given in terms of the Beta function by

$$
\mu_{k}^{\prime}=\frac{B(p+k, q)}{B(p, q)}
$$

### 4.5 Probability Content

In order to find the probability content for a Beta distribution we form the cumulative distribution

$$
F(x)=\frac{1}{B(p, q)} \int_{0}^{x} t^{p-1}(1-t)^{q-1} d t=\frac{B_{x}(p, q)}{B(p, q)}=I_{x}(p, q)
$$

where both $B_{x}$ and $I_{x}$ seems to be called the incomplete Beta function in the literature.
The incomplete Beta function $I_{x}$ is connected to the binomial distribution for integer values of $a$ by

$$
1-I_{x}(a, b)=I_{1-x}(b, a)=(1-x)^{a+b-1} \sum_{i=0}^{a-1}\binom{a+b-1}{i}\left(\frac{x}{1-x}\right)^{i}
$$

or expressed in the opposite direction

$$
\sum_{s=a}^{n}\binom{n}{s} p^{s}(1-p)^{n-s}=I_{p}(a, n-a+1)
$$

Also to the negative binomial distribution there is a connection by the relation

$$
\sum_{s=a}^{n}\binom{n+s-1}{s} p^{n} q^{s}=I_{q}(a, n)
$$

The incomplete Beta function is also connected to the probability content of Student's $t$-distribution and the $F$-distribution. See further section 42.7 for more information on $I_{x}$.

### 4.6 Random Number Generation

In order to obtain random numbers from a Beta distribution we first single out a few special cases.

For $p=1$ and/or $q=1$ we may easily solve the equation $F(x)=\xi$ where $F(x)$ is the cumulative function and $\xi$ a uniform random number between zero and one. In these cases

$$
\begin{aligned}
p=1 & \Rightarrow x=1-\xi^{1 / q} \\
q=1 & \Rightarrow x=\xi^{1 / p}
\end{aligned}
$$

For $p$ and $q$ half-integers we may use the relation to the chi-square distribution by forming the ratio

$$
\frac{y_{m}}{y_{m}+y_{n}}
$$

with $y_{m}$ and $y_{n}$ two independent random numbers from chi-square distributions with $m=$ $2 p$ and $n=2 q$ degrees of freedom, respectively.

Yet another way of obtaining random numbers from a Beta distribution valid when $p$ and $q$ are both integers is to take the $\ell$ :th out of $k(1 \leq \ell \leq k)$ independent uniform random numbers between zero and one (sorted in ascending order). Doing this we obtain a Beta distribution with parameters $p=\ell$ and $q=k+1-\ell$. Conversely, if we want to generate random numbers from a Beta distribution with integer parameters $p$ and $q$ we could use this technique with $\ell=p$ and $k=p+q-1$. This last technique implies that for low integer values of $p$ and $q$ simple code may be used, e.g. for $p=2$ and $q=1$ we may simply take $\max \left(\xi_{1}, \xi_{2}\right)$ i.e. the maximum of two uniform random numbers.

## 5 Binomial Distribution

### 5.1 Introduction

The Binomial distribution is given by

$$
p(r ; N, p)=\binom{N}{r} p^{r}(1-p)^{N-r}
$$

where the variable $r$ with $0 \leq r \leq N$ and the parameter $N(N>0)$ are integers and the parameter $p(0 \leq p \leq 1)$ is a real quantity.

The distribution describes the probability of exactly $r$ successes in $N$ trials if the probability of a success in a single trial is $p$ (we sometimes also use $q=1-p$, the probability for a failure, for convenience). It was first presented by Jacques Bernoulli in a work which was posthumously published [24].

### 5.2 Moments

The expectation value, variance, third and fourth moment are given by

$$
\begin{aligned}
E(r) & =N p \\
V(r) & =N p(1-p)=N p q \\
\mu_{3} & =N p(1-p)(1-2 p)=N p q(q-p) \\
\mu_{4} & =N p(1-p)[1+3 p(1-p)(N-2)]=N p q[1+3 p q(N-2)]
\end{aligned}
$$

Central moments of higher orders may be obtained by the recursive formula

$$
\mu_{r+1}=p q\left\{N r \mu_{r-1}+\frac{\partial \mu_{r}}{\partial p}\right\}
$$

starting with $\mu_{0}=1$ and $\mu_{1}=0$.
The coefficients of skewness and kurtosis are given by

$$
\gamma_{1}=\frac{q-p}{\sqrt{N p q}} \quad \text { and } \quad \gamma_{2}=\frac{1-6 p q}{N p q}
$$

### 5.3 Probability Generating Function

The probability generating function is given by

$$
G(z)=E\left(z^{r}\right)=\sum_{r=0}^{N} z^{r}\binom{N}{r} p^{r}(1-p)^{N-r}=(p z+q)^{N}
$$

and the characteristic function thus by

$$
\phi(t)=G\left(e^{\imath t}\right)=\left(q+p e^{\imath t}\right)^{N}
$$

### 5.4 Cumulative Function

For fixed $N$ and $p$ one may easily construct the cumulative function $P(r)$ by a recursive formula, see section on random numbers below.

However, an interesting and useful relation exist between $P(r)$ and the incomplete Beta function $I_{x}$ namely

$$
P(k)=\sum_{r=0}^{k} p(r ; N, p)=I_{1-p}(N-k, k+1)
$$

For further information on $I_{x}$ see section 42.7.

### 5.5 Random Number Generation

In order to achieve random numbers from a binomial distribution we may either

- Generate $N$ uniform random numbers and accumulate the number of such that are less or equal to $p$, or
- Use the cumulative technique, i.e. construct the cumulative, distribution, function and by use of this and one uniform random number obtain the required random number, or
- for larger values of $N$, say $N>100$, use an approximation to the normal distribution with mean $N p$ and variance $N p q$.
Except for very small values of $N$ and very high values of $p$ the cumulative technique is the fastest for numerical calculations. This is especially true if we proceed by constructing the cumulative vector once for all ${ }^{2}$ (as opposed to making this at each call) using the recursive formula

$$
p(i)=p(i-1) \frac{p}{q} \frac{N+1-i}{i}
$$

for $i=1,2, \ldots, N$ starting with $p(0)=q^{N}$.
However, using the relation given in the previous section with a well optimized code for the incomplete Beta function (see [10] or section 42.7) turns out to be a numerically more stable way of creating the cumulative distribution than a simple loop adding up the individual probabilities.

### 5.6 Estimation of Parameters

Experimentally the quantity $\frac{r}{N}$, the relative number of successes in $N$ trials, often is of more interest than $r$ itself. This variable has expectation $E\left(\frac{r}{N}\right)=p$ and variance $V\left(\frac{r}{N}\right)=\frac{p q}{N}$.

The estimated value for $p$ in an experiment giving $r$ successes in $N$ trials is $\hat{p}=\frac{r}{N}$.
If $p$ is unknown a unbiased estimate of the variance of a binomial distribution is given by

$$
V(r)=\frac{N}{N-1} N\left(\frac{r}{N}\right)\left(1-\frac{r}{N}\right)=\frac{N}{N-1} N \hat{p}(1-\hat{p})
$$

[^1]To find lower and upper confidence levels for $p$ we proceed as follows.

- For lower limits find a $p_{\text {low }}$ such that

$$
\sum_{r=k}^{N}\binom{N}{r} p_{\text {low }}^{r}\left(1-p_{\text {low }}\right)^{N-r}=1-\alpha
$$

or expressed in terms of the incomplete Beta function $1-I_{1-p}(N-k+1, k)=1-\alpha$

- for upper limits find a $p_{u p}$ such that

$$
\sum_{r=0}^{k}\binom{N}{r} p_{u p}^{r}\left(1-p_{u p}\right)^{N-r}=1-\alpha
$$

which is equivalent to $I_{1-p}(N-k, k+1)=1-\alpha$ i.e. $I_{p}(k+1, N-k)=\alpha$.
As an example we take an experiment with $N=10$ where a certain number of successes $0 \leq k \leq N$ have been observed. The confidence levels corresponding to $90 \%, 95 \%, 99 \%$ as well as the levels corresponding to one, two and three standard deviations for a normal distribution ( $84.13 \%, 97.72 \%$ and $99.87 \%$ probability content) are given below.

|  | Lower confidence levels |  |  |  |  |  | Upper confidence levels |  |  |  |  |  |  |
| ---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| k | $-3 \sigma$ | $99 \%$ | $-2 \sigma$ | $95 \%$ | $90 \%$ | $-\sigma$ | $\hat{p}$ | $-\sigma$ | $90 \%$ | $95 \%$ | $-2 \sigma$ | $99 \%$ | $-3 \sigma$ |
| 0 |  |  |  |  |  |  | 0.00 | 0.17 | 0.21 | 0.26 | 0.31 | 0.37 | 0.48 |
| 1 | 0.00 | 0.00 | 0.00 | 0.01 | 0.01 | 0.02 | 0.10 | 0.29 | 0.34 | 0.39 | 0.45 | 0.50 | 0.61 |
| 2 | 0.01 | 0.02 | 0.02 | 0.04 | 0.05 | 0.07 | 0.20 | 0.41 | 0.45 | 0.51 | 0.56 | 0.61 | 0.71 |
| 3 | 0.02 | 0.05 | 0.06 | 0.09 | 0.12 | 0.14 | 0.30 | 0.51 | 0.55 | 0.61 | 0.66 | 0.70 | 0.79 |
| 4 | 0.05 | 0.09 | 0.12 | 0.15 | 0.19 | 0.22 | 0.40 | 0.60 | 0.65 | 0.70 | 0.74 | 0.78 | 0.85 |
| 5 | 0.10 | 0.15 | 0.18 | 0.22 | 0.27 | 0.30 | 0.50 | 0.70 | 0.73 | 0.78 | 0.82 | 0.85 | 0.90 |
| 6 | 0.15 | 0.22 | 0.26 | 0.30 | 0.35 | 0.40 | 0.60 | 0.78 | 0.81 | 0.85 | 0.88 | 0.91 | 0.95 |
| 7 | 0.21 | 0.30 | 0.34 | 0.39 | 0.45 | 0.49 | 0.70 | 0.86 | 0.88 | 0.91 | 0.94 | 0.95 | 0.98 |
| 8 | 0.29 | 0.39 | 0.44 | 0.49 | 0.55 | 0.59 | 0.80 | 0.93 | 0.95 | 0.96 | 0.98 | 0.98 | 0.99 |
| 9 | 0.39 | 0.50 | 0.55 | 0.61 | 0.66 | 0.71 | 0.90 | 0.98 | 0.99 | 0.99 | 1.00 | 1.00 | 1.00 |
| 10 | 0.52 | 0.63 | 0.69 | 0.74 | 0.79 | 0.83 | 1.00 |  |  |  |  |  |  |

### 5.7 Probability Content

It is sometimes of interest to judge the significance level of a certain outcome given the hypothesis that $p=\frac{1}{2}$. If $N$ trials are made and we find $k$ successes (let's say $k<N / 2$ else use $N-k$ instead of $k$ ) we want to estimate the probability to have $k$ or fewer successes plus the probability for $N-k$ or more successes. Since the assumption is that $p=\frac{1}{2}$ we want the two-tailed probability content.

To calculate this either sum the individual probabilities or use the relation to the incomplete beta function. The former may seem more straightforward but the latter may be computationally easier given a routine for the incomplete beta function. If $k=N / 2$ we watch up not to add the central term twice (in this case the requested probability is $100 \%$ anyway). In the table below we show such confidence levels in \% for values of $N$ ranging from 1 to 20. E.g. the probability to observe 3 successes (or failures) or less and 12 failures (or successes) or more for $n=15$ is $3.52 \%$.

|  | $k$ |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $N$ | 0 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | 10 |
| 1 | 100.00 |  |  |  |  |  |  |  |  |  |  |
| 2 | 50.00 | 100.00 |  |  |  |  |  |  |  |  |  |
| 3 | 25.00 | 100.00 |  |  |  |  |  |  |  |  |  |
| 4 | 12.50 | 62.50 | 100.00 |  |  |  |  |  |  |  |  |
| 5 | 6.25 | 37.50 | 100.00 |  |  |  |  |  |  |  |  |
| 6 | 3.13 | 21.88 | 68.75 | 100.00 |  |  |  |  |  |  |  |
| 7 | 1.56 | 12.50 | 45.31 | 100.00 |  |  |  |  |  |  |  |
| 8 | 0.78 | 7.03 | 28.91 | 72.66 | 100.00 |  |  |  |  |  |  |
| 9 | 0.39 | 3.91 | 17.97 | 50.78 | 100.00 |  |  |  |  |  |  |
| 10 | 0.20 | 2.15 | 10.94 | 34.38 | 75.39 | 100.00 |  |  |  |  |  |
| 11 | 0.10 | 1.17 | 6.54 | 22.66 | 54.88 | 100.00 |  |  |  |  |  |
| 12 | 0.05 | 0.63 | 3.86 | 14.60 | 38.77 | 77.44 | 100.00 |  |  |  |  |
| 13 | 0.02 | 0.34 | 2.25 | 9.23 | 26.68 | 58.11 | 100.00 |  |  |  |  |
| 14 | 0.01 | 0.18 | 1.29 | 5.74 | 17.96 | 42.40 | 79.05 | 100.00 |  |  |  |
| 15 | 0.01 | 0.10 | 0.74 | 3.52 | 11.85 | 30.18 | 60.72 | 100.00 |  |  |  |
| 16 | 0.00 | 0.05 | 0.42 | 2.13 | 7.68 | 21.01 | 45.45 | 80.36 | 100.00 |  |  |
| 17 | 0.00 | 0.03 | 0.23 | 1.27 | 4.90 | 14.35 | 33.23 | 62.91 | 100.00 |  |  |
| 18 | 0.00 | 0.01 | 0.13 | 0.75 | 3.09 | 9.63 | 23.79 | 48.07 | 81.45 | 100.00 |  |
| 19 | 0.00 | 0.01 | 0.07 | 0.44 | 1.92 | 6.36 | 16.71 | 35.93 | 64.76 | 100.00 |  |
| 20 | 0.00 | 0.00 | 0.04 | 0.26 | 1.18 | 4.14 | 11.53 | 26.32 | 50.34 | 82.38 | 100.00 |

## 6 Binormal Distribution

### 6.1 Introduction

As a generalization of the normal or Gauss distribution to two dimensions we define the binormal distribution as

$$
f\left(x_{1}, x_{2}\right)=\frac{1}{2 \pi \sigma_{1} \sigma_{2} \sqrt{1-\rho^{2}}} \cdot e^{-\frac{1}{2\left(1-\rho^{2}\right)}\left(\left(\frac{x_{1}-\mu_{1}}{\sigma_{1}}\right)^{2}+\left(\frac{x_{2}-\mu_{2}}{\sigma_{2}}\right)^{2}-2 \rho \cdot \frac{x_{1}-\mu_{1} \cdot \frac{x_{2}-\mu_{2}}{\sigma_{2}}}{\sigma_{2}}\right)}
$$

where $\mu_{1}$ and $\mu_{2}$ are the expectation values of $x_{1}$ and $x_{2}, \sigma_{1}$ and $\sigma_{2}$ their standard deviations and $\rho$ the correlation coefficient between them. Putting $\rho=0$ we see that the distribution becomes the product of two one-dimensional Gauss distributions.


Figure 2: Binormal distribution
In figure 2 we show contours for a standardized Binormal distribution i.e putting $\mu_{1}=$ $\mu_{2}=0$ and $\sigma_{1}=\sigma_{2}=1$ (these parameters are anyway shift- and scale-parameters only). In the example shown $\rho=0.5$. Using standardized variables the contours range from a perfect circle for $\rho=0$ to gradually thinner ellipses in the $\pm 45^{\circ}$ direction as $\rho \rightarrow \pm 1$. The contours shown correspond to the one, two, and three standard deviation levels. See section on probability content below for details.

### 6.2 Conditional Probability Density

The conditional density of the binormal distribution is given by

$$
\begin{aligned}
f(x \mid y) & =f(x, y) / f(y)= \\
& =\frac{1}{\sqrt{2 \pi} \sigma_{x} \sqrt{1-\rho^{2}}} \exp \left\{-\frac{1}{2 \sigma_{x}^{2}\left(1-\rho^{2}\right)}\left[x-\left(\mu_{x}+\frac{\rho \sigma_{x}}{\sigma_{y}}\left(y-\mu_{y}\right)\right)\right]^{2}\right\}= \\
& =N\left(\mu_{x}+\rho \frac{\sigma_{x}}{\sigma_{y}}\left(y-\mu_{y}\right), \sigma_{x}^{2}\left(1-\rho^{2}\right)\right)
\end{aligned}
$$

which is seen to be a normal distribution which for $\rho=0$ is, as expected, given by $N\left(\mu_{x}, \sigma_{x}^{2}\right)$ but generally has a mean shifted from $\mu_{x}$ and a variance which is smaller than $\sigma_{x}^{2}$.

### 6.3 Characteristic Function

The characteristic function of the binormal distribution is given by

$$
\begin{aligned}
\phi\left(t_{1}, t_{2}\right) & =E\left(e^{\imath t_{1} x_{1}+\imath t_{2} x_{2}}\right)=\int_{-\infty}^{\infty} \int_{-\infty}^{\infty} e^{\imath t_{1} x_{1}+\imath t_{2} x_{2}} f\left(x_{1}, x_{2}\right) d x_{1} d x_{2}= \\
& =\exp \left\{\imath t_{1} \mu_{1}+\imath t_{2} \mu_{2}+\frac{1}{2}\left[\left(\imath t_{1}\right)^{2} \sigma_{1}^{2}+\left(\imath t_{2}\right)^{2} \sigma_{2}^{2}+2\left(\imath t_{1}\right)\left(\imath t_{2}\right) \rho \sigma_{1} \sigma_{2}\right]\right\}
\end{aligned}
$$

which shows that if the correlation coefficient $\rho$ is zero then the characteristic function factorizes i.e. the variables are independent. This is a unique property of the normal distribution since in general $\rho=0$ does not imply independence.

### 6.4 Moments

To find bivariate moments of the binormal distribution the simplest, but still quite tedious, way is to use the characteristic function given above (see section 2.7.3).

Algebraic bivariate moments for the binormal distribution becomes somewhat complicated but normally they are of less interest than the central ones. Algebraic moments of the type $\mu_{0 k}^{\prime}$ and $\mu_{k 0}^{\prime}$ are, of course, equal to moments of the marginal one-dimensional normal distribution e.g. $\mu_{10}^{\prime}=\mu_{1}, \mu_{20}^{\prime}=\mu_{1}^{2}+\sigma_{1}^{2}$, and $\mu_{30}^{\prime}=\mu_{1}\left(2 \sigma_{1}^{2}+\mu_{1}^{2}\right)$ (for $\mu_{0 k}^{\prime}$ simply exchange the subscripts on $\mu$ and $\sigma$ ). Some other lower order algebraic bivariate moments are given by

$$
\begin{aligned}
& \mu_{11}^{\prime}=\mu_{1} \mu_{2}+\rho \sigma_{1} \sigma_{2} \\
& \mu_{12}^{\prime}=2 \rho \sigma_{1} \sigma_{2} \mu_{2}+\sigma_{2}^{2} \mu_{1}+\mu_{2}^{2} \mu_{1} \\
& \mu_{22}^{\prime}=\sigma_{1}^{2} \sigma_{2}^{2}+\sigma_{1}^{2} \mu_{2}^{2}+\sigma_{2}^{2} \mu_{1}^{2}+\mu_{1}^{2} \mu_{2}^{2}+2 \rho^{2} \sigma_{1}^{2} \sigma_{2}^{2}+4 \rho \sigma_{1} \sigma_{2} \mu_{1} \mu_{2}
\end{aligned}
$$

Beware of the somewhat confusing notation where $\mu$ with two subscripts denotes bivariate moments while $\mu$ with one subscript denotes expectation values.

Lower order central bivariate moments $\mu_{k \ell}$, arranged in matrix form, are given by

|  | $\ell=0$ | $\ell=1$ | $\ell=2$ | $\ell=3$ | $\ell=4$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $k=0$ | 1 | 0 | $\sigma_{2}^{2}$ | 0 | $3 \sigma_{2}^{4}$ |
| $k=1$ | 0 | $\rho \sigma_{1} \sigma_{2}$ | 0 | $3 \rho \sigma_{1} \sigma_{2}^{3}$ | 0 |
| $k=2$ | $\sigma_{1}^{2}$ | 0 | $\sigma_{1}^{2} \sigma_{2}^{2}\left(2 \rho^{2}+1\right)$ | 0 | $3 \sigma_{1}^{2} \sigma_{2}^{4}\left(4 \rho^{2}+1\right)$ |
| $k=3$ | 0 | $3 \rho \sigma_{1}^{3} \sigma_{2}$ | 0 | $3 \rho \sigma_{1}^{3} \sigma_{2}^{3}\left(2 \rho^{2}+3\right)$ | 0 |
| $k=4$ | $3 \sigma_{1}^{4}$ | 0 | $3 \sigma_{1}^{4} \sigma_{2}^{2}\left(4 \rho^{2}+1\right)$ | 0 | $3 \sigma_{1}^{4} \sigma_{2}^{4}\left(8 \rho^{4}+24 \rho^{2}+3\right)$ |

### 6.5 Box-Muller Transformation

Recall that if we have a distribution in one set of variables $\left\{x_{1}, x_{2}, \ldots, x_{n}\right\}$ and want to change variables to another set $\left\{y_{1}, y_{2}, \ldots, y_{n}\right\}$ the distribution in the new variables are given by

$$
f\left(y_{1}, y_{2}, \ldots, y_{n}\right)=\left\|\begin{array}{cccc}
\frac{\partial x_{1}}{\partial y_{1}} & \frac{\partial x_{1}}{\partial y_{2}} & \cdots & \frac{\partial x_{1}}{\partial y_{n}} \\
\frac{\partial x_{2}}{\partial y_{1}} & \frac{\partial x_{2}}{\partial y_{2}} & \cdots & \frac{\partial x_{2}}{\partial y_{n}} \\
\vdots & \vdots & \ddots & \vdots \\
\frac{\partial x_{n}}{\partial y_{1}} & \frac{\partial x_{n}}{\partial y_{2}} & \cdots & \frac{\partial x_{n}}{\partial y_{n}}
\end{array}\right\| f\left(x_{1}, x_{2}, \ldots, x_{n}\right)
$$

where the symbol $\|J\|$ denotes the absolute value of the determinant of the Jacobian $J$.
Let $x_{1}$ and $x_{2}$ be two independent stochastic variables from a uniform distribution between zero and one and define

$$
\begin{aligned}
& y_{1}=\sqrt{-2 \ln x_{1}} \sin 2 \pi x_{2} \\
& y_{2}=\sqrt{-2 \ln x_{1}} \cos 2 \pi x_{2}
\end{aligned}
$$

Note that with the definition above $-\infty<y_{1}<\infty$ and $-\infty<y_{2}<\infty$. In order to obtain the joint probability density function in $y_{1}$ and $y_{2}$ we need to calculate the Jacobian matrix

$$
\frac{\partial\left(x_{1}, x_{2}\right)}{\partial\left(y_{1}, y_{2}\right)}=\left(\begin{array}{ll}
\frac{\partial x_{1}}{\partial y_{1}} & \frac{\partial x_{1}}{\partial y_{2}} \\
\frac{\partial x_{2}}{\partial y_{1}} & \frac{\partial x_{2}}{\partial y_{2}}
\end{array}\right)
$$

In order to obtain these partial derivatives we express $x_{1}$ and $x_{2}$ in $y_{1}$ and $y_{2}$ by rewriting the original equations.

$$
\begin{aligned}
y_{1}^{2}+y_{2}^{2} & =-2 \ln x_{1} \\
\frac{y_{1}}{y_{2}} & =\tan 2 \pi x_{2}
\end{aligned}
$$

which implies

$$
\begin{aligned}
& x_{1}=e^{-\frac{1}{2}\left(y_{1}^{2}+y_{2}^{2}\right)} \\
& x_{2}=\frac{1}{2 \pi} \arctan \left(\frac{y_{1}}{y_{2}}\right)
\end{aligned}
$$

Then the Jacobian matrix becomes

$$
\frac{\partial\left(x_{1}, x_{2}\right)}{\partial\left(y_{1}, y_{2}\right)}=\left(\begin{array}{cc}
-y_{1} e^{-\frac{1}{2}\left(y_{1}^{2}+y_{2}^{2}\right)} & -y_{2} e^{-\frac{1}{2}\left(y_{1}^{2}+y_{2}^{2}\right)} \\
\frac{1}{2 \pi y_{2}} \cos ^{2} \arctan \left(\frac{y_{1}}{y_{2}}\right) & -\frac{y_{1}}{2 \pi y_{2}^{2}} \cos ^{2} \arctan \left(\frac{y_{1}}{y_{2}}\right)
\end{array}\right)
$$

The distribution $f\left(y_{1}, y_{2}\right)$ is given by

$$
f\left(y_{1}, y_{2}\right)=\left\|\frac{\partial\left(x_{1}, x_{2}\right)}{\partial\left(y_{1}, y_{2}\right)}\right\| f\left(x_{1}, x_{2}\right)
$$

where $f\left(x_{1}, x_{2}\right)$ is the uniform distribution in $x_{1}$ and $x_{2}$. Now $f\left(x_{1}, x_{2}\right)=1$ in the interval $0 \leq x_{1} \leq 1$ and $0 \leq x_{2} \leq 1$ and zero outside this region. and the absolute value of the determinant of the Jacobian is

$$
\left\|\frac{\partial\left(x_{1}, x_{2}\right)}{\partial\left(y_{1}, y_{2}\right)}\right\|=\frac{1}{2 \pi} e^{-\frac{1}{2}\left(y_{1}^{2}+y_{2}^{2}\right)}\left(\frac{y_{1}^{2}}{y_{2}^{2}}+1\right) \cos ^{2} \arctan \left(\frac{y_{1}}{y_{2}}\right)
$$

but

$$
\left(\frac{y_{1}^{2}}{y_{2}^{2}}+1\right) \cos ^{2} \arctan \left(\frac{y_{1}}{y_{2}}\right)=\left(\tan ^{2} 2 \pi x_{2}+1\right) \cos ^{2} 2 \pi x_{2}=1
$$

and thus

$$
f\left(y_{1}, y_{2}\right)=\frac{1}{2 \pi} e^{-\frac{1}{2}\left(y_{1}^{2}+y_{2}^{2}\right)}=\frac{1}{\sqrt{2 \pi}} e^{-\frac{y_{1}^{2}}{2}} \frac{1}{\sqrt{2 \pi}} e^{-\frac{y_{2}^{2}}{2}}
$$

i.e. the product of two standard normal distributions.

Thus the result is that $y_{1}$ and $y_{2}$ are distributed as two independent standard normal variables. This is a well known method, often called the Box-Muller transformation, used in order to achieve pseudorandom numbers from the standard normal distribution given a uniform pseudorandom number generator (see below). The method was introduced by G. E. P. Box and M. E. Muller [25].

### 6.6 Probability Content

In figure 2 contours corresponding to one, two, and three standard deviations were shown. The projection on each axis for e.g. the one standard deviation contour covers the range $-1 \leq x_{i} \leq 1$ and contains a probability content of $68.3 \%$ which is well known from the one-dimensional case.

More generally, for a contour corresponding to $z$ standard deviations the contour has the equation

$$
\frac{\left(x_{1}+x_{2}\right)^{2}}{1+\rho}+\frac{\left(x_{1}-x_{2}\right)^{2}}{1-\rho}=2 z^{2}
$$

i.e. the major and minor semi-axes are $z \sqrt{1+\rho}$ and $z \sqrt{1-\rho}$, respectively. The function value at the contour is given by

$$
f\left(x_{1}, x_{2}\right)=\frac{1}{2 \pi \sqrt{1-\rho^{2}}} \exp \left\{-\frac{z^{2}}{2}\right\}
$$

Expressed in polar coordinates $(r, \phi)$ the contour is described by

$$
r^{2}=\frac{z^{2}\left(1-\rho^{2}\right)}{1-2 \rho \sin \phi \cos \phi}
$$

While the projected probability contents follow the usual figures for one-dimensional normal distributions the joint probability content within each ellipse is smaller. For the
one, two, and three standard deviation contours the probability content, regardless of the correlation coefficient $\rho$, inside the ellipse is approximately $39.3 \%, 86.5 \%$, and $98.9 \%$. If we would like to find the ellipse with a joint probability content of $68.3 \%$ we must chose $z \approx 1.5$ (for a content of $95.5 \%$ use $z \approx 2.5$ and for $99.7 \%$ use $z \approx 3.4$ ). Se further discussion on probability content for a multinormal distribution in section 28.3.

### 6.7 Random Number Generation

The joint distribution of $y_{1}$ and $y_{2}$ in section 6.5 above is a binormal distribution having $\rho=0$. For arbitrary correlation coefficients $\rho$ the binormal distribution is given by

$$
f\left(x_{1}, x_{2}\right)=\frac{1}{2 \pi \sigma_{1} \sigma_{2} \sqrt{1-\rho^{2}}} \cdot e^{-\frac{1}{2\left(1-\rho^{2}\right)}\left(\left(\frac{x_{1}-\mu_{1}}{\sigma_{1}}\right)^{2}+\left(\frac{x_{2}-\mu_{2}}{\sigma_{2}}\right)^{2}-2 \rho \cdot \frac{x_{1}-\mu_{1}}{\sigma_{1}} \cdot \frac{x_{2}-\mu_{2}}{\sigma_{2}}\right)}
$$

where $\mu_{1}$ and $\mu_{2}$ are the expectation values of $x_{1}$ and $x_{2}, \sigma_{1}$ and $\sigma_{2}$ their standard deviations and $\rho$ the correlation coefficient between them.

Variables distributed according to the binormal distribution may be obtained by transforming the two independent numbers $y_{1}$ and $y_{2}$ found in the section 6.5 either as

$$
\begin{aligned}
& z_{1}=\mu_{1}+\sigma_{1}\left(y_{1} \sqrt{1-\rho^{2}}+y_{2} \rho\right) \\
& z_{2}=\mu_{2}+\sigma_{2} y_{2}
\end{aligned}
$$

or as

$$
\begin{aligned}
& z_{1}=\mu_{1}+\frac{\sigma_{1}}{\sqrt{2}}\left(y_{1} \sqrt{1+\rho}+y_{2} \sqrt{1-\rho}\right) \\
& z_{2}=\mu_{2}+\frac{\sigma_{2}}{\sqrt{2}}\left(y_{1} \sqrt{1+\rho}-y_{2} \sqrt{1-\rho}\right)
\end{aligned}
$$

which can be proved by expressing $y_{1}$ and $y_{2}$ as functions of $z_{1}$ and $z_{2}$ and evaluate

$$
f\left(z_{1}, z_{2}\right)=\left\|\frac{\partial\left(y_{1}, y_{2}\right)}{\partial\left(z_{1}, z_{2}\right)}\right\| f\left(y_{1}, y_{2}\right)=\left\|\begin{array}{ll}
\frac{\partial y_{1}}{\partial z_{1}} & \frac{\partial y_{1}}{\partial z_{2}} \\
\frac{\partial y_{2}}{\partial z_{1}} & \frac{\partial y_{2}}{\partial z_{2}}
\end{array}\right\| f\left(y_{1}, y_{2}\right)
$$

In the first case

$$
\begin{aligned}
& y_{1}=\frac{1}{\sqrt{1-\rho^{2}}}\left(\frac{z_{1}-\mu_{1}}{\sigma_{1}}-\rho \frac{z_{2}-\mu_{2}}{\sigma_{2}}\right) \\
& y_{2}=\frac{z_{2}-\mu_{2}}{\sigma_{2}}
\end{aligned}
$$

and in the second case

$$
\begin{aligned}
& y_{1}=\frac{\sqrt{2}}{2 \sqrt{1+\rho}}\left(\frac{z_{1}-\mu_{1}}{\sigma_{1}}+\frac{z_{2}-\mu_{2}}{\sigma_{2}}\right) \\
& y_{2}=\frac{\sqrt{2}}{2 \sqrt{1-\rho}}\left(\frac{z_{1}-\mu_{1}}{\sigma_{1}}-\frac{z_{2}-\mu_{2}}{\sigma_{2}}\right)
\end{aligned}
$$

In both cases the absolute value of the determinant of the Jacobian is $1 / \sigma_{1} \sigma_{2} \sqrt{1-\rho^{2}}$ and we get

$$
f\left(z_{1}, z_{2}\right)=\frac{1}{\sigma_{1} \sigma_{2} \sqrt{1-\rho^{2}}} \cdot \frac{1}{\sqrt{2 \pi}} e^{-\frac{y_{1}^{2}}{2}} \cdot \frac{1}{\sqrt{2 \pi}} e^{-\frac{y_{2}^{2}}{2}}=\frac{1}{2 \pi \sigma_{1} \sigma_{2} \sqrt{1-\rho^{2}}} \cdot e^{-\frac{1}{2}\left(y_{1}^{2}+y_{2}^{2}\right)}
$$

Inserting the relations expressing $y_{1}$ and $y_{2}$ in $z_{1}$ and $z_{2}$ in the exponent we finally obtain the binormal distribution in both cases.

Thus we have found methods which given two independent uniform pseudorandom numbers between zero and one supplies us with a pair of numbers from a binormal distribution with arbitrary means, standard deviations and correlation coefficient.

## 7 Cauchy Distribution

### 7.1 Introduction

The Cauchy distribution is given by

$$
f(x)=\frac{1}{\pi} \cdot \frac{1}{1+x^{2}}
$$

and is defined for $-\infty<x<\infty$. It is a symmetric unimodal distribution as is shown in figure 3.


Figure 3: Graph of the Cauchy distribution
The distribution is named after the famous french mathematician Augustin Louis Cauchy (1789-1857) who was a professor at École Polytechnique in Paris from 1816. He was one of the most productive mathematicians which have ever existed.

### 7.2 Moments

This probability density function is peculiar inasmuch as it has undefined expectation value and all higher moments diverge. For the expectation value the integral

$$
E(x)=\frac{1}{\pi} \int_{-\infty}^{\infty} \frac{x}{1+x^{2}} d x
$$

is not completely convergent, i.e.

$$
\lim _{a \rightarrow \infty, b \rightarrow \infty} \frac{1}{\pi} \int_{-a}^{b} \frac{x}{1+x^{2}} d x
$$

does not exist. However, the principal value

$$
\lim _{a \rightarrow \infty} \frac{1}{\pi} \int_{-a}^{a} \frac{x}{1+x^{2}} d x
$$

does exist and is equal to zero. Anyway the convention is to regard the expectation value of the Cauchy distribution as undefined.

Other measures of location and dispersion which are useful in the case of the Cauchy distribution is the median and the mode which are at $x=0$ and the half-width at halfmaximum which is 1 (half-maxima at $x= \pm 1$ ).

### 7.3 Normalization

In spite of the somewhat awkward property of not having any moments the distribution at least fulfil the normalization requirement for a proper probability density function i.e.

$$
\mathcal{N}=\int_{-\infty}^{\infty} f(x) d x=\frac{1}{\pi} \int_{-\infty}^{\infty} \frac{1}{1+x^{2}} d x=\frac{1}{\pi} \int_{-\pi / 2}^{\pi / 2} \frac{1}{1+\tan ^{2} \phi} \cdot \frac{d \phi}{\cos ^{2} \phi}=1
$$

where we have made the substitution $\tan \phi=x$ in order to simplify the integration.

### 7.4 Characteristic Function

The characteristic function for the Cauchy distribution is given by

$$
\begin{aligned}
\phi(t) & =\int_{-\infty}^{\infty} e^{\imath t x} f(x) d x=\frac{1}{\pi} \int_{-\infty}^{\infty} \frac{\cos t x+\imath \sin t x}{1+x^{2}} d x= \\
& =\frac{1}{\pi}\left(\int_{0}^{\infty} \frac{\cos t x}{1+x^{2}} d x+\int_{-\infty}^{0} \frac{\cos t x}{1+x^{2}} d x+\int_{0}^{\infty} \frac{\imath \sin t x}{1+x^{2}} d x+\int_{-\infty}^{0} \frac{\imath \sin t x}{1+x^{2}} d x\right)= \\
& =\frac{2}{\pi} \int_{0}^{\infty} \frac{\cos t x}{1+x^{2}} d x=e^{-|t|}
\end{aligned}
$$

where we have used that the two sine integrals are equal but with opposite sign whereas the two cosine integrals are equal. The final integral we have taken from standard integral tables. Note that the characteristic function has no derivatives at $t=0$ once again telling us that the distribution has no moments.

### 7.5 Location and Scale Parameters

In the form given above the Cauchy distribution has no parameters. It is useful, however, to introduce location $\left(x_{0}\right)$ and scale $(\Gamma>0)$ parameters writing

$$
f\left(x ; x_{0}, \Gamma\right)=\frac{1}{\pi} \cdot \frac{\Gamma}{\Gamma^{2}+\left(x-x_{0}\right)^{2}}
$$

where $x_{0}$ is the mode of the distribution and $\Gamma$ the half-width at half-maximum (HWHM). Including these two parameters the characteristic function is modified to

$$
\phi(t)=e^{i t x_{0}-\Gamma|t|}
$$

### 7.6 Breit-Wigner Distribution

In this last form we recognize the Breit-Wigner formula, named after the two physicists Gregory Breit and Eugene Wigner, which arises in physics e.g. in the description of the cross section dependence on energy (mass) for two-body resonance scattering. Resonances like e.g. the $\Delta^{++}$in $\pi^{+} p$ scattering or the $\rho$ in $\pi \pi$ scattering can be quite well described in terms of the Cauchy distribution. This is the reason why the Cauchy distribution in physics often is referred to as the Breit-Wigner distribution. However, in more elaborate physics calculations the width may be energy-dependent in which case things become more complicated.

### 7.7 Comparison to Other Distributions

The Cauchy distribution is often compared to the normal (or Gaussian) distribution with mean $\mu$ and standard deviation $\sigma>0$

$$
f(x ; \mu, \sigma)=\frac{1}{\sigma \sqrt{2 \pi}} e^{-\frac{1}{2}\left(\frac{x-\mu}{\sigma}\right)^{2}}
$$

and the double-exponential distribution with mean $\mu$ and slope parameter $\lambda>0$

$$
f(x ; \mu, \lambda)=\frac{\lambda}{2} e^{-\lambda|x-\mu|}
$$

These are also examples of symmetric unimodal distributions. The Cauchy distribution has longer tails than the double-exponential distribution which in turn has longer tails than the normal distribution. In figure 4 we compare the Cauchy distribution with the standard normal ( $\mu=0$ and $\sigma=1$ ) and the double-exponential distributions $(\lambda=1)$ for $x>0$.

The normal and double-exponential distributions have well defined moments. Since they are symmetric all central moments of odd order vanish while central moments of even order are given by $\mu_{2 n}=(2 n)!\sigma^{2 n} / 2^{n} n!$ (for $n \geq 0$ ) for the normal and by $\mu_{n}=n!/ \lambda^{n}$ (for even $n$ ) for the double-exponential distribution. E.g. the variances are $\sigma^{2}$ and $2 / \lambda^{2}$ and the fourth central moments $3 \sigma^{4}$ and $24 / \lambda^{4}$, respectively.

The Cauchy distribution is related to Student's $t$-distribution with $n$ degrees of freedom (with $n$ a positive integer)

$$
f(t ; n)=\frac{\Gamma\left(\frac{n+1}{2}\right)}{\sqrt{n \pi} \Gamma\left(\frac{n}{2}\right)}\left(1+\frac{t^{2}}{n}\right)^{-\frac{n+1}{2}}=\frac{\left(1+\frac{t^{2}}{n}\right)^{-\frac{n+1}{2}}}{\sqrt{n} B\left(\frac{1}{2}, \frac{n}{2}\right)}
$$

where $\Gamma(x)$ is the Euler gamma-function not no be mixed up with the width parameter for the Cauchy distribution used elsewhere in this section. $B$ is the beta-function defined in terms of the $\Gamma$-function as $B(p, q)=\frac{\Gamma(p) \Gamma(q)}{\Gamma(p+q)}$. As can be seen the Cauchy distribution arises


Figure 4: Comparison between the Cauchy distribution, the standard normal distribution, and the double-exponential distribution
as the special case where $n=1$. If we change variable to $x=t / \sqrt{n}$ and put $m=\frac{n+1}{2}$ the Student's $t$-distribution becomes

$$
f(x ; m)=\frac{k}{\left(1+x^{2}\right)^{m}} \quad \text { with } \quad k=\frac{\Gamma(m)}{\Gamma\left(\frac{1}{2}\right) \Gamma\left(m-\frac{1}{2}\right)}=\frac{1}{B\left(m-\frac{1}{2}, \frac{1}{2}\right)}
$$

where $k$ is simply a normalization constant. Here it is easier to see the more general form of this distribution which for $m=1$ gives the Cauchy distribution. The requirement $n \geq 1$ corresponds to $m$ being a half-integer $\geq 1$ but we could even allow for $m$ being a real number.

As for the Cauchy distribution the Student's $t$-distribution have problems with divergent moments and moments of order $\geq n$ does not exist. Below this limit odd central moments are zero (the distribution is symmetric) and even central moments are given by

$$
\mu_{2 r}=n^{r} \frac{\Gamma\left(r+\frac{1}{2}\right) \Gamma\left(\frac{n}{2}-r\right)}{\Gamma\left(\frac{1}{2}\right) \Gamma\left(\frac{n}{2}\right)}=n^{r} \frac{B\left(r+\frac{1}{2}, \frac{n}{2}-r\right)}{B\left(\frac{1}{2}, \frac{n}{2}\right)}
$$

for $r$ a positive integer $(2 r<n)$. More specifically the expectation value is $E(t)=0$, the variance $V(t)=\frac{n}{n-2}$ and the fourth central moment is given by $\mu_{4}=\frac{3 n^{2}}{(n-2)(n-4)}$ when they exist. As $n \rightarrow \infty$ the Student's $t$-distribution approaches a standard normal distribution.

### 7.8 Truncation

In order to avoid the long tails of the distribution one sometimes introduces a truncation. This, of course, also cures the problem with the undefined mean and divergent higher
moments. For a symmetric truncation $-X \leq x \leq X$ we obtain the renormalized probability density function

$$
f(x)=\frac{1}{2 \arctan X} \cdot \frac{1}{1+x^{2}}
$$

which has expectation value $E(x)=0$, variance $V(x)=\frac{X}{\arctan X}-1$, third central moment $\mu_{3}=0$ and fourth central moment $\mu_{4}=\frac{X}{\arctan X}\left(\frac{X^{2}}{3}-1\right)+1$. The fraction of the original Cauchy distribution within the symmetric interval is $f=\frac{2}{\pi} \arctan X$. We will, however, not make any truncation of the Cauchy distribution in the considerations made in this note.

### 7.9 Sum and Average of Cauchy Variables

In most cases one would expect the sum and average of many variables drawn from the same population to approach a normal distribution. This follows from the famous Central Limit Theorem. However, due to the divergent variance of the Cauchy distribution the requirements for this theorem to hold is not fulfilled and thus this is not the case here. We define

$$
S_{n}=\sum_{i=1}^{n} x_{i} \quad \text { and } \quad \bar{S}_{n}=\frac{1}{n} S_{n}
$$

with $x_{i}$ independent variables from a Cauchy distribution.
The characteristic function of a sum of independent random variables is equal to the product of the individual characteristic functions and hence

$$
\Phi(t)=\phi(t)^{n}=e^{-n|t|}
$$

for $S_{n}$. Turning this into a probability density function we get (putting $x=S_{n}$ for convenience)

$$
\begin{aligned}
f(x) & =\frac{1}{2 \pi} \int_{-\infty}^{\infty} \Phi(t) e^{-\imath x t} d t=\frac{1}{2 \pi} \int_{-\infty}^{\infty} e^{-(\imath x t+n|t|)} d t=\frac{1}{2 \pi}\left(\int_{-\infty}^{0} e^{n t-\imath x t} d t+\int_{0}^{\infty} e^{-\imath x t-n t} d t\right)= \\
& =\frac{1}{2 \pi}\left(\left[\frac{e^{t(n-\imath x)}}{n-\imath x}\right]_{-\infty}^{0}+\left[\frac{e^{-t(2 x+n)}}{-n-\imath x}\right]_{0}^{\infty}\right)=\frac{1}{2 \pi}\left(\frac{1}{n-\imath x}+\frac{1}{n+\imath x}\right)=\frac{1}{\pi} \cdot \frac{n}{n^{2}+x^{2}}
\end{aligned}
$$

This we recognize as a Cauchy distribution with scale parameter $\Gamma=n$ and thus for each additional Cauchy variable the HWHM increases by one unit.

Moreover, the probability density function of $\bar{S}_{n}$ is given by

$$
f\left(\bar{S}_{n}\right)=\frac{d S_{n}}{d \bar{S}_{n}} f\left(S_{n}\right)=\frac{1}{\pi} \cdot \frac{1}{1+\bar{S}_{n}^{2}}
$$

i.e. the somewhat amazing result is that the average of any number of independent random variables from a Cauchy distribution is also distributed according to the Cauchy distribution.

### 7.10 Estimation of the Median

For the Cauchy distribution the sample mean is not a consistent estimator of the median of the distribution. In fact, as we saw in the previous section, the sample mean is itself distributed according to the Cauchy distribution and therefore has divergent variance. However, the sample median for a sample of $n$ independent observations from a Cauchy distribution is a consistent estimator of the true median.

In the table below we give the expectations and variances of the sample mean and sample median estimators for the normal, double-exponential and Cauchy distributions (see above for definitions of distributions). Sorting all the observations the median is taken as the value for the central observation for odd $n$ and as the average of the two central values for even $n$. The variance of the sample mean is simply the variance of the distribution divided by the sample size $n$. For large $n$ the variance of the sample median $m$ is given by $V(m)=1 / 4 n f^{2}$ where $f$ is the function value at the median.

| Distribution | $E(\bar{x})$ | $V(\bar{x})$ | $E(m)$ | $V(m)$ |
| :--- | :---: | :---: | :---: | :---: |
| Normal | $\mu$ | $\frac{\sigma^{2}}{n}$ | $\mu$ | $\frac{\pi \sigma^{2}}{2 n}$ |
| Double-exponential | $\mu$ | $\frac{2}{n \lambda^{2}}$ | $\mu$ | $\frac{1}{n \lambda^{2}}$ |
| Cauchy | undef. | $\infty$ | $x_{0}$ | $\frac{\pi^{2} \Gamma^{2}}{4 n}$ |

For a normal distribution the sample mean is superior to the median as an estimator of the mean (i.e. it has the smaller variance). However, the double-exponential distribution is an example of a distribution where the sample median is the best estimator of the mean of the distribution. In the case of the Cauchy distribution only the median works of the above alternatives but even better is a proper Maximum Likelihood estimator. In the case of the normal and double-exponential the mean and median, respectively, are identical to the maximum likelihood estimators but for the Cauchy distribution such an estimator may not be expressed in a simple way.

The large $n$ approximation for the variance of the sample median gives conservative estimates for lower values of $n$ in the case of the normal distribution. Beware, however, that for the Cauchy and the double-exponential distributions it is not conservative but gives too small values. Calculating the standard deviation this is within $10 \%$ of the true value already at $n=5$ for the normal distribution whereas for the Cauchy distribution this is true at about $n=20$ and for the double-exponential distribution only at about $n=60$.

### 7.11 Estimation of the HWHM

To find an estimator for the half-width at half-maximum is not trivial. It implies binning the data, finding the maximum and then locating the positions where the curve is at halfmaximum. Often it is preferable to fit the probability density function to the observations in such a case.

However, it turns out that another measure of dispersion the so called semi-interquartile range can be used as an estimator. The semi-interquartile range is defined as half the difference between the upper and the lower quartiles. The quartiles are the values which divide the probability density function into four parts with equal probability content, i.e. $25 \%$ each. The second quartile is thus identical to the median. The definition is thus

$$
\mathcal{S}=\frac{1}{2}\left(Q_{3}-Q_{1}\right)=\Gamma
$$

where $Q_{1}$ is the lower and $Q_{3}$ the upper quartile which for the Cauchy distribution is equal to $x_{0}-\Gamma$ and $x_{0}+\Gamma$, respectively. As is seen $\mathcal{S}=H W H M=\Gamma$ and thus this estimator may be used in order to estimate $\Gamma$.

We gave above the large $n$ approximation for the variance of the median. The median and the quartiles are examples of the more general concept quantiles. Generally the large $n$ approximation for the variance of a quantile $\mathcal{Q}$ is given by $V(\mathcal{Q})=p q / n f^{2}$ where $f$ is the ordinate at the quantile and $p$ and $q=1-p$ are the probability contents above and below the quantile, respectively. The covariance between two quantiles $\mathcal{Q}_{1}$ and $\mathcal{Q}_{2}$ is, with similar notations, given by $\operatorname{Cov}\left(\mathcal{Q}_{1}, \mathcal{Q}_{2}\right)=p_{2} q_{1} / n f_{1} f_{2}$ where $\mathcal{Q}_{1}$ should be the leftmost quantile.

For large $n$ the variance of the semi-interquartile range for a sample of size $n$ is thus found by error propagation inserting the formulæ above

$$
V(\mathcal{S})=\frac{1}{4}\left(V\left(Q_{1}\right)+V\left(Q_{3}\right)-2 \operatorname{Cov}\left(Q_{1}, Q_{3}\right)\right)=\frac{1}{64 n}\left(\frac{3}{f_{1}^{2}}+\frac{3}{f_{3}^{2}}-\frac{2}{f_{1} f_{3}}\right)=\frac{1}{16 n f_{1}^{2}}=\frac{\pi^{2} \Gamma^{2}}{4 n}
$$

where $f_{1}$ and $f_{3}$ are the function values at the lower and upper quartile which are both equal to $1 / 2 \pi \Gamma$. This turns out to be exactly the same as the variance we found for the median in the previous section.

After sorting the sample the quartiles are determined by extrapolation between the two observations closest to the quartile. In the case where $n+2$ is a multiple of 4 i.e. the series $n=2,6,10 \ldots$ the lower quartile is exactly at the $\frac{n+2}{4}$ :th observation and the upper quartile at the $\frac{3 n+2}{4}$ :th observation. In the table below we give the expectations and variances of the estimator of $\mathcal{S}$ as well as the variance estimator $s^{2}$ for the normal, double-exponential and Cauchy distributions. The variance estimator $s^{2}$ and its variance are given by

$$
s^{2}=\frac{1}{n-1} \sum_{i=1}^{n}\left(x_{i}-\bar{x}\right)^{2} \quad \text { and } \quad V\left(s^{2}\right)=\frac{\mu_{4}-\mu_{2}^{2}}{n}+\frac{2 \mu_{2}^{2}}{n(n-1)}
$$

with $\mu_{2}$ and $\mu_{4}$ the second and fourth central moments. The expectation value of $s^{2}$ is equal to the variance and thus it is a unbiased estimator.

| Distribution | HWHM | $E\left(s^{2}\right)$ | $V\left(s^{2}\right)$ | $E(\mathcal{S})$ | $V(\mathcal{S})$ |
| :--- | :---: | :---: | :---: | :---: | :---: |
| Normal | $\sigma \sqrt{2 \ln 2}$ | $\sigma^{2}$ | $\frac{2 \sigma^{4}}{n-1}$ | $0.6745 \sigma$ | $\frac{1}{16 n f\left(Q_{1}\right)^{2}}$ |
| Double-exponential | $\frac{\ln 2}{\lambda}$ | $\frac{2}{\lambda^{2}}$ | $\frac{20}{n \lambda^{4}} \alpha$ | $\frac{\ln 2}{\lambda}$ | $\frac{1}{n \lambda^{2}}$ |
| Cauchy | $\Gamma$ | $\infty$ | $\infty$ | $\Gamma$ | $\frac{\pi^{2} \Gamma^{2}}{4 n}$ |

In this table $\alpha=1+\frac{0.4}{n-1}$ if we include the second term in the expression of $V\left(s^{2}\right)$ above and $\alpha=1$ otherwise. It can be seen that the double-exponential distribution also has $H W H M=\mathcal{S}$ but for the normal distribution $H W H M \approx 1.1774 \sigma$ as compared to $\mathcal{S} \approx$ $0.6745 \sigma$.

For the three distributions tested the semi-interquartile range estimator is biased. In the case of the normal distribution the values are approaching the true value from below while for the Cauchy and double-exponential distributions from above. The large $n$ approximation for $V(\mathcal{S})$ is conservative for the normal and double-exponential distribution but not conservative for the Cauchy distribution. In the latter case the standard deviation of $\mathcal{S}$ is within $10 \%$ of the true value for $n>50$ but for small values of $n$ it is substantially larger than given by the formula. The estimated value for $\Gamma$ is less than $10 \%$ too big for $n>25$.

### 7.12 Random Number Generation

In order to generate pseudorandom numbers from a Cauchy distribution we may solve the equation $F(x)=\xi$ where $F(x)$ is the cumulative distribution function and $\xi$ is a uniform pseudorandom number between 0 and 1 . This means solving for $x$ in the equation

$$
F(x)=\frac{\Gamma}{\pi} \int_{-\infty}^{x} \frac{1}{\Gamma^{2}+\left(t-x_{0}\right)^{2}} d t=\xi
$$

If we make the substitution $\tan \phi=\left(t-x_{0}\right) / \Gamma$ using that $d \phi / \cos ^{2} \phi=d t / \Gamma$ we obtain

$$
\frac{1}{\pi} \arctan \left(\frac{x-x_{0}}{\Gamma}\right)+\frac{1}{2}=\xi
$$

which finally gives

$$
x=x_{0}+\Gamma \tan \left(\pi\left(\xi-\frac{1}{2}\right)\right)
$$

as a pseudorandom number from a Cauchy distribution. One may easily see that it is equivalent to use

$$
x=x_{0}+\Gamma \tan (2 \pi \xi)
$$

which is a somewhat simpler expression.
An alternative method (see also below) to achieve random numbers from a Cauchy distribution would be to use

$$
x=x_{0}+\Gamma \frac{z_{1}}{z_{2}}
$$

where $z_{1}$ and $z_{2}$ are two independent random numbers from a standard normal distribution. However, if the standard normal random numbers are achieved through the Box-Muller transformation then $z_{1} / z_{2}=\tan 2 \pi \xi$ and we are back to the previous method.

In generating pseudorandom numbers one may, if profitable, avoid the tangent by
a Generate in $u$ and $v$ two random numbers from a uniform distribution between -1 and 1.
b If $u^{2}+v^{2}>1$ (outside circle with radius one in $u v$-plane) go back to a.
c Obtain $x=x_{0}+\Gamma \frac{u}{v}$ as a random number from a Cauchy distribution.

### 7.13 Physical Picture

A physical picture giving rise to the Cauchy distribution is as follows: Regard a plane in which there is a point source which emits particles isotropically in the plane (either in the full $2 \pi$ region or in one hemisphere $\pi$ radians wide). The source is at the $x$-coordinate $x_{0}$ and the particles are detected in a detector extending along a line $\Gamma$ length units from the source. This scenario is depicted in figure 5


Figure 5: Physical scenario leading to a Cauchy distribution
The distribution in the variable $x$ along the detector will then follow the Cauchy distribution. As can be seen by pure geometrical considerations this is in accordance with the result above where pseudorandom numbers from a Cauchy distribution could be obtained by $x=x_{0}+\Gamma \tan \phi$, i.e. $\tan \phi=\frac{x-x_{0}}{\Gamma}$, with $\phi$ uniformly distributed between $-\frac{\pi}{2}$ and $\frac{\pi}{2}$.

To prove this let us start with the distribution in $\phi$

$$
f(\phi)=\frac{1}{\pi} \quad \text { for } \quad-\frac{\pi}{2} \leq \phi \leq \frac{\pi}{2}
$$

To change variables from $\phi$ to $x$ requires the derivative $d \phi / d x$ which is given by

$$
\frac{d \phi}{d x}=\frac{\cos ^{2} \phi}{\Gamma}=\frac{1}{\Gamma} \cos ^{2} \arctan \left(\frac{x-x_{0}}{\Gamma}\right)
$$

Note that the interval from $-\frac{\pi}{2}$ to $\frac{\pi}{2}$ in $\phi$ maps onto the interval $-\infty<x<\infty$. We get

$$
\begin{aligned}
f(x) & =\left|\frac{d \phi}{d x}\right| f(\phi)=\frac{1}{\pi \Gamma} \cos ^{2} \arctan \left(\frac{x-x_{0}}{\Gamma}\right)=\frac{1}{\pi \Gamma} \cos ^{2} \phi= \\
& =\frac{1}{\pi \Gamma} \cdot \frac{\Gamma^{2}}{\Gamma^{2}+\left(x-x_{0}\right)^{2}}=\frac{1}{\pi} \cdot \frac{\Gamma}{\Gamma^{2}+\left(x-x_{0}\right)^{2}}
\end{aligned}
$$

i.e. the Cauchy distribution.

It is just as easy to make the proof in the reversed direction, i.e. given a Cauchy distribution in $x$ one may show that the $\phi$-distribution is uniform between $-\frac{\pi}{2}$ and $\frac{\pi}{2}$.

### 7.14 Ratio Between Two Standard Normal Variables

As mentioned above the Cauchy distribution also arises if we take the ratio between two standard normal variables $z_{1}$ and $z_{2}$, viz.

$$
x=x_{0}+\Gamma \frac{z_{1}}{z_{2}} .
$$

In order to deduce the distribution in $x$ we first introduce a dummy variable $y$ which we simply take as $z_{2}$ itself. We then make a change of variables from $z_{1}$ and $z_{2}$ to $x$ and $y$. The transformation is given by

$$
\begin{aligned}
x & =x_{0}+\Gamma \frac{z_{1}}{z_{2}} \\
y & =z_{2}
\end{aligned}
$$

or if we express $z_{1}$ and $z_{2}$ in $x$ and $y$

$$
\begin{aligned}
& z_{1}=y\left(x-x_{0}\right) / \Gamma \\
& z_{2}=y
\end{aligned}
$$

The distribution in $x$ and $y$ is given by

$$
f(x, y)=\left\|\frac{\partial\left(z_{1}, z_{2}\right)}{\partial(x, y)}\right\| f\left(z_{1}, z_{2}\right)
$$

where the absolute value of the determinant of the Jacobian is equal to $y / \Gamma$ and $f\left(z_{1}, z_{2}\right)$ is the product of two independent standard normal distributions. We get

$$
f(x, y)=\frac{y}{\Gamma} \cdot \frac{1}{2 \pi} e^{-\frac{1}{2}\left(z_{1}^{2}+z_{2}^{2}\right)}=\frac{y}{2 \pi \Gamma} e^{-\frac{1}{2}\left(\frac{y^{2}\left(x-x_{0}\right)^{2}}{\Gamma^{2}}+y^{2}\right)}
$$

In order to obtain the marginal distribution in $x$ we integrate over $y$

$$
f(x)=\int_{-\infty}^{\infty} f(x, y) d y=\frac{1}{2 \pi \Gamma} \int_{-\infty}^{\infty} y e^{-\alpha y^{2}} d y
$$

where we have put

$$
\alpha=\frac{1}{2}\left(\left(\frac{x-x_{0}}{\Gamma}\right)^{2}+1\right)
$$

for convenience. If we make the substitution $z=y^{2}$ we get

$$
f(x)=2 \frac{1}{2 \pi \Gamma} \int_{0}^{\infty} e^{-\alpha z} \frac{d z}{2}=\frac{1}{2 \pi \Gamma \alpha}
$$

Note that the first factor of 2 comes from the fact that the region $-\infty<y<\infty$ maps twice onto the region $0<z<\infty$. Finally

$$
f(x)=\frac{1}{2 \pi \Gamma \alpha}=\frac{1}{2 \pi \Gamma} \cdot \frac{2}{\left(\frac{x-x_{0}}{\Gamma}\right)^{2}+1}=\frac{1}{\pi} \cdot \frac{\Gamma}{\left(x-x_{0}\right)^{2}+\Gamma^{2}}
$$

i.e. a Cauchy distribution.

## 8 Chi-square Distribution

### 8.1 Introduction

The chi-square distribution is given by

$$
f(x ; n)=\frac{\left(\frac{x}{2}\right)^{\frac{n}{2}-1} e^{-\frac{x}{2}}}{2 \Gamma\left(\frac{n}{2}\right)}
$$

where the variable $x \geq 0$ and the parameter $n$, the number of degrees of freedom, is a positive integer. In figure 6 the distribution is shown for $n$-values of $1,2,5$ and 10 . For $n \geq 2$ the distribution has a maximum at $n-2$.


Figure 6: Graph of chi-square distribution for some values of $n$

### 8.2 Moments

Algebraic moments of order $k$ are given by

$$
\begin{aligned}
\mu_{k}^{\prime} & =E\left(x^{k}\right)=\frac{1}{2 \Gamma\left(\frac{n}{2}\right)} \int_{0}^{\infty} x^{k}\left(\frac{x}{n}\right)^{\frac{n}{2}-1} e^{-\frac{x^{2}}{2}} d x=\frac{2^{k}}{\Gamma\left(\frac{n}{2}\right)} \int_{0}^{\infty} y^{\frac{n}{2}-1+k} e^{-y} d y=\frac{2^{k} \Gamma\left(\frac{n}{2}+k\right)}{\Gamma\left(\frac{n}{2}\right)}= \\
& =2^{k} \cdot \frac{n}{2}\left(\frac{n}{2}+1\right) \cdots\left(\frac{n}{2}+k-2\right)\left(\frac{n}{2}+k-1\right)=n(n+2)(n+4) \cdots(n+2 k-2)
\end{aligned}
$$

e.g. the first algebraic moment which is the expectation value is equal to $n$. A recursive formula to calculate algebraic moments is thus given by

$$
\mu_{k}^{\prime}=\mu_{k-1}^{\prime} \cdot(n+2 k-2)
$$

where we may start with $\mu_{0}^{\prime}=1$ to find the expectation value $\mu_{1}^{\prime}=n, \mu_{2}^{\prime}=n(n+1)$ etc.

From this we may calculate the central moments which for the lowest orders become $\mu_{2}=2 n, \mu_{3}=8 n, \mu_{4}=12 n(n+4), \mu_{5}=32 n(5 n+12)$ and $\mu_{6}=40 n\left(3 n^{2}+52 n+96\right)$. The coefficients of skewness and kurtosis thus becomes $\gamma_{1}=2 \sqrt{2 / n}$ and $\gamma_{2}=12 / n$.

The fact that the expectation value of a chi-square distribution equals the number of degrees of freedom has led to a bad habit to give the ratio between a found chi-square value and the number of degrees of freedom. This is, however, not a very good variable and it may be misleading. We strongly recommend that one always should give both the chi-square value and degrees of freedom e.g. as $\chi^{2} /$ n.d.f. $=9.7 / 5$.

To judge the quality of the fit we want a better measure. Since the exact sampling distribution is known one should stick to the chi-square probability as calculated from an integral of the tail i.e. given a specific chi-square value for a certain number of degrees of freedom we integrate from this value to infinity (see below).

As an illustration we show in figure 7 the chi-square probability for constant ratios of $\chi^{2} /$ n.d.f.


Figure 7: Chi-square probability for constant ratios of $\chi^{2} /$ n.d.f.
Note e.g. that for few degrees of freedom we may have an acceptable chi-square value even for larger ratios.

### 8.3 Characteristic Function

The characteristic function for a chi-square distribution with $n$ degrees of freedom is given by

$$
\phi(t)=E\left(e^{\imath t x}\right)=\frac{1}{2 \Gamma\left(\frac{n}{2}\right)} \int_{0}^{\infty}\left(\frac{x}{2}\right)^{\frac{n}{2}-1} e^{-\left(\frac{1}{2}-\imath t\right) x} d x=\frac{1}{2 \Gamma\left(\frac{n}{2}\right)} \int_{0}^{\infty}\left(\frac{y}{1-2 \imath t}\right)^{\frac{n}{2}-1} e^{-y} \frac{d y}{\frac{1}{2}-\imath t}=
$$

$$
=\frac{1}{\Gamma\left(\frac{n}{2}\right)(1-2 t)^{\frac{n}{2}}} \int_{0}^{\infty} y^{\frac{n}{2}-1} e^{-y} d y=(1-2 t t)^{-\frac{n}{2}}
$$

### 8.4 Cumulative Function

The cumulative, or distribution, function for a chi-square distribution with $n$ degrees of freedom is given by

$$
\begin{aligned}
F(x) & =\frac{1}{2 \Gamma\left(\frac{n}{2}\right)} \int_{0}^{x}\left(\frac{x}{2}\right)^{\frac{n}{2}-1} e^{-\frac{x}{2}} d x=\frac{1}{2 \Gamma\left(\frac{n}{2}\right)} \int_{0}^{\frac{x}{2}} y^{\frac{n}{2}-1} e^{-y} 2 d y= \\
& =\frac{\gamma\left(\frac{n}{2}, \frac{x}{2}\right)}{\Gamma\left(\frac{n}{2}\right)}=P\left(\frac{n}{2}, \frac{x}{2}\right)
\end{aligned}
$$

where $P\left(\frac{n}{2}, \frac{x}{2}\right)$ is the incomplete Gamma function (see section 42.5). In this calculation we have made the simple substitution $y=x / 2$ in simplifying the integral.

### 8.5 Origin of the Chi-square Distribution

If $z_{1}, z_{2}, \ldots, z_{n}$ are $n$ independent standard normal random variables then $\sum_{i=1}^{n} z_{i}^{2}$ is distributed as a chi-square variable with $n$ degrees of freedom.

In order to prove this first regard the characteristic function for the square of a standard normal variable

$$
E\left(e^{\imath t z^{2}}\right)=\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} e^{-\frac{z^{2}}{2}(1-2 \imath t)} d z=\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} e^{-\frac{y^{2}}{2}} \frac{d y}{\sqrt{1-2 \imath t}}=\frac{1}{\sqrt{1-2 \imath t}}
$$

where we made the substitution $y=z \sqrt{1-2 \imath t}$.
For a sum of $n$ such independent variables the characteristic function is then given by

$$
\phi(t)=(1-2 i t)^{-\frac{n}{2}}
$$

which we recognize as the characteristic function for a chi-square distribution with $n$ degrees of freedom.

This property implies that if $x$ and $y$ are independently distributed according to the chisquare distribution with $n$ and $m$ degrees of freedom, respectively, then $x+y$ is distributed as a chi-square variable with $m+n$ degrees of freedom.

Indeed the requirement that all $z$ 's come from a standard normal distribution is more than what is needed. The result is the same if all observations $x_{i}$ come from different normal populations with means $\mu_{i}$ and variance $\sigma_{i}^{2}$ if we in each case calculate a standardized variable by subtracting the mean and dividing with the standard deviation i.e taking $z_{i}=\left(x_{i}-\mu_{i}\right) / \sigma_{i}$.

### 8.6 Approximations

For large number of degrees of freedom $n$ the chi-square distribution may be approximated by a normal distribution. There are at least three different approximations. Firstly we may naïvely construct a standardized variable

$$
z_{1}=\frac{x-E(x)}{\sqrt{V(x)}}=\frac{x-n}{\sqrt{2 n}}
$$

which would tend to normality as $n$ increases. Secondly an approximation, due to R. A. Fisher, is that the quantity

$$
z_{2}=\sqrt{2 x}-\sqrt{2 n-1}
$$

approaches a standard normal distribution faster than the standardized variable. Thirdly a transformation, due to E. B. Wilson and M. M. Hilferty, is that the cubic root of $x / n$ is closely distributed as a standard normal distribution using

$$
z_{3}=\frac{\left(\frac{x}{n}\right)^{\frac{1}{3}}-\left(1-\frac{2}{9 n}\right)}{\sqrt{\frac{2}{9 n}}}
$$

The second approximation is probably the most well known but the latter is approaching normality even faster. In fact there are even correction factors which may be applied to $z_{3}$ to give an even more accurate approximation (see e.g. [26])

$$
z_{4}=z_{3}+h_{n}=z_{3}+\frac{60}{n} h_{60}
$$

with $h_{60}$ given for values of $z_{2}$ from -3.5 to 3.5 in steps of 0.5 (in this order the values of $h_{60}$ are $-0.0118,-0.0067,-0.0033,-0.0010,0.0001,0.0006,0.0006,0.0002,-0.0003,-0.0006$, $-0.0005,0.0002,0.0017,0.0043$, and 0.0082 ).

To compare the quality of all these approximations we calculate the maximum deviation between the cumulative function for the true chi-square distribution and each of these approximations for $n=30$ and $n=100$. The results are shown in the table below. Normally one accepts $z_{2}$ for $n>100$ while $z_{3}$, and certainly $z_{4}$, are even better already for $n>30$.

| Approximation | $n=30$ | $n=100$ |
| :---: | :--- | :--- |
| $z_{1}$ | 0.034 | 0.019 |
| $z_{2}$ | 0.0085 | 0.0047 |
| $z_{3}$ | 0.00039 | 0.00011 |
| $z_{4}$ | 0.000044 | 0.000035 |

### 8.7 Random Number Generation

As we saw above the sum of $n$ independent standard normal random variables gave a chi-square distribution with $n$ degrees of freedom. This may be used as a technique to produce pseudorandom numbers from a chi-square distribution. This required a generator for standard normal random numbers and may be quite slow. However, if we make use of
the Box-Muller transformation in order to obtain the standard normal random numbers we may simplify the calculations.

First we recall the Box-Muller transformation which given two pseudorandom numbers uniformly distributed between zero and one through the transformation

$$
\begin{aligned}
& z_{1}=\sqrt{-2 \ln \xi_{1}} \cos 2 \pi \xi_{2} \\
& z_{2}=\sqrt{-2 \ln \xi_{1}} \sin 2 \pi \xi_{2}
\end{aligned}
$$

gives, in $z_{1}$ and $z_{2}$, two independent pseudorandom numbers from a standard normal distribution.

Adding $n$ such squared random numbers implies that

$$
\begin{aligned}
y_{2 k} & =-2 \ln \left(\xi_{1} \cdot \xi_{2} \cdots \xi_{k}\right) \\
y_{2 k+1} & =-2 \ln \left(\xi_{1} \cdot \xi_{2} \cdots \xi_{k}\right)-2 \ln \xi_{k+1} \cos ^{2} 2 \pi \xi_{k+2}
\end{aligned}
$$

for $k$ a positive integer will be distributed as chi-square variable with even or odd number of degrees of freedom. In this manner a lot of unnecessary operations are avoided.

Since the chi-square distribution is a special case of the Gamma distribution we may also use a generator for this distribution.

### 8.8 Confidence Intervals for the Variance

If $x_{1}, x_{2}, \ldots, x_{n}$ are independent normal random variables from a $N\left(\mu, \sigma^{2}\right)$ distribution then $\frac{(n-1) s^{2}}{\sigma^{2}}$ is distributed according to the chi-square distribution with $n-1$ degrees of freedom.

A $1-\alpha$ confidence interval for the variance is then given by

$$
\frac{(n-1) s^{2}}{\chi_{1-\alpha / 2, n-1}^{2}} \leq \sigma^{2} \leq \frac{(n-1) s^{2}}{\chi_{\alpha / 2, n-1}^{2}}
$$

where $\chi_{\alpha, n}$ is the chi-square value for a distribution with $n$ degrees of freedom for which the probability to be greater or equal to this value is given by $\alpha$. See also below for calculations of the probability content of the chi-square distribution.

### 8.9 Hypothesis Testing

Let $x_{1}, x_{2}, \ldots, x_{n}$ be $n$ independent normal random variables distributed according to a $N\left(\mu, \sigma^{2}\right)$ distribution. To test the null hypothesis $H_{0}: \sigma^{2}=\sigma_{0}^{2}$ versus $H_{1}: \sigma^{2} \neq \sigma_{0}^{2}$ at the $\alpha$ level of significance, we would reject the null hypothesis if $(n-1) s^{2} / \sigma_{0}^{2}$ is less than $\chi_{\alpha / 2, n-1}^{2}$ or greater than $\chi_{1-\alpha / 2, n-1}^{2}$.

### 8.10 Probability Content

In testing hypotheses using the chi-square distribution we define $x_{\alpha}=\chi_{\alpha, n}^{2}$ from

$$
F\left(x_{\alpha}\right)=\int_{0}^{x_{\alpha}} f(x ; n) d x=1-\alpha
$$

i.e. $\alpha$ is the probability that a variable distributed according to the chi-square distribution with $n$ degrees of freedom exceeds $x_{\alpha}$.

This formula can be used in order to determine confidence levels for certain values of $\alpha$. This is what is done in producing the tables which is common in all statistics text-books. However, more often the equation is used in order to calculate the confidence level $\alpha$ given an experimentally determined chi-square value $x_{\alpha}$.

In calculating the probability content of a chi-square distribution we differ between the case with even and odd number of degrees of freedom. This is described in the two following subsections.

Note that one may argue that it is as unlikely to obtain a very small chi-square value as a very big one. It is customary, however, to use only the upper tail in calculation of significance levels. A too small chi-square value is regarded as not a big problem. However, in such a case one should be somewhat critical since it indicates that one either is cheating, are using selected (biased) data or has (undeliberately) overestimated measurement errors (e.g. included systematic errors).

To proceed in calculating the cumulative function we write

$$
1-\alpha=F\left(x_{\alpha}\right)=\frac{1}{2 \Gamma\left(\frac{n}{2}\right)} \int_{0}^{x_{\alpha}}\left(\frac{x}{2}\right)^{\frac{n}{2}-1} e^{-\frac{x}{2}} d x=\frac{1}{\Gamma\left(\frac{n}{2}\right)} \int_{0}^{x_{\alpha} / 2} z^{\frac{n}{2}-1} e^{-z} d z=P\left(\frac{n}{2}, \frac{x_{\alpha}}{2}\right)
$$

where we have made the substitution $z=x / 2$. From this we see that we may use the incomplete Gamma function $P$ (see section 42.5) in evaluating probability contents but for historical reasons we have solved the problem by considering the cases with even and odd degrees of freedom separately as is shown in the next two subsections.

Although we prefer exact routines to calculate the probability in each specific case a classical table may sometimes be useful. In table 1 on page 171 we show percentage points, i.e. points where the cumulative probability is $1-\alpha$, for different degrees of freedom.

It is sometimes of interest e.g. when rejecting a hypothesis using a chi-square test to scrutinize extremely small confidence levels. In table 2 on page 172 we show this for confidence levels down to $10^{-12}$ as chi-square values. In table 3 on page 173 we show the same thing in terms of chi-square over degrees of freedom ratios (reluctantly since we do not like such ratios). As discussed in section 8.2 we see, perhaps even more clearly, that for few degrees of freedom the ratios may be very high while for large number of degrees of freedom this is not the case for the same confidence level.

### 8.11 Even Number of Degrees of Freedom

With even $n$ the power of $z$ in the last integral in the formula for $F\left(x_{\alpha}\right)$ above is an integer. From standard integral tables we find

$$
\int x^{m} e^{a x} d x=e^{a x} \sum_{r=0}^{m}(-1)^{r} \frac{m!x^{m-r}}{(m-r)!a^{r+1}}
$$

where, in our case, $a=-1$. Putting $m=\frac{n}{2}-1$ and using this integral we obtain

$$
\begin{aligned}
1-\alpha & =\frac{1}{\Gamma\left(\frac{n}{2}\right)} \int_{0}^{x_{\alpha} / 2} z^{\frac{n}{2}-1} e^{-z} d z=\left[\frac{1}{m!} e^{-z} \sum_{r=0}^{m}(-1)^{r} \frac{m!z^{m-r}}{(m-r)!(-1)^{r+1}}\right]_{0}^{\frac{x_{\alpha}}{2}} \\
& =1-e^{-\frac{x_{\alpha}}{2}} \sum_{r=0}^{m} \frac{x_{\alpha}^{m-r}}{2^{m-r}(m-r)!}=1-e^{-\frac{x_{\alpha}}{2}} \sum_{r=0}^{\frac{n}{2}-1} \frac{x_{\alpha}^{r}}{2^{r} r!}
\end{aligned}
$$

a result which indeed is identical to the formula for $P(n, x)$ for integer $n$ given on page 160 .

### 8.12 Odd Number of Degrees of Freedom

In the case of odd number of degrees of freedom we make the substitution $z^{2}=x$ yielding

$$
\begin{aligned}
1-\alpha & =F\left(x_{\alpha}\right)=\frac{1}{2 \Gamma\left(\frac{n}{2}\right)} \int_{0}^{x_{\alpha}}\left(\frac{x}{2}\right)^{\frac{n}{2}-1} e^{-\frac{x}{2}} d x=\frac{1}{2^{\frac{n}{2}} \Gamma\left(\frac{n}{2}\right)} \int_{0}^{\sqrt{x_{\alpha}}}\left(z^{2}\right)^{\frac{n}{2}-1} e^{-\frac{z^{2}}{2}} 2 z d z= \\
& =\frac{1}{2^{\frac{n}{2}-1} \Gamma\left(\frac{n}{2}\right)} \int_{0}^{\sqrt{x_{\alpha}}}\left(z^{2}\right)^{\frac{n-1}{2}} e^{-\frac{z^{2}}{2}} d z=\frac{1}{2^{m-\frac{1}{2}} \Gamma\left(m+\frac{1}{2}\right)} \int_{0}^{\sqrt{x_{\alpha}}} z^{2 m} e^{-\frac{z^{2}}{2}} d z
\end{aligned}
$$

where we have put $m=\frac{n-1}{2}$ which for odd $n$ is an integer. By partial integration in $m$ steps

$$
\begin{aligned}
\int z^{2 m} e^{-\frac{z^{2}}{2}} d z & =\int z^{2 m-1} z e^{-\frac{z^{2}}{2}} d z=-z^{2 m-1} e^{-\frac{z^{2}}{2}}+(2 m-1) \int z^{2 m-2} e^{-\frac{z^{2}}{2}} d z \\
\int z^{2 m-2} e^{-\frac{z^{2}}{2}} d z & =-z^{2 m-3} e^{-\frac{z^{2}}{2}}+(2 m-3) \int z^{2 m-4} e^{-\frac{z^{2}}{2}} d z \\
& \vdots \\
\int z^{4} e^{-\frac{z^{2}}{2}} d z & =-z^{3} e^{-\frac{z^{2}}{2}}+3 \int z^{2} e^{-\frac{z^{2}}{2}} d z \\
\int z^{2} e^{-\frac{z^{2}}{2}} d z & =-z e^{-\frac{z^{2}}{2}}+\int e^{-\frac{z^{2}}{2}} d z
\end{aligned}
$$

we obtain

$$
\int z^{2 m} e^{-\frac{z^{2}}{2}}=(2 m-1)!!\int e^{-\frac{z^{2}}{2}} d z-\sum_{r=0}^{m-1} \frac{(2 m-1)!!}{(2 r+1)!!} z^{2 r+1} e^{-\frac{z^{2}}{2}}
$$

Applying this to our case gives

$$
1-\alpha=\frac{1}{2^{m-\frac{1}{2}} \Gamma\left(m+\frac{1}{2}\right)}\left((2 m-1)!!\int_{0}^{\sqrt{x_{\alpha}}} e^{-\frac{z^{2}}{2}} d z-\left[\sum_{r=0}^{m-1} \frac{(2 m-1)!!}{(2 r+1)!!} z^{2 r+1} e^{-\frac{z^{2}}{2}}\right]_{0}^{\sqrt{x_{\alpha}}}\right)=
$$

$$
\begin{aligned}
& =\sqrt{\frac{2}{\pi}}\left(\int_{0}^{\sqrt{x_{\alpha}}} e^{-\frac{z^{2}}{2}} d z-\left[\sum_{r=0}^{m-1} \frac{1}{(2 r+1)!!} z^{2 r+1} e^{-\frac{z^{2}}{2}}\right]_{0}^{\sqrt{x_{\alpha}}}\right)= \\
& =2 G\left(\sqrt{x_{\alpha}}\right)-1-\sqrt{\frac{2 x_{\alpha}}{\pi}} e^{-\frac{x_{\alpha}}{2}} \sum_{r=0}^{m-1} \frac{x_{\alpha}^{r}}{(2 r+1)!!}
\end{aligned}
$$

where $G(z)$ is the integral of the standard normal distribution from $-\infty$ to $z$. Here we have used $\Gamma\left(m+\frac{1}{2}\right)=\frac{(2 m-1)!!}{2^{m}} \sqrt{\pi}$ in order to simplify the coefficient. This result may be compared to the formula given on page 160 for the incomplete Gamma function when the first argument is a half-integer.

### 8.13 Final Algorithm

The final algorithm to evaluate the probability content from $-\infty$ to $x$ for a chi-square distribution with $n$ degrees of freedom is

- For $n$ even:
- Put $m=\frac{n}{2}-1$.
- Set $u_{0}=1, s=0$ and $i=0$.
- For $i=0,1, \ldots, m$ set $s=s+u_{i}, i=i+1$ and $u_{i}=u_{i-1} \cdot \frac{x}{2 i}$.
- $\alpha=s \cdot e^{-\frac{x}{2}}$.
- For $n$ odd:
- Put $m=\frac{n-1}{2}$.
- Set $u_{0}=1, s=0$ and $i=0$.
- For $i=0,1, \ldots, m-1$ set $s=s+u_{i}, i=i+1$ and $u_{i}=u_{i-1} \cdot \frac{x}{2 i+1}$.
- $\alpha=2-2 G(\sqrt{x})+\sqrt{\frac{2 x}{\pi}} e^{-\frac{x}{2}} \cdot s$.


### 8.14 Chi Distribution

Sometimes, but less often, the chi distribution i.e. the distribution of $y=\sqrt{x}$ is used. By a simple change of variables this distribution is given by

$$
f(y)=\left|\frac{d x}{d y}\right| f\left(y^{2}\right)=2 y \cdot \frac{1}{2}\left(\frac{y^{2}}{2}\right)^{\frac{n}{2}-1} \frac{e^{-\frac{y^{2}}{2}}}{\Gamma\left(\frac{n}{2}\right)}=y^{n-1}\left(\frac{1}{2}\right)^{\frac{n}{2}-1} \frac{e^{-\frac{y^{2}}{2}}}{\Gamma\left(\frac{n}{2}\right)}
$$

In figure 8 the chi distribution is shown for $n$-values of $1,2,5$, and 10 . The mode of the distribution is at $\sqrt{n-1}$.

The cumulative function for the chi distribution becomes

$$
F(y)=\frac{\left(\frac{1}{2}\right)^{\frac{n}{2}-1}}{\Gamma\left(\frac{n}{2}\right)} \int_{0}^{y} x^{n-1} e^{-\frac{x^{2}}{2}} d x=P\left(\frac{n}{2}, \frac{y^{2}}{2}\right)
$$



Figure 8: Graph of chi distribution for some values of $n$
and algebraic moments are given by

$$
\mu_{k}^{\prime}=\frac{\left(\frac{1}{2}\right)^{\frac{n}{2}-1}}{\Gamma\left(\frac{n}{2}\right)} \int_{0}^{\infty} y^{k} y^{n-1} e^{-\frac{y^{2}}{2}} d y=\frac{2^{\frac{k}{2}} \Gamma\left(\frac{n}{2}+\frac{k}{2}\right)}{\Gamma\left(\frac{n}{2}\right)}
$$

## 9 Compound Poisson Distribution

### 9.1 Introduction

The compound Poisson distribution describes the branching process for Poisson variables and is given by

$$
p(r ; \mu, \lambda)=\sum_{n=0}^{\infty} \frac{(n \mu)^{r} e^{-n \mu}}{r!} \frac{\lambda^{n} e^{-\lambda}}{n!}
$$

where the integer variable $r \geq 0$ and the parameters $\mu$ and $\lambda$ are positive real quantities.

### 9.2 Branching Process

The distribution describes the branching of $n$ Poisson variables $n_{i}$ all with mean $\mu$ where $n$ is also distributed according to the Poisson distribution with mean $\lambda$ i.e.

$$
r=\sum_{i=1}^{n} n_{i} \quad \text { with } \quad p\left(n_{i}\right)=\frac{\mu^{n_{i}} e^{-\mu}}{n_{i}!} \quad \text { and } \quad p(n)=\frac{\lambda^{n} e^{-\lambda}}{n!}
$$

and thus

$$
p(r)=\sum_{n=0}^{\infty} p(r \mid n) p(n)
$$

Due to the so called addition theorem (see page 121) for Poisson variables with mean $\mu$ the sum of $n$ such variables are distributed as a Poisson variable with mean $n \mu$ and thus the distribution given above results.

### 9.3 Moments

The expectation value and variance of the Compound Poisson distribution are given by

$$
E(r)=\lambda \mu \quad \text { and } \quad V(r)=\lambda \mu(1+\mu)
$$

while higher moments gets slightly more complicated:

$$
\begin{aligned}
\mu_{3}= & \mu \lambda\left\{\mu+(\mu+1)^{2}\right\} \\
\mu_{4}= & \mu \lambda\left\{\mu^{3}+6 \mu^{2}+7 \mu+1+3 \mu \lambda(1+\mu)^{2}\right\} \\
\mu_{5}= & \mu \lambda\left\{\mu^{4}+10 \mu^{3}+25 \mu^{2}+15 \mu+1+10 \mu \lambda(\mu+1)\left(\mu+(1+\mu)^{2}\right)\right\} \\
\mu_{6}= & \mu \lambda\left\{\mu^{5}+15 \mu^{4}+65 \mu^{3}+90 \mu^{2}+31 \mu+1\right. \\
& \left.\quad+5 \mu \lambda\left(5 \mu^{4}+33 \mu^{3}+61 \mu^{2}+36 \mu+5\right)+15 \mu^{2} \lambda^{2}(\mu+1)^{3}\right\}
\end{aligned}
$$

### 9.4 Probability Generating Function

The probability generating function of the compound Poisson distribution is given by

$$
G(z)=\exp \left\{-\lambda+\lambda e^{-\mu+\mu z}\right\}
$$

This is easily found by using the rules for branching processes where the probability generating function (p.g.f.) is given by

$$
G(z)=G_{P}\left(G_{P}(z)\right)
$$

where $G_{P}(z)$ is the p.g.f. for the Poisson distribution.

### 9.5 Random Number Generation

Using the basic definition we may proceed by first generate a random number $n$ from a Poisson distribution with mean $\lambda$ and then another one with mean $n \mu$.

For fixed $\mu$ and $\lambda$ it is, however, normally much faster to prepare a cumulative vector for values ranging from zero up to the point where computer precision gives unity and then use this vector for random number generation. Using a binary search technique one may allow for quite long vectors giving good precision without much loss in efficiency.

## 10 Double-Exponential Distribution

### 10.1 Introduction

The Double-exponential distribution is given by

$$
f(x ; \mu, \lambda)=\frac{\lambda}{2} e^{-\lambda|x-\mu|}
$$

where the variable $x$ is a real number as is the location parameter $\mu$ while the parameter $\lambda$ is a real positive number.

The distribution is sometimes called the Laplace distribution after the french astronomer, mathematician and physicist marquis Pierre Simon de Laplace (1749-1827). It is a symmetric distribution whose tails fall off less sharply than the Gaussian distribution but faster than the Cauchy distribution. It has a cusp, discontinuous first derivative, at $x=\mu$.

The distribution has an interesting feature inasmuch as the best estimator for the mean $\mu$ is the median and not the sample mean. See further the discussion in section 7 on the Cauchy distribution where the Double-exponential distribution is discussed in some detail.

### 10.2 Moments

For the Double-exponential distribution central moments are more easy to determine than algebraic moments (the mean is $\mu_{1}^{\prime}=\mu$ ). They are given by

$$
\begin{aligned}
\mu_{n} & =\int_{-\infty}^{\infty}(x-\mu)^{n} f(x) d x=\frac{\lambda}{2}\left\{\int_{-\infty}^{\mu}(x-\mu)^{n} e^{-\lambda(\mu-x)}+\int_{\mu}^{\infty}(x-\mu)^{n} e^{-\lambda(x-\mu)}\right\}= \\
& =\frac{1}{2}\left\{\int_{-\infty}^{0}\left(-\frac{y}{\lambda}\right)^{n} e^{-y} d y+\int_{0}^{\infty}\left(\frac{y}{\lambda}\right)^{n} e^{-y} d y\right\}=\frac{n!}{2 \lambda^{n}}+(-1)^{n} \frac{n!}{2 \lambda^{n}}
\end{aligned}
$$

i.e. odd moments vanish as they should due to the symmetry of the distribution and even moments are given by the simple relation $\mu_{n}=n!/ \lambda^{n}$. From this one easily finds that the coefficient of skewness is zero and the coefficient of kurtosis 3 .

If required algebraic moments may be calculated from the central moments especially the lowest order algebraic moments become

$$
\mu_{1}^{\prime}=\mu, \quad \mu_{2}^{\prime}=\frac{2}{\lambda^{2}}+\mu^{2}, \quad \mu_{3}^{\prime}=\frac{6 \mu}{\lambda^{2}}+\mu^{3}, \quad \text { and } \quad \mu_{4}^{\prime}=\frac{24}{\lambda^{4}}+\frac{12 \mu^{2}}{\lambda^{2}}+\mu^{4}
$$

but more generally

$$
\mu_{n}^{\prime}=\sum_{r=0}^{n}\binom{n}{r} \mu^{r} \mu_{n-r}
$$

### 10.3 Characteristic Function

The characteristic function which generates central moments is given by

$$
\phi_{x-\mu}(t)=\frac{\lambda^{2}}{\lambda^{2}+t^{2}}
$$

from which we may find the characteristic function which generates algebraic moments

$$
\phi_{x}(t)=E\left(e^{\imath t x}\right)=e^{\imath t \mu} E\left(e^{\imath t(x-\mu)}\right)=e^{\imath t \mu} \phi_{x-\mu}(t)=e^{\imath t \mu} \frac{\lambda^{2}}{\lambda^{2}+t^{2}}
$$

Sometimes an alternative which generates the sequence $\mu_{1}^{\prime}, \mu_{2}, \mu_{3}, \ldots$ is given as

$$
\phi(t)=\imath t \mu+\frac{\lambda^{2}}{\lambda^{2}+t^{2}}
$$

### 10.4 Cumulative Function

The cumulative function, or distribution function, for the Double-exponential distribution is given by

$$
F(x)= \begin{cases}\frac{1}{2} e^{-\lambda(\mu-x)} & \text { if } \quad x \leq \mu \\ 1-\frac{1}{2} e^{-\lambda(x-\mu)} & \text { if } \quad x>\mu\end{cases}
$$

From this we see not only the obvious that the median is at $x=\mu$ but also that the lower and upper quartile is located at $\mu \mp \ln 2 / \lambda$.

### 10.5 Random Number Generation

Given a uniform random number between zero and one in $\xi$ a random number from a Double-exponential distribution is given by solving the equation $F(x)=\xi$ for $x$ giving

$$
\begin{array}{lll}
\text { For } & \xi \leq \frac{1}{2} & x=\mu+\ln (2 \xi) / \lambda \\
\text { for } & \xi>\frac{1}{2} & x=\mu-\ln (2-2 \xi) / \lambda
\end{array}
$$

## 11 Doubly Non-Central F-Distribution

### 11.1 Introduction

If $x_{1}$ and $x_{2}$ are independently distributed according to two non-central chi-square distributions with $n_{1}$ and $n_{2}$ degrees of freedom and non-central parameters $\lambda_{1}$ and $\lambda_{2}$, respectively, then the variable

$$
F^{\prime}=\frac{x_{1} / n_{1}}{x_{2} / n_{2}}
$$

is said to have a doubly non-central $F$-distribution with $n_{1}, n_{2}$ degrees of freedom (positive integers) and non-centrality parameters $\lambda_{1}, \lambda_{2}$ (both $\geq 0$ ).

This distribution may be written as

$$
f\left(x ; n_{1}, n_{2}, \lambda_{1}, \lambda_{2}\right)=\frac{n_{1}}{n_{2}} e^{-\frac{\lambda}{2}} \sum_{r=0}^{\infty} \sum_{s=0}^{\infty} \frac{\left(\frac{\lambda_{1}}{2}\right)^{r}}{r!} \frac{\left(\frac{\lambda_{2}}{2}\right)^{s}}{s!} \frac{\left(\frac{n_{1} x}{n_{2}}\right)^{\frac{n_{1}}{2}+r-1}}{\left(1+\frac{n_{1} x}{n_{2}}\right)^{\frac{n}{2}+r+s}} \frac{1}{B\left(\frac{n_{1}}{2}+r, \frac{n_{2}}{2}+s\right)}
$$

where we have put $n=n_{1}+n_{2}$ and $\lambda=\lambda_{1}+\lambda_{2}$. For $\lambda_{2}=0$ we obtain the (singly) non-central $F$-distribution (see section 32) and if also $\lambda_{1}=0$ we are back to the ordinary variance ratio, or $F$-, distribution (see section 16 ).

With four parameters a variety of shapes are possible. As an example figure 9 shows the doubly non-central $F$-distribution for the case with $n_{1}=10, n_{2}=5$ and $\lambda_{1}=10$ varying $\lambda_{2}$ from zero (an ordinary non-central $F$-distribution) to five.


Figure 9: Examples of doubly non-central $F$-distributions

### 11.2 Moments

Algebraic moments of this distributions become

$$
\begin{aligned}
E\left(x^{k}\right) & =\left(\frac{n_{2}}{n_{1}}\right)^{k} e^{-\frac{\lambda}{2}} \sum_{r=0}^{\infty} \frac{\left(\frac{\lambda_{1}}{2}\right)^{r}}{r!} \frac{\Gamma\left(\frac{n_{1}}{2}+r+k\right)}{\Gamma\left(\frac{n_{1}}{2}+r\right)} \sum_{s=0}^{\infty} \frac{\left(\frac{\lambda_{2}}{2}\right)^{s}}{s!} \frac{\Gamma\left(\frac{n_{2}}{2}+s-k\right)}{\Gamma\left(\frac{n_{2}}{2}+s\right)}= \\
& =\left(\frac{n_{2}}{n_{1}}\right)^{k} e^{-\frac{\lambda}{2}} \sum_{r=0}^{\infty} \frac{\left(\frac{\lambda_{1}}{2}\right)^{r}}{r!}\left(\frac{n_{1}}{2}+r+k-1\right) \cdots\left(\frac{n_{1}}{2}+r\right) .
\end{aligned}
$$

$$
\cdot \sum_{s=0}^{\infty} \frac{\left(\frac{\lambda_{2}}{2}\right)^{s}}{s!}\left(\frac{n_{2}}{2}+s-1\right)^{-1} \cdots\left(\frac{n_{2}}{2}+s-k\right)^{-1}
$$

The $r$-sum involved, with a polynomial in the numerator, is quite easily solvable giving similar expressions as for the (singly) non-central $F$-distribution. The $s$-sums, however, with polynomials in the denominator give rise to confluent hypergeometric functions $M(a, b ; x)$ (see appendix B). Lower order algebraic moments are given by

$$
\begin{aligned}
& E(x)= e^{-\frac{\lambda_{2}}{2}} \frac{n}{m} \cdot \frac{m+\lambda}{n-2} M\left(\frac{n-2}{2}, \frac{n}{2} ; \frac{\lambda_{2}}{2}\right) \\
& E\left(x^{2}\right)= e^{-\frac{\lambda_{2}}{2}}\left(\frac{n}{m}\right)^{2} \frac{\lambda^{2}+(2 \lambda+m)(m+2)}{(n-2)(n-4)} M\left(\frac{n-4}{2}, \frac{n}{2} ; \frac{\lambda_{2}}{2}\right) \\
& E\left(x^{3}\right)= e^{-\frac{\lambda_{2}}{2}}\left(\frac{n}{m}\right)^{3} \frac{\lambda^{3}+3(m+4) \lambda^{2}+(3 \lambda+m)(m+2)(m+4)}{(n-2)(n-4)(n-6)} M\left(\frac{n-6}{2}, \frac{n}{2} ; \frac{\lambda_{2}}{2}\right) \\
& E\left(x^{4}\right)=e^{-\frac{\lambda_{2}}{2}}\left(\frac{n}{m}\right)^{4} \frac{\lambda^{4}+(m+6)\left\{4 \lambda^{3}+(m+4)\left[6 \lambda^{2}+(4 \lambda+m)(m+2)\right]\right\}}{(n-2)(n-4)(n-6)(n-8)} . \\
& \cdot M\left(\frac{n-8}{2}, \frac{n}{2} ; \frac{\lambda_{2}}{2}\right)
\end{aligned}
$$

### 11.3 Cumulative Distribution

The cumulative, or distribution, function may be deduced by simple integration

$$
\begin{aligned}
F(x) & =\frac{n_{1}}{n_{2}} e^{-\frac{\lambda}{2}} \sum_{r=0}^{\infty} \sum_{s=0}^{\infty} \frac{\left(\frac{\lambda_{1}}{2}\right)^{r}}{r!} \frac{\left(\frac{\lambda_{2}}{2}\right)^{s}}{s!} \frac{1}{B\left(\frac{n_{1}}{2}+r, \frac{n_{2}}{2}+s\right)} \int_{0}^{x} \frac{\left(\frac{n_{1} u}{n_{2}}\right)^{\frac{n_{1}}{2}+r-1}}{\left(1+\frac{n_{1} u}{n_{2}}\right)^{\frac{n}{2}+r+s}} d u= \\
& =e^{-\frac{\lambda}{2}} \sum_{r=0}^{\infty} \sum_{s=0}^{\infty} \frac{\left(\frac{\lambda_{1}}{2}\right)^{r}}{r!} \frac{\left(\frac{\lambda_{2}}{2}\right)^{s}}{s!} \frac{B_{q}\left(\frac{n_{1}}{2}+r, \frac{n_{2}}{2}+s\right)}{B\left(\frac{n_{1}}{2}+r, \frac{n_{2}}{2}+s\right)}= \\
& =e^{-\frac{\lambda}{2}} \sum_{r=0}^{\infty} \sum_{s=0}^{\infty} \frac{\left(\frac{\lambda_{1}}{2}\right)^{r}}{r!} \frac{\left(\frac{\lambda_{2}}{2}\right)^{s}}{s!} I_{q}\left(\frac{n_{1}}{2}+r, \frac{n_{2}}{2}+s\right)
\end{aligned}
$$

with

$$
q=\frac{\frac{n_{1} x}{n_{2}}}{1+\frac{n_{1} x}{n_{2}}}
$$

### 11.4 Random Number Generation

Random numbers from a doubly non-central $F$-distribution is easily obtained using the definition in terms of the ratio between two independent random numbers from non-central chi-square distributions. This ought to be sufficient for most applications but if needed more efficient techniques may easily be developed e.g. using more general techniques.

## 12 Doubly Non-Central $t$-Distribution

### 12.1 Introduction

If $x$ and $y$ are independent and $x$ is normally distributed with mean $\delta$ and unit variance while $y$ is distributed according a non-central chi-square distribution with $n$ degrees of freedom and non-centrality parameter $\lambda$ then the variable

$$
t=x / \sqrt{y / n}
$$

is said to have a doubly non-central t-distribution with $n$ degrees of freedom (positive integer) and non-centrality parameters $\delta$ and $\lambda$ (with $\lambda \geq 0$ ).

This distribution may be expressed as

$$
f(t ; n, \delta, \lambda)=\frac{e^{-\frac{\delta^{2}}{2}} e^{-\frac{\lambda}{2}}}{\sqrt{n \pi}} \sum_{r=0}^{\infty} \frac{\left(\frac{\lambda}{2}\right)^{r}}{r!} \frac{1}{\Gamma\left(\frac{n}{2}+r\right)} \sum_{s=0}^{\infty} \frac{(t \delta)^{s}}{s!\left(\frac{n}{2}\right)^{\frac{s}{2}}}\left(1+\frac{t^{2}}{n}\right)^{-\left(\frac{n+s+1}{2}+r\right)} \Gamma\left(\frac{n+s+1}{2}+r\right)
$$

For $\lambda=0$ we obtain the (singly) non-central $t$-distribution (see section 33) and if also $\delta=0$ we are back to the ordinary $t$-distribution (see section 38).

Examples of doubly non-central $t$-distributions are shown in figure 9 for the case with $n=10$ and $\delta^{2}=5$ varying $\lambda$ from zero (an ordinary non-central $t$-distribution) to ten.


Figure 10: Examples of doubly non-central $t$-distributions

### 12.2 Moments

Algebraic moments may be deduced from the expression

$$
\begin{aligned}
E\left(t^{k}\right) & =\frac{e^{-\frac{\delta^{2}}{2}} e^{-\frac{\lambda}{2}}}{\sqrt{n \pi}} \sum_{r=0}^{\infty} \frac{\left(\frac{\lambda}{2}\right)^{r}}{r!} \frac{1}{\Gamma\left(\frac{n}{2}+r\right)} \sum_{s=0}^{\infty} \frac{\delta^{s}}{s!n^{\frac{s}{2}}} 2^{\frac{s}{2}} \Gamma\left(\frac{n+s+1}{2}+r\right) \int_{-\infty}^{\infty} \frac{t^{s+k}}{\left(1+\frac{t^{2}}{n}\right)^{\frac{n+s+1}{2}+r}} d t= \\
& =\frac{e^{-\frac{\delta^{2}}{2}} e^{-\frac{\lambda}{2}}}{\sqrt{\pi}} \sum_{r=0}^{\infty} \frac{\left(\frac{\lambda}{2}\right)^{r}}{r!} \frac{1}{\Gamma\left(\frac{n}{2}+r\right)} \sum_{s=0}^{\infty} \frac{\delta^{s}}{s!} 2^{\frac{s}{2}} n^{\frac{k}{2}} \Gamma\left(\frac{s+k+1}{2}\right) \Gamma\left(\frac{n-k}{2}+r\right)
\end{aligned}
$$

where the sum should be taken for even values of $s+k i . e$. for even (odd) orders sum only over even (odd) $s$-values.

Differing between moments of even and odd order the following expressions for lower order algebraic moments of the doubly non-central $t$-distribution may be expressed in terms of the confluent hypergeometric function $M(a, b ; x)$ (see appendix B for details) as

$$
\begin{aligned}
E(t) & =\delta \sqrt{\frac{n}{2}} e^{-\frac{\lambda}{2}} \frac{\Gamma\left(\frac{n-1}{2}\right)}{\Gamma\left(\frac{n}{2}\right)} M\left(\frac{n-1}{2}, \frac{n}{2} ; \frac{\lambda}{2}\right) \\
E\left(t^{2}\right) & =\frac{n e^{-\frac{\lambda}{2}}\left(\delta^{2}+1\right)}{n-2} M\left(\frac{n-2}{2}, \frac{n}{2} ; \frac{\lambda}{2}\right) \\
E\left(t^{3}\right) & =\delta\left(\delta^{2}+3\right) \sqrt{\frac{n^{3}}{8}} e^{-\frac{\lambda}{2}} \frac{\Gamma\left(\frac{n-3}{2}\right)}{\Gamma\left(\frac{n}{2}\right)} M\left(\frac{n-3}{2}, \frac{n}{2} ; \frac{\lambda}{2}\right) \\
E\left(t^{4}\right) & =\frac{n^{2}}{(n-2)(n-4)}\left(\delta^{4}+6 \delta^{2}+3\right) e^{-\frac{\lambda}{2}} M\left(\frac{n-4}{2}, \frac{n}{2} ; \frac{\lambda}{2}\right)
\end{aligned}
$$

### 12.3 Cumulative Distribution

The cumulative, distribution, function is given by

$$
\begin{aligned}
F(t) & =\frac{e^{-\frac{\delta^{2}}{2}} e^{-\frac{\lambda}{2}}}{\sqrt{n \pi}} \sum_{r=0}^{\infty} \frac{\left(\frac{\lambda}{2}\right)^{r}}{r!} \frac{1}{\Gamma\left(\frac{n}{2}+r\right)} \sum_{s=0}^{\infty} \frac{\delta^{s}}{s!\left(\frac{n}{2}\right)^{\frac{s}{2}}} \Gamma\left(\frac{n+s+1}{2}+r\right) \int_{-\infty}^{t} \frac{u^{s}}{\left(1+\frac{u^{2}}{n}\right)^{\frac{n+s+1}{2}+r}} d u= \\
& =\frac{e^{-\frac{\delta^{2}}{2}} e^{-\frac{\lambda}{2}}}{\sqrt{\pi}} \sum_{r=0}^{\infty} \frac{\left(\frac{\lambda}{2}\right)^{r}}{r!} \sum_{s=0}^{\infty} \frac{\delta^{s}}{s!} 2^{\frac{s}{2}-1} \Gamma\left(\frac{s+1}{2}\right)\left\{s_{1}+s_{2} I_{q}\left(\frac{s+1}{2}, \frac{n}{2}+r\right)\right\}
\end{aligned}
$$

where $q=\left(t^{2} / n\right) /\left(1+t^{2} / n\right)$ and $s_{1}, s_{2}$ are signs differing between cases with positive or negative $t$ as well as odd or even $s$ in the summation. More specific, the sign $s_{1}$ is -1 if $s$ is odd and +1 if it is even while $s_{2}$ is +1 unless $t<0$ and $s$ is even in which case it is -1 .

### 12.4 Random Number Generation

Random numbers from a doubly non-central $t$-distribution is easily obtained with the definition given above using random numbers from a normal distribution and a non-central chi-square distribution. This ought to be sufficient for most applications but if needed more efficient techniques may easily be developed e.g. using more general techniques.

## 13 Error Function

### 13.1 Introduction

A function, related to the probability content of the normal distribution, which often is referred to is the error function

$$
\operatorname{erf} z=\frac{2}{\sqrt{\pi}} \int_{0}^{z} e^{-t^{2}} d t
$$

and its complement

$$
\operatorname{erfc} z=\frac{2}{\sqrt{\pi}} \int_{z}^{\infty} e^{-t^{2}} d t=1-\operatorname{erf} z
$$

These functions may be defined for complex arguments, for many relations concerning the error function see [27], but here we are mainly interested in the function for real positive values of $z$. However, sometimes one may still want to define the function values for negative real values of $z$ using symmetry relations

$$
\begin{aligned}
\operatorname{erf}(-z) & =-\operatorname{erf}(z) \\
\operatorname{erfc}(-z) & =1-\operatorname{erf}(-z)=1+\operatorname{erf}(z)
\end{aligned}
$$

### 13.2 Probability Density Function

As is seen the error function erf is a distribution (or cumulative) function and the corresponding probability density function is given by

$$
f(z)=\frac{2}{\sqrt{\pi}} e^{-z^{2}}
$$

If we make the transformation $z=(x-\mu) / \sigma \sqrt{2}$ we obtain a folded normal distribution

$$
f(x ; \mu, \sigma)=\frac{1}{\sigma} \sqrt{\frac{2}{\pi}} e^{-\frac{1}{2}\left(\frac{x-\mu}{\sigma}\right)^{2}}
$$

where the function is defined for $x>\mu$ corresponding to $z>0, \mu$ may be any real number while $\sigma>0$.

This implies that $\operatorname{er} f(z / \sqrt{2})$ is equal to the symmetric integral of a standard normal distribution between $-z$ and $z$.

The error function may also be expressed in terms of the incomplete Gamma function

$$
\operatorname{erf} x=\frac{2}{\sqrt{\pi}} \int_{0}^{x} e^{-t^{2}} d t=P\left(\frac{1}{2}, x^{2}\right)
$$

defined for $x \geq 0$.

## 14 Exponential Distribution

### 14.1 Introduction

The exponential distribution is given by

$$
f(x ; \alpha)=\frac{1}{\alpha} e^{-\frac{x}{\alpha}}
$$

where the variable $x$ as well as the parameter $\alpha$ is positive real quantities.
The exponential distribution occur in many different connections such as the radioactive or particle decays or the time between events in a Poisson process where events happen at a constant rate.

### 14.2 Cumulative Function

The cumulative (distribution) function is

$$
F(x)=\int_{0}^{x} f(x) d x=1-e^{-\frac{x}{\alpha}}
$$

and it is thus straightforward to calculate the probability content in any given situation. E.g. we find that the median and the lower and upper quartiles are at

$$
\mathcal{M}=\alpha \ln 2 \approx 0.693 \alpha, \quad \mathcal{Q}_{1}=-\alpha \ln \frac{3}{4} \approx 0.288 \alpha, \quad \text { and } \quad \mathcal{Q}_{3}=\alpha \ln 4 \approx 1.386 \alpha
$$

### 14.3 Moments

The expectation value, variance, and lowest order central moments are given by

$$
\begin{gathered}
E(x)=\alpha, \quad V(x)=\alpha^{2}, \quad \mu_{3}=2 \alpha^{3}, \quad \mu_{4}=9 \alpha^{4}, \\
\mu_{5}=44 \alpha^{5}, \quad \mu_{6}=265 \alpha^{6}, \quad \mu_{7}=1854 \alpha^{7}, \quad \text { and } \mu_{8}=14833 \alpha^{8}
\end{gathered}
$$

More generally algebraic moments are given by

$$
\mu_{n}^{\prime}=\alpha^{n} n!
$$

Central moments thereby becomes

$$
\mu_{n}=\alpha^{n} n!\sum_{m=0}^{n} \frac{(-1)^{m}}{m!} \rightarrow \frac{\alpha^{n} n!}{e}=\frac{\mu_{n}^{\prime}}{e} \quad \text { when } \quad n \rightarrow \infty
$$

the approximation is, in fact, quite good already for $n=5$ where the absolute error is $0.146 \alpha^{5}$ and the relative error $0.3 \%$.

### 14.4 Characteristic Function

The characteristic function of the exponential distribution is given by

$$
\phi(t)=E\left(e^{\imath t x}\right)=\frac{1}{\alpha} \int_{0}^{\infty} e^{\left(\imath t-\frac{1}{\alpha}\right) x} d x=\frac{1}{1-\imath t \alpha}
$$

### 14.5 Random Number Generation

The most common way to achieve random numbers from an exponential distribution is to use the inverse to the cumulative distribution such that

$$
x=F^{-1}(\xi)=-\alpha \ln (1-\xi)=-\alpha \ln \xi^{\prime}
$$

where $\xi$ is a uniform random number between zero and one (aware not to include exactly zero in the range) and so is, of course, also $\xi^{\prime}=1-\xi$.

There are, however, alternatives some of which may be of some interest and useful if the penalty of using the logarithm would be big on any system [28].

### 14.5.1 Method by von Neumann

The first of these is due to J. von Neumann [29] and is as follows (with different $\xi$ 's denoting independent uniform random numbers between zero and one)
i Set $a=0$.
ii Generate $\xi$ and put $\xi_{0}=\xi$.
iii Generate $\xi^{*}$ and if $\xi^{*}<\xi$ then go to vi.
iv Generate $\xi$ and if $\xi<\xi^{*}$ then go to iii.
v Put $a=a+1$ and go to ii.
vi Put $x=\alpha\left(a+\xi_{0}\right)$ as a random number from an exponential distribution.

### 14.5.2 Method by Marsaglia

The second technique is attributed to G. Marsaglia [30].

- Prepare

$$
p_{n}=1-e^{-n} \quad \text { and } \quad q_{n}=\frac{1}{e-1}\left(\frac{1}{1!}+\frac{1}{2!}+\cdots+\frac{1}{n!}\right)
$$

for $n=1,2, \ldots$ until the largest representable fraction below one is exceeded in both vectors.
i Put $i=0$ and generate $\xi$.
ii If $\xi>p_{i+1}$ put $i=i+1$ and perform this step again.
iii Put $k=1$, generate $\xi$ and $\xi^{*}$, and set $\xi_{\text {min }}=\xi^{*}$.
iv If $\xi \leq q_{k}$ then go to vi else set $k=k+1$.
$\mathbf{v}$ Generate a new $\xi^{*}$ and if $\xi^{*}<\xi_{\text {min }}$ set $\xi_{\text {min }}=\xi^{*}$ and go to iv.
vi Put $x=\alpha\left(i+\xi_{\min }\right)$ as an exponentially distributed random number.

### 14.5.3 Method by Ahrens

The third method is due to J. H. Ahrens [28]

- Prepare

$$
q_{n}=\frac{\ln 2}{1!}+\frac{(\ln 2)^{2}}{2!}+\cdots+\frac{(\ln 2)^{n}}{n!}
$$

for $n=1,2, \ldots$ until the largest representable fraction less than one is exceeded.
i Put $a=0$ and generate $\xi$.
ii If $\xi<\frac{1}{2}$ set $a=a+\ln 2=a+q_{1}, \quad \xi=2 \xi$ and perform this step again.
iii Set $\xi=2 \xi-1$ and if $\xi \leq \ln 2=q_{1}$ then exit with $x=\alpha(a+\xi)$ else put $i=2$ and generate $\xi_{\text {min }}$.
iv Generate $\xi$ and put $\xi_{\min }=\xi$ if $\xi<\xi_{\min }$ then if $\xi>q_{i}$ put $i=i+1$ and perform this step again else exit with $x=\alpha\left(a+q_{1} \xi_{\text {min }}\right)$.

Of these three methods the method by Ahrens is the fastest. This is much due to the fact that the average number of uniform random numbers consumed by the three methods is 1.69 for Ahrens, 3.58 for Marsaglia, and 4.31 for von Neumann. The method by Ahrens is often as fast as the direct logarithm method on many computers.

## 15 Extreme Value Distribution

### 15.1 Introduction

The extreme value distribution is given by

$$
f(x ; \mu, \sigma)=\frac{1}{\sigma} \exp \left\{\mp \frac{x-\mu}{\sigma}-e^{\mp \frac{x-\mu}{\sigma}}\right\}
$$

where the upper sign is for the maximum and the lower sign for the minimum (often only the maximum is considered). The variable $x$ and the parameter $\mu$ (the mode) are real numbers while $\sigma$ is a positive real number. The distribution is sometimes referred to as the Fisher-Tippett distribution (type I), the log-Weibull distribution, or the Gumbel distribution after E. J. Gumbel (1891-1966).

The extreme value distribution gives the limiting distribution for the largest or smallest elements of a set of independent observations from a distribution of exponential type (normal, gamma, exponential, etc.).

A normalized form, useful to simplify calculations, is obtained by making the substitution to the variable $z= \pm \frac{x-\mu}{\sigma}$ which has the distribution

$$
g(z)=e^{-z-e^{-z}}
$$

In figure 11 we show the distribution in this normalized form. The shape corresponds to the case for the maximum value while the distribution for the minimum value would be mirrored in $z=0$.


Figure 11: The normalized Extreme Value Distribution

### 15.2 Cumulative Distribution

The cumulative distribution for the extreme value distribution is given by

$$
F(x)=\int_{-\infty}^{x} f(u) d u=\int_{-\infty}^{ \pm \frac{x-\mu}{\sigma}} g(z) d z=G\left( \pm \frac{x-\mu}{\sigma}\right)
$$

where $G(z)$ is the cumulative function of $g(z)$ which is given by

$$
G(z)=\int_{-\infty}^{z} e^{-u-e^{-u}} d u=\int_{e^{-z}}^{\infty} e^{-y} d y=e^{-e^{-z}}
$$

where we have made the substitution $y=e^{-u}$ in simplifying the integral. From this, and using $x=\mu \pm \sigma z$, we find the position of the median and the lower and upper quartile as

$$
\begin{aligned}
\mathcal{M} & =\mu \mp \sigma \ln \ln 2 \approx \mu \pm 0.367 \sigma, \\
\mathcal{Q}_{1} & =\mu \mp \sigma \ln \ln 4 \approx \mu \mp 0.327 \sigma, \quad \text { and } \\
\mathcal{Q}_{3} & =\mu \mp \sigma \ln \ln \frac{4}{3} \approx \mu \pm 1.246 \sigma
\end{aligned}
$$

### 15.3 Characteristic Function

The characteristic function of the extreme value distribution is given by

$$
\begin{aligned}
\phi(t) & =E\left(e^{\imath t x}\right)=\int_{-\infty}^{\infty} e^{\imath t x} \frac{1}{\sigma} \exp \left\{\mp \frac{x-\mu}{\sigma}-e^{\mp \frac{x-\mu}{\sigma}}\right\} d x= \\
& =\mp \int_{0}^{\infty} e^{\imath t(\mu \mp \sigma \ln z)} z e^{-z}\left(\mp \frac{\sigma d z}{z}\right)=e^{\imath t \mu} \int_{0}^{\infty} z^{\mp \imath t \sigma} e^{-z} d z=e^{\imath t \mu} \Gamma(1 \mp \imath t \sigma)
\end{aligned}
$$

where we have made the substitution $z=\exp (\mp(x-\mu) / \sigma)$ i.e. $x=\mu \mp \sigma \ln z$ and thus $d x=\mp \sigma d z / z$ to achieve an integral which could be expressed in terms of the Gamma function (see section 42.2). As a check we may calculate the first algebraic moment, the mean, by

$$
\mu_{1}^{\prime}=\left.\frac{1}{\imath} \frac{d \phi(t)}{d t}\right|_{t=0}=\frac{1}{\imath}[\imath \mu \Gamma(1)+\Gamma(1) \psi(1)(\mp \imath \sigma)]=\mu \pm \sigma \gamma
$$

Here $\psi(1)=-\gamma$ is the digamma function, see section 42.3, and $\gamma$ is Euler's constant. Similarly higher moments could be obtained from the characteristic function or, perhaps even easier, we may find cumulants from the cumulant generating function $\ln \phi(t)$. In the section below, however, moments are determined by a more direct technique.

### 15.4 Moments

Algebraic moments for $f(x)$ are given by

$$
E\left(x^{n}\right)=\int_{-\infty}^{\infty} x^{n} f(x) d x=\int_{-\infty}^{\infty}(\mu \pm \sigma z)^{n} g(z) d z
$$

which are related to moments of $g(z)$

$$
E\left(z^{n}\right)=\int_{-\infty}^{\infty} z^{n} e^{-z-e^{-z}} d z=\int_{0}^{\infty}(-\ln y)^{n} e^{-y} d y
$$

The first six such integrals, for $n$ values from 1 to 6 , are given by

$$
\begin{aligned}
& \int_{0}^{\infty}(-\ln x) e^{-x} d x=\gamma \\
& \int_{0}^{\infty}(-\ln x)^{2} e^{-x} d x=\gamma^{2}+\frac{\pi^{2}}{6} \\
& \int_{0}^{\infty}(-\ln x)^{3} e^{-x} d x= \gamma^{3}+\frac{\gamma \pi^{2}}{2}+2 \zeta_{3} \\
& \int_{0}^{\infty} \\
& \int_{0}^{\infty}(-\ln x)^{4} e^{-x} d x=\gamma^{4}+\gamma^{2} \pi^{2}+\frac{3 \pi^{4}}{20}+8 \gamma \zeta_{3} \\
& \int_{0}^{\infty}(-\ln x)^{5} e^{-x} d x=\gamma^{5}+\frac{5 \gamma^{3} \pi^{2}}{3}+\frac{3 \gamma \pi^{4}}{4}+20 \gamma^{2} \zeta_{3}+\frac{10 \pi^{2} \zeta_{3}}{3}+24 \zeta_{5} \\
& \int_{0}^{\infty}= \gamma^{6}+\frac{5 \gamma^{4} \pi^{2}}{2}+\frac{9 \gamma^{2} \pi^{4}}{4}+\frac{61 \pi^{6}}{168}+40 \gamma^{3} \zeta_{3}+ \\
& \int_{0}^{\infty}(-\ln x)^{6} e^{-x} d x+40 \zeta_{3}^{2}+144 \gamma \zeta_{5}
\end{aligned}
$$

corresponding to the six first algebraic moments of $g(z)$. Here $\gamma$ is Euler's (or EulerMascheroni) constant

$$
\gamma=\lim _{n \rightarrow \infty}\left(\sum_{k=1}^{n} \frac{1}{k}-\ln n\right)=0.5772156649015328606065120 \ldots
$$

and $\zeta_{n}$ is a short hand notation for Riemann's zeta-function $\zeta(n)$ given by

$$
\zeta(z)=\sum_{k=1}^{\infty} \frac{1}{k^{z}}=\frac{1}{\Gamma(z)} \int_{0}^{\infty} \frac{x^{z-1}}{e^{x}-1} d x \quad \text { for } \quad z>1
$$

(see also [31]). For $z$ an even integer we may use

$$
\zeta(2 n)=\frac{2^{2 n-1} \pi^{2 n}\left|B_{2 n}\right|}{(2 n)!} \quad \text { for } \quad n=1,2, \ldots
$$

where $B_{2 n}$ are the Bernoulli numbers given by $B_{2}=\frac{1}{6}, B_{4}=-\frac{1}{30}, B_{6}=\frac{1}{42}, B_{8}=-\frac{1}{30}$ etc (see table 4 on page 174 for an extensive table of the Bernoulli numbers). This implies $\zeta_{2}=\frac{\pi^{2}}{6}, \zeta_{4}=\frac{\pi^{4}}{90}, \zeta_{6}=\frac{\pi^{6}}{945}$ etc.

For odd integer arguments no similar relation as for even integers exists but evaluating the sum of reciprocal powers the two numbers needed in the calculations above are given
by $\zeta_{3}=1.20205690315959428540 \ldots$ and $\zeta_{5}=1.03692775514336992633 \ldots$. The number $\zeta_{3}$ is sometimes referred to as Apéry's constant after the person who in 1979 showed that it is an irrational number (but sofar it is not known if it is also transcendental) [32].

Using the algebraic moments of $g(z)$ as given above we may find the low order central moments of $g(z)$ as

$$
\begin{aligned}
\mu_{2} & =\frac{\pi^{2}}{6}=\zeta_{2} \\
\mu_{3} & =2 \zeta_{3} \\
\mu_{4} & =3 \pi^{4} / 20 \\
\mu_{5} & =\frac{10 \pi^{2} \zeta_{3}}{3}+24 \zeta_{5} \\
\mu_{6} & =\frac{61 \pi^{6}}{168}+40 \zeta_{3}^{2}
\end{aligned}
$$

and thus the coefficients of skewness $\gamma_{1}$ and kurtosis $\gamma_{2}$ are given by

$$
\begin{aligned}
\gamma_{1} & =\mu_{3} / \mu_{2}^{\frac{3}{2}}=12 \sqrt{6} \zeta_{3} / \pi^{3} \approx 1.13955 \\
\gamma_{2} & =\mu^{4} / \mu_{2}^{2}-3=2.4
\end{aligned}
$$

Algebraic moments of $f(x)$ may be found from this with some effort. Central moments are simpler being connected to those for $g(z)$ through the relation $\mu_{n}(x)=( \pm 1)^{n} \sigma^{n} \mu_{n}(z)$.

In particular the expectation value and the variance of $f(x)$ are given by

$$
\begin{aligned}
& E(x)=\mu \pm \sigma E(z)=\mu \pm \sigma \gamma \\
& V(x)=\sigma^{2} V(z)=\frac{\sigma^{2} \pi^{2}}{6}
\end{aligned}
$$

The coefficients of skewness (except for a sign $\pm 1$ ) and kurtosis are the same as for $g(z)$.

### 15.5 Random Number Generation

Using the expression for the cumulative distribution we may use a random number $\xi$, uniformly distributed between zero and one, to obtain a random number from the extreme value distribution by

$$
G(z)=e^{-e^{-z}}=\xi \quad \Rightarrow \quad z=-\ln (-\ln \xi)
$$

which gives a random number from the normalized function $g(z)$. A random number from $f(x)$ is then easily obtained by $x=\mu \pm \sigma z$.

## $16 \quad$ F-distribution

### 16.1 Introduction

The $F$-distribution is given by

$$
f(F ; m, n)=\frac{m^{\frac{m}{2}} n^{\frac{n}{2}} \Gamma\left(\frac{m+n}{2}\right)}{\Gamma\left(\frac{m}{2}\right) \Gamma\left(\frac{n}{2}\right)} \cdot \frac{F^{\frac{m}{2}-1}}{(m F+n)^{\frac{m+n}{2}}}=\frac{m^{\frac{m}{2}} n^{\frac{n}{2}}}{B\left(\frac{m}{2}, \frac{n}{2}\right)} \cdot \frac{F^{\frac{m}{2}-1}}{(m F+n)^{\frac{m+n}{2}}}
$$

where the parameters $m$ and $n$ are positive integers, degrees of freedom and the variable $F$ is a positive real number. The functions $\Gamma$ and $B$ are the usual Gamma and Beta functions. The distribution is often called the Fisher $F$-distribution, after the famous british statistician Sir Ronald Aylmer Fisher (1890-1962), sometimes the Snedecor $F$-distribution and sometimes the Fisher-Snedecor $F$-distribution. In figure 12 we show the $F$-distribution for low values of $m$ and $n$.


Figure 12: The $F$-distribution (a) for $m=10$ and $n=1,2, \ldots, 10$ and (b) for $m=$ $1,2, \ldots, 10$ and $n=10$

For $m \leq 2$ the distribution has its maximum at $F=0$ and is monotonically decreasing. Otherwise the distribution has the mode at

$$
F_{\text {mode }}=\frac{m-2}{m} \cdot \frac{n}{n+2}
$$

This distribution is also known as the variance-ratio distribution since it, as will be shown below, describes the distribution of the ratio of the estimated variances from two independent samples from normal distributions with equal variance.

### 16.2 Relations to Other Distributions

For $m=1$ we obtain a $t^{2}$-distribution, the distribution of the square of a variable distributed according to Student's $t$-distribution. As $n \rightarrow \infty$ the quantity $m F$ approaches a chi-square distribution with $m$ degrees of freedom.

For large values of $m$ and $n$ the $F$-distribution tends to a normal distribution. There are several approximations found in the literature all of which are better than a simpleminded standardized variable. One is

$$
z_{1}=\frac{\sqrt{2 n-1} \frac{m F}{n}-\sqrt{2 m-1}}{\sqrt{1+\frac{m F}{n}}}
$$

and an even better choice is

$$
z_{2}=\frac{F^{\frac{1}{3}}\left(1-\frac{2}{9 n}\right)-\left(1-\frac{2}{9 m}\right)}{\sqrt{\frac{2}{9 m}+F^{\frac{2}{3}} \cdot \frac{2}{9 n}}}
$$

For large values of $m$ and $n$ also the distribution in the variable $z=\frac{\ln F}{2}$, the distribution of which is known as the Fisher $z$-distribution, is approximately normal with mean $\frac{1}{2}\left(\frac{1}{n}-\frac{1}{m}\right)$ and variance $\frac{1}{2}\left(\frac{1}{m}+\frac{1}{n}\right)$. This approximation is, however, not as good as $z_{2}$ above.

## $16.3 \quad 1 / \mathrm{F}$

If $F$ is distributed according to the $F$-distribution with $m$ and $n$ degrees of freedom then $\frac{1}{F}$ has the $F$-distribution with $n$ and $m$ degrees of freedom. This is easily verified by a change of variables. Putting $G=\frac{1}{F}$ we have

$$
f(G)=\left|\frac{d F}{d G}\right| f(F)=\frac{1}{G^{2}} \cdot \frac{m^{\frac{m}{2}} n^{\frac{n}{2}}}{B\left(\frac{m}{2}, \frac{n}{2}\right)} \cdot \frac{\left(\frac{1}{G}\right)^{\frac{m}{2}-1}}{\left(\frac{m}{G}+n\right)^{\frac{m+n}{2}}}=\frac{m^{\frac{m}{2}} n^{\frac{n}{2}}}{B\left(\frac{m}{2}, \frac{n}{2}\right)} \cdot \frac{G^{\frac{n}{2}-1}}{(m+n G)^{\frac{m+n}{2}}}
$$

which is seen to be identical to a $F$-distribution with $n$ and $m$ degrees of freedom for $G=\frac{1}{F}$.

### 16.4 Characteristic Function

The characteristic function for the $F$-distribution may be expressed in terms of the confluent hypergeometric function $M$ (see section 43.3) as

$$
\phi(t)=E\left(e^{\imath F t}\right)=M\left(\frac{m}{2},-\frac{n}{2} ;-\frac{n}{m} \imath t\right)
$$

### 16.5 Moments

Algebraic moments are given by

$$
\mu_{r}^{\prime}=\frac{m^{\frac{m}{2}} n^{\frac{n}{2}}}{B\left(\frac{m}{2}, \frac{n}{2}\right)} \int_{0}^{\infty} \frac{F^{\frac{m}{2}-1+r}}{(m F+n)^{\frac{m+n}{2}}} d F=\left(\frac{m}{n}\right)^{\frac{m}{2}} \frac{1}{B\left(\frac{m}{2}, \frac{n}{2}\right)} \int_{0}^{\infty} \frac{F^{\frac{m}{2}-1+r}}{\left(\frac{m F}{n}+1\right)^{\frac{m+n}{2}}} d F=
$$

$$
\begin{aligned}
& =\left(\frac{m}{n}\right)^{\frac{m}{2}} \frac{1}{B\left(\frac{m}{2}, \frac{n}{2}\right)} \int_{0}^{\infty} \frac{\left(\frac{u n}{m}\right)^{\frac{m}{2}-1+r}}{(u+1)^{\frac{m+n}{2}}} \frac{n}{m} d u=\left(\frac{n}{m}\right)^{r} \frac{B\left(\frac{m}{2}+r, \frac{n}{2}-r\right)}{B\left(\frac{m}{2}, \frac{n}{2}\right)}= \\
& =\left(\frac{n}{m}\right)^{r} \cdot \frac{\Gamma\left(\frac{m}{2}+r\right) \Gamma\left(\frac{n}{2}-r\right)}{\Gamma\left(\frac{m}{2}\right) \Gamma\left(\frac{n}{2}\right)}
\end{aligned}
$$

and are defined for $r<\frac{n}{2}$. This may be written

$$
\mu_{r}^{\prime}=\left(\frac{n}{m}\right)^{r} \cdot \frac{\frac{m}{2}\left(\frac{m}{2}+1\right) \cdots\left(\frac{m}{2}+r-1\right)}{\left(\frac{n}{2}-r\right)\left(\frac{n}{2}-r+1\right) \cdots\left(\frac{n}{2}-1\right)}
$$

a form which may be more convenient for computations especially when $m$ or $n$ are large. A recursive formula to obtain the algebraic moments would thus be

$$
\mu_{r}^{\prime}=\mu_{r-1}^{\prime} \cdot\left(\frac{n}{m}\right) \cdot \frac{\frac{m}{2}+r-1}{\frac{n}{2}-r}
$$

starting with $\mu_{0}^{\prime}=1$.
The first algebraic moment, the mean, becomes

$$
E(F)=\frac{n}{n-2} \quad \text { for } \quad n>2
$$

and the variance is given by

$$
V(F)=\frac{2 n^{2}(m+n-2)}{m(n-2)^{2}(n-4)} \quad \text { for } \quad n>4
$$

### 16.6 F-ratio

Regard $F=\frac{u / m}{v / n}$ where $u$ and $v$ are two independent variables distributed according to the chi-square distribution with $m$ and $n$ degrees of freedom, respectively.

The independence implies that the joint probability function in $u$ and $v$ is given by the product of the two chi-square distributions

$$
f(u, v ; m, n)=\left(\frac{\left(\frac{u}{2}\right)^{\frac{m}{2}-1} e^{-\frac{u}{2}}}{2 \Gamma\left(\frac{m}{2}\right)}\right)\left(\frac{\left(\frac{v}{2}\right)^{\frac{n}{2}-1} e^{-\frac{v}{2}}}{2 \Gamma\left(\frac{n}{2}\right)}\right)
$$

If we change variables to $x=\frac{u / m}{v / n}$ and $y=v$ the distribution in $x$ and $y$ becomes

$$
f(x, y ; m, n)=\left\|\frac{\partial(u, v)}{\partial(x, y)}\right\| f(u, v ; m, n)
$$

The determinant of the Jacobian of the transformation is $\frac{y m}{n}$ and thus we have

$$
f(x, y ; m, n)=\frac{y m}{n}\left(\frac{\left(\frac{x y m}{n}\right)^{\frac{m}{2}-1} e^{-\frac{x y m}{2 n}}}{2^{\frac{m}{2}} \Gamma\left(\frac{m}{2}\right)}\right)\left(\frac{y^{\frac{n}{2}-1} e^{-\frac{y}{2}}}{2^{\frac{n}{2}} \Gamma\left(\frac{n}{2}\right)}\right)
$$

Finally, since we are interested in the marginal distribution in $x$ we integrate over $y$

$$
\begin{aligned}
f(x ; m, n) & =\int_{0}^{\infty} f(x, y ; m, n) d y=\frac{\left(\frac{m}{n}\right)^{\frac{m}{2}} x^{\frac{m}{2}-1}}{2^{\frac{m+n}{2}} \Gamma\left(\frac{m}{2}\right) \Gamma\left(\frac{n}{2}\right)} \int_{0}^{\infty} y^{\frac{m+n}{2}-1} e^{-\frac{y}{2}\left(\frac{x m}{n}+1\right)} d y= \\
& =\frac{\left(\frac{m}{n}\right)^{\frac{m}{2}} x^{\frac{m}{2}-1}}{2^{\frac{m+n}{2}} \Gamma\left(\frac{m}{2}\right) \Gamma\left(\frac{n}{2}\right)} \cdot \frac{2^{\frac{m+n}{2}} \Gamma\left(\frac{m+n}{2}\right)}{\left(\frac{x m}{n}+1\right)^{\frac{m+n}{2}}}=\frac{\left(\frac{m}{n}\right)^{\frac{m}{2}}}{B\left(\frac{m}{2}, \frac{n}{2}\right)} \cdot \frac{x^{\frac{m}{2}-1}}{\left(\frac{x m}{n}+1\right)^{\frac{m+n}{2}}}
\end{aligned}
$$

which with $x=F$ is the $F$-distribution with $m$ and $n$ degrees of freedom. Here we used the integral

$$
\int_{0}^{\infty} t^{z-1} e^{-\alpha t} d t=\frac{\Gamma(z)}{\alpha^{z}}
$$

in simplifying the expression.

### 16.7 Variance Ratio

A practical example where the $F$-distribution is applicable is when estimates of the variance for two independent samples from normal distributions

$$
s_{1}^{2}=\sum_{i=1}^{m} \frac{\left(x_{i}-\bar{x}\right)^{2}}{m-1} \quad \text { and } \quad s_{2}^{2}=\sum_{i=1}^{n} \frac{\left(y_{i}-\bar{y}\right)^{2}}{n-1}
$$

have been made. In this case $s_{1}^{2}$ and $s_{2}^{2}$ are so called normal theory estimates of $\sigma_{1}^{2}$ and $\sigma_{2}^{2}$ i.e. $(m-1) s_{1}^{2} / \sigma_{1}^{2}$ and $(n-1) s_{2}^{2} / \sigma_{2}^{2}$ are distributed according to the chi-square distribution with $m-1$ and $n-1$ degrees of freedom, respectively.

In this case the quantity

$$
F=\frac{s_{1}^{2}}{\sigma_{1}^{2}} \cdot \frac{\sigma_{2}^{2}}{s_{2}^{2}}
$$

is distributed according to the $F$-distribution with $m-1$ and $n-1$ degrees of freedom. If the true variances of the two populations are indeed the same then the variance ratio $s_{1}^{2} / s_{2}^{2}$ have the $F$-distribution. We may thus use this ratio to test the null hypothesis $H_{0}: \sigma_{1}^{2}=\sigma_{2}^{2}$ versus the alternative $H_{1}: \sigma_{1}^{2} \neq \sigma_{2}^{2}$ using the $F$-distribution. We would reject the null hypotheses at the $\alpha$ confidence level if the $F$-ratio is less than $F_{1-\alpha / 2, m-1, n-1}$ or greater than $F_{\alpha / 2, m-1, n-1}$ where $F_{\alpha, m, n}$ is defined by

$$
\int_{0}^{F_{\alpha, m, n}} f(F ; m, n) d F=1-\alpha
$$

i.e. $\alpha$ is the probability content of the distribution above the value $F_{\alpha, m-1, n-1}$. Note that the following relation between $F$-values corresponding to the same upper and lower confidence levels is valid

$$
F_{1-\alpha, m, n}=\frac{1}{F_{\alpha, n, m}}
$$

### 16.8 Analysis of Variance

As a simple example, which is often called analysis of variance, we regard $n$ observations of a dependent variable $x$ with overall mean $\bar{x}$ divided into $k$ classes on an independent variable. The mean in each class is denoted $\bar{x}_{j}$ for $j=1,2, \ldots, k$. In each of the $k$ classes there are $n_{j}$ observations together adding up to $n$, the total number of observations. Below we denote by $x_{j i}$ the $i$ :th observation in class $j$.

Rewrite the total sum of squares of the deviations from the mean

$$
\begin{aligned}
S S_{x} & =\sum_{j=1}^{k} \sum_{i=1}^{n_{j}}\left(x_{j i}-\bar{x}\right)^{2}=\sum_{j=1}^{k} \sum_{i=1}^{n_{j}}\left(\left(x_{j i}-\bar{x}_{j}\right)+\left(\bar{x}_{j}-\bar{x}\right)\right)^{2}= \\
& =\sum_{j=1}^{k} \sum_{i=1}^{n_{j}}\left[\left(x_{j i}-\bar{x}_{j}\right)^{2}+\left(\bar{x}_{j}-\bar{x}\right)^{2}+2\left(x_{j i}-\bar{x}_{j}\right)\left(\bar{x}_{j}-\bar{x}\right)\right]= \\
& =\sum_{j=1}^{k} \sum_{i=1}^{n_{j}}\left(x_{j i}-\bar{x}_{j}\right)^{2}+\sum_{j=1}^{k} \sum_{i=1}^{n_{j}}\left(\bar{x}_{j}-\bar{x}\right)^{2}+2 \sum_{j=1}^{k}\left(\bar{x}_{j}-\bar{x}\right) \sum_{i=1}^{n_{j}}\left(x_{j i}-\bar{x}_{j}\right)= \\
& =\sum_{j=1}^{k} \sum_{i=1}^{n_{j}}\left(x_{j i}-\bar{x}_{j}\right)^{2}+\sum_{j=1}^{k} n_{j}\left(\bar{x}_{j}-\bar{x}\right)^{2}=S S_{\text {within }}+S S_{\text {between }}
\end{aligned}
$$

i.e. the total sum of squares is the sum of the sum of squares within classes and the sum of squares between classes. Expressed in terms of variances

$$
n V(x)=\sum_{j=1}^{k} n_{j} V_{j}(x)+\sum_{j=1}^{k} n_{j}\left(\bar{x}_{j}-\bar{x}\right)^{2}
$$

If the variable $x$ is independent on the classification then the variance within groups and the variance between groups are both estimates of the same true variance. The quantity

$$
F=\frac{S S_{\text {between }} /(k-1)}{S S_{\text {within }} /(n-k)}
$$

is then distributed according to the $F$-distribution with $k-1$ and $n-k$ degrees of freedom. This may then be used in order to test the hypothesis of no dependence. A too high $F$-value would be unlikely and thus we can choose a confidence level at which we would reject the hypothesis of no dependence of $x$ on the classification.

Sometimes one also defines $\eta^{2}=S S_{\text {between }} / S S_{x}$, the proportion of variance explained, as a measure of the strength of the effects of classes on the variable $x$.

### 16.9 Calculation of Probability Content

In order to set confidence levels for the $F$-distribution we need to evaluate the cumulative function i.e. the integral

$$
1-\alpha=\int_{0}^{F_{\alpha}} f(F ; m, n) d F
$$

where we have used the notation $F_{\alpha}$ instead of $F_{\alpha, m, n}$ for convenience.

$$
\begin{aligned}
1-\alpha & =\frac{m^{\frac{m}{2}} n^{\frac{n}{2}}}{B\left(\frac{m}{2}, \frac{n}{2}\right)} \int_{0}^{F_{\alpha}} \frac{F^{\frac{m}{2}-1}}{(m F+n)^{\frac{m+n}{2}}} d F=\frac{\left(\frac{m}{n}\right)^{\frac{m}{2}}}{B\left(\frac{m}{2}, \frac{n}{2}\right)} \int_{0}^{F_{\alpha}} \frac{F^{\frac{m}{2}-1}}{\left(\frac{m F}{n}+1\right)^{\frac{m+n}{2}}} d F= \\
& =\frac{\left(\frac{m}{n}\right)^{\frac{m}{2}}}{B\left(\frac{m}{2}, \frac{n}{2}\right)} \int_{0}^{\frac{m F_{\alpha}}{n}} \frac{\left(\frac{u n}{m}\right)^{\frac{m}{2}-1}}{(u+1)^{\frac{m+n}{2}}} \frac{n}{m} d u=\frac{1}{B\left(\frac{m}{2}, \frac{n}{2}\right)} \int_{0}^{\frac{m F_{\alpha}}{n}} \frac{u^{\frac{m}{2}-1}}{(1+u)^{\frac{m+n}{2}}} d u
\end{aligned}
$$

where we made the substitution $u=\frac{m F}{n}$. The last integral we recognize as the incomplete Beta function $B_{x}$ defined for $0 \leq x \leq 1$ as

$$
B_{x}(p, q)=\int_{0}^{x} t^{p-1}(1-t)^{q-1} d t=\int_{0}^{\frac{x}{1-x}} \frac{u^{p-1}}{(1+u)^{p+q}} d u
$$

where we made the substitution $u=\frac{t}{1-t}$ i.e. $t=\frac{u}{1+u}$. We thus obtain

$$
1-\alpha=\frac{B_{x}\left(\frac{m}{2}, \frac{n}{2}\right)}{B\left(\frac{m}{2}, \frac{n}{2}\right)}=I_{x}\left(\frac{m}{2}, \frac{n}{2}\right)
$$

with $\frac{x}{1-x}=\frac{m F_{\alpha}}{n}$ i.e. $x=\frac{m F_{\alpha}}{n+m F_{\alpha}}$. The variable $x$ thus has a Beta distribution. Note that also $I_{x}(a, b)$ is called the incomplete Beta function (for historical reasons discussed below but see also section 42.7).

### 16.9.1 The Incomplete Beta function

In order to evaluate the incomplete Beta function we may use the serial expansion

$$
B_{x}(p, q)=x^{p}\left[\frac{1}{p}+\frac{1-q}{p+1} x+\frac{(1-q)(2-q)}{2!(p+2)} x^{2}+\ldots+\frac{(1-q)(2-q) \cdots(n-q)}{n!(p+n)} x^{n}+\ldots\right]
$$

For integer values of $q$ corresponding to even values of $n$ the sum may be stopped at $n=q-1$ since all remaining terms will be identical to zero in this case.

We may express the sum with successive terms expressed recursively in the previous term

$$
B_{x}(p, q)=x^{p} \sum_{r=0}^{\infty} t_{r} \quad \text { with } \quad t_{r}=t_{r-1} \cdot \frac{x(r-q)(p+r-1)}{r(p+r)} \quad \text { starting with } \quad t_{0}=\frac{1}{p}
$$

The sum normally converges quite fast but beware that e.g. for $p=q=\frac{1}{2}(m=$ $n=1$ ) the convergence is very slow. Also some cases with $q$ very big but $p$ small seem pathological since in these cases big terms with alternate signs cancel each other causing roundoff problems. It seems preferable to keep $q<p$ to assure faster convergence. This may be done by using the relation

$$
B_{x}(p, q)=B_{1}(q, p)-B_{1-x}(q, p)
$$

which if inserted in the formula for $1-\alpha$ gives

$$
1-\alpha=\frac{B_{1}\left(\frac{n}{2}, \frac{m}{2}\right)-B_{1-x}\left(\frac{n}{2}, \frac{m}{2}\right)}{B\left(\frac{m}{2}, \frac{n}{2}\right)} \Rightarrow \alpha=\frac{B_{1-x}\left(\frac{n}{2}, \frac{m}{2}\right)}{B\left(\frac{m}{2}, \frac{n}{2}\right)}=I_{1-x}\left(\frac{n}{2}, \frac{m}{2}\right)
$$

since $B_{1}(p, q)=B(p, q)=B(q, p)$.
A numerically better way to evaluate the incomplete Beta function $I_{x}(a, b)$ is by the continued fraction formula [10]

$$
I_{x}(a, b)=\frac{x^{a}(1-x)^{b}}{a B(a, b)}\left[\frac{1}{1+} \frac{d_{1}}{1+} \frac{d_{2}}{1+} \cdots\right]
$$

Here

$$
d_{2 m+1}=-\frac{(a+m)(a+b+m) x}{(a+2 m)(a+2 m+1)} \quad \text { and } \quad d_{2 m}=\frac{m(b-m) x}{(a+2 m-1)(a+2 m)}
$$

and the formula converges rapidly for $x<(a+1) /(a+b+1)$. For other $x$-values the same formula may be used after applying the symmetry relation

$$
I_{x}(a, b)=1-I_{1-x}(b, a)
$$

### 16.9.2 Final Formulæ

Using the serial expression for $B_{x}$ given in the previous subsection the probability content of the F-distribution may be calculated. The numerical situation is, however, not ideal. For integer $a$ - or $b$-values ${ }^{3}$ the following relation to the binomial distribution valid for integer values of $a$ is useful

$$
1-I_{x}(a, b)=I_{1-x}(b, a)=\sum_{i=0}^{a-1}\binom{a+b-1}{i} x^{i}(1-x)^{a+b-1-i}
$$

Our final formulæ are taken from [26], using $x=\frac{n}{n+m F}$ (note that this is one minus our previous definition of $x$ ),

- Even $m$ :

$$
\begin{aligned}
1-\alpha= & x^{\frac{n}{2}} \cdot\left[1+\frac{n}{2}(1-x)+\frac{n(n+1)}{2 \cdot 4}(1-x)^{2}+\ldots\right. \\
& \left.\ldots+\frac{n(n+2) \ldots(m+n-4)}{2 \cdot 4 \ldots(m-2)}(1-x)^{\frac{m-2}{2}}\right]
\end{aligned}
$$

- Even $n$ :

$$
\begin{aligned}
1-\alpha= & 1-(1-x)^{\frac{m}{2}}\left[1+\frac{m}{2} x+\frac{m(m+2)}{2 \cdot 4} x^{2}+\ldots\right. \\
& \left.\ldots+\frac{m(m+2) \ldots(m+n-4)}{2 \cdot 4 \ldots(n-2)} x^{\frac{n-2}{2}}\right]
\end{aligned}
$$

[^2]- Odd $m$ and $n$ :

$$
\begin{aligned}
1-\alpha= & 1-A+\beta \quad \text { with } \\
A= & \frac{2}{\pi}\left[\theta+\sin \theta\left(\cos \theta+\frac{2}{3} \cos ^{3} \theta+\ldots\right.\right. \\
& \left.\left.\ldots+\frac{2 \cdot 4 \ldots(n-3)}{1 \cdot 3 \ldots(n-2)} \cos ^{n-2} \theta\right)\right] \quad \text { for } n>1 \quad \text { and } \\
\beta= & \frac{2}{\sqrt{\pi}} \cdot \frac{\Gamma\left(\frac{n+1}{2}\right)}{\Gamma\left(\frac{n}{2}\right)} \cdot \sin \theta \cdot \cos ^{n} \theta \cdot\left[1+\frac{n+1}{3} \sin ^{2} \theta+\ldots\right. \\
& \left.\ldots+\frac{(n+1)(n+3) \ldots(m+n-4)}{3 \cdot 5 \ldots(n-2)} \sin ^{m-3} \theta\right] \text { for } m>1 \quad \text { where } \\
\theta= & \arctan \sqrt{\frac{n F}{m}} \text { and } \frac{\Gamma\left(\frac{n+1}{2}\right)}{\Gamma\left(\frac{n}{2}\right)}=\frac{(n-1)!!}{(n-2)!!} \cdot \frac{1}{\sqrt{\pi}}
\end{aligned}
$$

If $n=1$ then $A=2 \theta / \pi$ and if $m=1$ then $\beta=0$.

- For large values of $m$ and $n$ we use an approximation using the standard normal distribution where

$$
z=\frac{F^{\frac{1}{3}}\left(1-\frac{2}{9 n}\right)-\left(1-\frac{2}{9 m}\right)}{\sqrt{\frac{2}{9 m}+F^{\frac{2}{3}} \cdot \frac{2}{9 n}}}
$$

is approximately distributed according to the standard normal distribution. Confidence levels are obtained by

$$
1-\alpha=\frac{1}{\sqrt{2 \pi}} \int_{z}^{\infty} e^{-\frac{x^{2}}{2}} d x
$$

In table 5 on page 175 we show some percentage points for the $F$-distribution. Here $n$ is the degrees of freedom of the greater mean square and $m$ the degrees of freedom for the lesser mean square. The values express the values of $F$ which would be exceeded by pure chance in $10 \%, 5 \%$ and $1 \%$ of the cases, respectively.

### 16.10 Random Number Generation

Following the definition the quantity

$$
F=\frac{y_{m} / m}{y_{n} / n}
$$

where $y_{n}$ and $y_{m}$ are two variables distributed according to the chi-square distribution with $n$ and $m \mathrm{x}$ degrees of freedom respectively follows the F-distribution. We may thus use this relation inserting random numbers from chi-square distributions (see section 8.7).

## 17 Gamma Distribution

### 17.1 Introduction

The Gamma distribution is given by

$$
f(x ; a, b)=a(a x)^{b-1} e^{-a x} / \Gamma(b)
$$

where the parameters $a$ and $b$ are positive real quantities as is the variable $x$. Note that the parameter $a$ is simply a scale factor.

For $b \leq 1$ the distribution is J-shaped and for $b>1$ it is unimodal with its maximum at $x=\frac{b-1}{a}$.

In the special case where $b$ is a positive integer this distribution is often referred to as the Erlangian distribution.

For $b=1$ we obtain the exponential distribution and with $a=\frac{1}{2}$ and $b=\frac{n}{2}$ with $n$ an integer we obtain the chi-squared distribution with $n$ degrees of freedom.

In figure 13 we show the Gamma distribution for $b$-values of 2 and 5 .


Figure 13: Examples of Gamma distributions

### 17.2 Derivation of the Gamma Distribution

For integer values of $b$, i.e. for Erlangian distributions, we may derive the Gamma distribution from the Poisson assumptions. For a Poisson process where events happen at a rate of $\lambda$ the number of events in a time interval $t$ is given by Poisson distribution

$$
P(r)=\frac{(\lambda t)^{r} e^{-\lambda t}}{r!}
$$

The probability that the $k$ :th event occur at time $t$ is then given by

$$
\sum_{r=0}^{k-1} P(r)=\sum_{r=0}^{k-1} \frac{(\lambda t)^{r} e^{-\lambda t}}{r!}
$$

i.e. the probability that there are at least $k$ events in the time $t$ is given by

$$
F(t)=\sum_{r=k}^{\infty} P(r)=1-\sum_{r=0}^{k-1} \frac{(\lambda t)^{r} e^{-\lambda t}}{r!}=\int_{0}^{\lambda t} \frac{z^{k-1} e^{-z}}{(k-1)!} d z=\int_{0}^{t} \frac{\lambda^{k} z^{k-1} e^{-\lambda z}}{(k-1)!} d z
$$

where the sum has been replaced by an integral (no proof given here) and the substitution $z=\lambda z$ made at the end. This is the cumulative Gamma distribution with $a=\lambda$ and $b=k$, i.e. the time distribution for the $k$ :th event follows a Gamma distribution. In particular we may note that the time distribution for the occurrence of the first event follows an exponential distribution.

The Erlangian distribution thus describes the time distribution for exponentially distributed events occurring in a series. For exponential processes in parallel the appropriate distribution is the hyperexponential distribution.

### 17.3 Moments

The distribution has expectation value, variance, third and fourth central moments given by

$$
E(x)=\frac{b}{a}, \quad V(x)=\frac{b}{a^{2}}, \quad \mu_{3}=\frac{2 b}{a^{3}}, \quad \text { and } \quad \mu_{4}=\frac{3 b(2+b)}{a^{4}}
$$

The coefficients of skewness and kurtosis is given by

$$
\gamma_{1}=\frac{2}{\sqrt{b}} \quad \text { and } \quad \gamma_{2}=\frac{6}{b}
$$

More generally algebraic moments are given by

$$
\begin{aligned}
\mu_{n}^{\prime} & =\int_{0}^{\infty} x^{n} f(x) d x=\frac{a^{b}}{\Gamma(b)} \int_{0}^{\infty} x^{n+b-1} e^{-a x} d x= \\
& =\frac{a^{b}}{\Gamma(b)} \int_{0}^{\infty}\left(\frac{y}{a}\right)^{n+b-1} e^{-y} \frac{d y}{a}=\frac{\Gamma(n+b)}{a^{n} \Gamma(b)}= \\
& =\frac{b(b+1) \cdots(b+n-1)}{a^{n}}
\end{aligned}
$$

where we have made the substitution $y=a x$ in simplifying the integral.

### 17.4 Characteristic Function

The characteristic function is

$$
\begin{aligned}
\phi(t) & =E\left(e^{\imath t x}\right)=\frac{a^{b}}{\Gamma(b)} \int_{0}^{\infty} x^{b-1} e^{-x(a-\imath t)} d x= \\
& =\frac{a^{b}}{\Gamma(b)} \cdot \frac{1}{(a-\imath t)^{b}} \int_{0}^{\infty} y^{b-1} e^{-y} d y=\left(1-\frac{\imath t}{a}\right)^{-b}
\end{aligned}
$$

where we made the transformation $y=x(a-\imath t)$ in evaluating the integral.

### 17.5 Probability Content

In order to calculate the probability content for a Gamma distribution we need the cumulative (or distribution) function

$$
\begin{aligned}
F(x) & =\int_{0}^{x} f(x) d x=\frac{a^{b}}{\Gamma(b)} \int_{0}^{x} u^{b-1} e^{-a u} d u= \\
& =\frac{a^{b}}{\Gamma(b)} \int_{0}^{a x}\left(\frac{v}{a}\right)^{b-1} e^{-v} \frac{d v}{a}=\frac{1}{\Gamma(b)} \int_{0}^{a x} v^{b-1} e^{-v} d v=\frac{\gamma(b, a x)}{\Gamma(b)}
\end{aligned}
$$

where $\gamma(b, a x)$ denotes the incomplete gamma function ${ }^{4}$.

### 17.6 Random Number Generation

### 17.6.1 Erlangian distribution

In the case of an Erlangian distribution ( $b$ a positive integer) we obtain a random number by adding $b$ independent random numbers from an exponential distribution i.e.

$$
x=-\ln \left(\xi_{1} \cdot \xi_{2} \cdot \ldots \cdot \xi_{b}\right) / a
$$

where all the $\xi_{i}$ are uniform random numbers in the interval from zero to one. Note that care must be taken if $b$ is large in which case the product of uniform random numbers may become zero due to machine precision. In such cases simply divide the product in pieces and add the logarithms afterwards.

### 17.6.2 General case

In a more general case we use the so called Johnk's algorithm
i Denote the integer part of $b$ with $i$ and the fractional part with $f$ and put $r=0$. Let $\xi$ denote uniform random numbers in the interval from zero to one.
ii If $i>0$ then put $r=-\ln \left(\xi_{1} \cdot \xi_{2} \cdot \ldots \cdot \xi_{i}\right)$.
iii If $f=0$ then go to vii.
iv Calculate $w_{1}=\xi_{i+1}^{1 / f}$ and $w_{2}=\xi_{i+2}^{1 /(1-f)}$.
$\mathbf{v}$ If $w_{1}+w_{2}>1$ then go back to iv.
vi Put $r=r-\ln \left(\xi_{i+3}\right) \cdot \frac{w_{1}}{w_{1}+w_{2}}$.
vii Quit with $r=r / a$.

[^3]
### 17.6.3 Asymptotic Approximation

For $b$ big, say $b>15$, we may use the Wilson-Hilferty approximation:
i Calculate $q=1+\frac{1}{9 b}+\frac{z}{3 \sqrt{b}}$ where $z$ is a random number from a standard normal distribution.
ii Calculate $r=b \cdot q^{3}$.
iii If $r<0$ then go back to $\mathbf{i}$.
iv Quit with $r=r / a$.

## 18 Generalized Gamma Distribution

### 18.1 Introduction

The Gamma distribution is often used to describe variables bounded on one side. An even more flexible version of this distribution is obtained by adding a third parameter giving the so called generalized Gamma distribution

$$
f(x ; a, b, c)=a c(a x)^{b c-1} e^{-(a x)^{c}} / \Gamma(b)
$$

where $a$ (a scale parameter) and $b$ are the same real positive parameters as is used for the Gamma distribution but a third parameter $c$ has been added ( $c=1$ for the ordinary Gamma distribution). This new parameter may in principle take any real value but normally we consider the case where $c>0$ or even $c \geq 1$. Put $|c|$ in the normalization for $f(x)$ if $c<0$.

According to Hegyi [33] this density function first appeared in 1925 when L. Amoroso used it in analyzing the distribution of economic income. Later it has been used to describe the sizes of grains produced in comminution and drop size distributions in sprays etc.

In figure 14 we show the generalized Gamma distribution for different values of $c$ for the case $a=1$ and $b=2$.


Figure 14: Examples of generalized Gamma distributions

### 18.2 Cumulative Function

The cumulative function is given by

$$
F(x)= \begin{cases}\gamma\left(b,(a x)^{c}\right) / \Gamma(b)=P\left(b,(a x)^{c}\right) & \text { if } c>0 \\ \Gamma\left(b,(a x)^{c}\right) / \Gamma(b)=1-P\left(b,(a x)^{c}\right) & \text { if } c<0\end{cases}
$$

where $P$ is the incomplete Gamma function.

### 18.3 Moments

Algebraic moments are given by

$$
\mu_{n}^{\prime}=\frac{1}{a^{n}} \cdot \frac{\Gamma\left(b+\frac{n}{c}\right)}{\Gamma(b)}
$$

For negative values of $c$ the moments are finite for ranks $n$ satisfying $n / c>-b$ (or even just avoiding the singularities $\left.\frac{1}{a}+\frac{n}{c} \neq 0,-1,-2 \ldots\right)$.

### 18.4 Relation to Other Distributions

The generalized Gamma distribution is a general form which for certain parameter combinations gives many other distributions as special cases. In the table below we indicate some such relations. For notations see the corresponding section.

| Distribution | a | b | c | Section |
| :--- | :---: | :---: | :---: | :---: |
| Generalized gamma | a | b | c | 18 |
| Gamma | a | b | 1 | 17 |
| Chi-squared | $\frac{1}{2}$ | $\frac{n}{2}$ | 1 | 8 |
| Exponential | $\frac{1}{\alpha}$ | 1 | 1 | 14 |
| Weibull | $\frac{1}{\sigma}$ | 1 | $\eta$ | 41 |
| Rayleigh | $\frac{1}{\alpha \sqrt{2}}$ | 1 | 2 | 37 |
| Maxwell | $\frac{1}{\alpha \sqrt{2}}$ | $\frac{3}{2}$ | 2 | 25 |
| Standard normal (folded) | $\frac{1}{\sqrt{2}}$ | $\frac{1}{2}$ | 2 | 34 |

In reference [33], where this distribution is used in the description of multiplicity distributions in high energy particle collisions, more examples on special cases as well as more details regarding the distribution are given.

## 19 Geometric Distribution

### 19.1 Introduction

The geometric distribution is given by

$$
p(r ; p)=p(1-p)^{r-1}
$$

where the integer variable $r \geq 1$ and the parameter $0<p<1$ (no need to include limits since this give trivial special cases). It expresses the probability of having to wait exactly $r$ trials before the first successful event if the probability of a success in a single trial is $p$ (probability of failure $q=1-p$ ). It is a special case of the negative binomial distribution (with $k=1$ ).

### 19.2 Moments

The expectation value, variance, third and fourth moment are given by

$$
E(r)=\frac{1}{p} \quad V(r)=\frac{1-p}{p^{2}} \quad \mu_{3}=\frac{(1-p)(2-p)}{p^{3}} \quad \mu_{4}=\frac{(1-p)\left(p^{2}-9 p+9\right)}{p^{4}}
$$

The coefficients of skewness and kurtosis is thus

$$
\gamma_{1}=\frac{2-p}{\sqrt{1-p}} \quad \text { and } \quad \gamma_{2}=\frac{p^{2}-6 p+6}{1-p}
$$

### 19.3 Probability Generating Function

The probability generating function is

$$
G(z)=E\left(z^{r}\right)=\sum_{r=1}^{\infty} z^{r} p(1-p)^{r-1}=\frac{p z}{1-q z}
$$

### 19.4 Random Number Generation

The cumulative distribution may be written

$$
P(k)=\sum_{r=1}^{k} p(r)=1-q^{k} \quad \text { with } \quad q=1-p
$$

which can be used in order to obtain a random number from a geometric distribution by generating uniform random numbers between zero and one until such a number (the $k:$ th) is above $q^{k}$.

A more straightforward technique is to generate uniform random numbers $\xi_{i}$ until we find a success where $\xi_{k} \leq p$.

These two methods are both very inefficient for low values of $p$. However, the first technique may be solved explicitly

$$
\sum_{r=1}^{k} P(r)=\xi \quad \Rightarrow \quad k=\frac{\ln \xi}{\ln q}
$$

which implies taking the largest integer less than $k+1$ as a random number from a geometric distribution. This method is quite independent of the value of $p$ and we found [14] that a reasonable breakpoint below which to use this technique is $p=0.07$ and use the first method mentioned above this limit. With such a method we do not gain by creating a cumulative vector for the random number generation as we do for many other discrete distributions.

## 20 Hyperexponential Distribution

### 20.1 Introduction

The hyperexponential distribution describes exponential processes in parallel and is given by

$$
f\left(x ; p, \lambda_{1}, \lambda_{2}\right)=p \lambda_{1} e^{-\lambda_{1} x}+q \lambda_{2} e^{-\lambda_{2} x}
$$

where the variable $x$ and the parameters $\lambda_{1}$ and $\lambda_{2}$ are positive real quantities and $0 \leq p \leq 1$ is the proportion for the first process and $q=1-p$ the proportion of the second.

The distribution describes the time between events in a process where the events are generated from two independent exponential distributions. For exponential processes in series we obtain the Erlangian distribution (a special case of the Gamma distribution).

The hyperexponential distribution is easily generalized to the case with $k$ exponential processes in parallel

$$
f(x)=\sum_{i=1}^{k} p_{i} \lambda_{i} e^{-\lambda_{i} x}
$$

where $\lambda_{i}$ is the slope and $p_{i}$ the proportion for each process (with the constraint that $\sum p_{i}=1$ ).

The cumulative (distribution) function is

$$
F(x)=p\left(1-e^{-\lambda_{1} x}\right)+q\left(1-e^{-\lambda_{2} x}\right)
$$

and it is thus straightforward to calculate the probability content in any given situation.

### 20.2 Moments

Algebraic moments are given by

$$
\mu_{n}^{\prime}=n!\left(\frac{p}{\lambda_{1}^{n}}+\frac{q}{\lambda_{2}^{n}}\right)
$$

Central moments becomes somewhat complicated but the second central moment, the variance of the distribution, is given by

$$
\mu_{2}=V(x)=\frac{p}{\lambda_{1}^{2}}+\frac{q}{\lambda_{2}^{2}}+p q\left(\frac{1}{\lambda_{1}}-\frac{1}{\lambda_{2}}\right)^{2}
$$

### 20.3 Characteristic Function

The characteristic function of the hyperexponential distribution is given by

$$
\phi(t)=\frac{p}{1-\frac{\imath t}{\lambda_{1}}}+\frac{q}{1-\frac{\imath t}{\lambda_{2}}}
$$

### 20.4 Random Number Generation

Generating two uniform random numbers between zero and one, $\xi_{1}$ and $\xi_{2}$, we obtain a random number from a hyperexponential distribution by

- If $\xi_{1} \leq p$ then put $x=-\frac{\ln \xi_{2}}{\lambda_{1}}$.
- If $\xi_{1}>p$ then put $x=-\frac{\ln \xi_{2}}{\lambda_{2}}$.
i.e. using $\xi_{1}$ we choose which of the two processes to use and with $\xi_{2}$ we generate an exponential random number for this process. The same technique is easily generalized to the case with $k$ processes.


## 21 Hypergeometric Distribution

### 21.1 Introduction

The Hypergeometric distribution is given by

$$
p(r ; n, N, M)=\frac{\binom{M}{r}\binom{N-M}{n-r}}{\binom{N}{n}}
$$

where the discrete variable $r$ has limits from $\max (0, n-N+M)$ to $\min (n, M)$ (inclusive). The parameters $n(1 \leq n \leq N), N(N \geq 1)$ and $M(M \geq 1)$ are all integers.

This distribution describes the experiment where elements are picked at random without replacement. More precisely, suppose that we have $N$ elements out of which $M$ has a certain attribute (and $N-M$ has not). If we pick $n$ elements at random without replacement $p(r)$ is the probability that exactly $r$ of the selected elements come from the group with the attribute.

If $N \gg n$ this distribution approaches a binomial distribution with $p=\frac{M}{N}$.
If instead of two groups there are $k$ groups with different attributes the generalized hypergeometric distribution

$$
p(\underline{r} ; n, N, \underline{M})=\frac{\prod_{i=1}^{k}\binom{M_{i}}{r_{i}}}{\binom{N}{n}}
$$

where, as before, $N$ is the total number of elements, $n$ the number of elements picked and $\underline{M}$ a vector with the number of elements of each attribute (whose sum should equal $N$ ). Here $n=\sum r_{i}$ and the limits for each $r_{k}$ is given by $\max \left(0, n-N+M_{k}\right) \leq r_{k} \leq \min \left(n, M_{k}\right)$.

### 21.2 Probability Generating Function

The Hypergeometric distribution is closely related to the hypergeometric function, see appendix B on page 167 , and the probability generating function is given by

$$
G(z)=\frac{\binom{N-M}{n}}{\binom{N}{n}}{ }_{2} F_{1}(-n,-M ; N-M-n+1 ; z)
$$

### 21.3 Moments

With the notation $p=\frac{M}{N}$ and $q=1-p$, i.e. the proportions of elements with and without the attribute, the expectation value, variance, third and fourth central moments are given by

$$
\begin{aligned}
E(r) & =n p \\
V(r) & =n p q \frac{N-n}{N-1} \\
\mu_{3} & =n p q(q-p) \frac{(N-n)(N-2 n)}{(N-1)(N-2)} \\
\mu_{4} & =n p q(N-n) \frac{N(N+1)-6 n(N-n)+3 p q\left(N^{2}(n-2)-N n^{2}+6 n(N-n)\right)}{(N-1)(N-2)(N-3)}
\end{aligned}
$$

For the generalized hypergeometric distribution using $p_{i}=M_{i} / N$ and $q_{i}=1-p_{i}$ we find moments of $r_{i}$ using the formulæ above regarding the group $i$ as having an attribute and all other groups as not having the attribute. the covariances are given by

$$
\operatorname{Cov}\left(r_{i}, r_{j}\right)=n p_{i} p_{j} \frac{N-n}{N-1}
$$

### 21.4 Random Number Generation

To generate random numbers from a hypergeometric distribution one may construct a routine which follow the recipe above by picking elements at random. The same technique may be applied for the generalized hypergeometric distribution. Such techniques may be sufficient for many purposes but become quite slow.

For the hypergeometric distribution a better choice is to construct the cumulative function by adding up the individual probabilities using the recursive formula

$$
p(r)=\frac{(M-r+1)(n-r+1)}{r(N-M-n+r)} p(r-1)
$$

for the appropriate $r$-range (see above) starting with $p\left(r_{\text {min }}\right)$. With the cumulative vector and one single uniform random number one may easily make a fast algorithm in order to obtain the required random number.

## 22 Logarithmic Distribution

### 22.1 Introduction

The logarithmic distribution is given by

$$
p(r ; p)=-\frac{(1-p)^{r}}{r \ln p}
$$

where the variable $r \geq 1$ is an integer and the parameter $0<p<1$ is a real quantity.
It is a limiting form of the negative binomial distribution when the zero class has been omitted and the parameter $k \rightarrow 0$ (see section 29.4.3).

### 22.2 Moments

The expectation value and variance are given by

$$
E(r)=-\frac{\alpha q}{p} \quad \text { and } \quad V(r)=-\frac{\alpha q(1+\alpha q)}{p^{2}}
$$

where we have introduced $q=1-p$ and $\alpha=1 / \ln p$ for convenience. The third and fourth central moments are given by

$$
\begin{aligned}
& \mu_{3}=-\frac{\alpha q}{p^{3}}\left(1+q+3 \alpha q+2 \alpha^{2} q^{2}\right) \\
& \mu_{4}=-\frac{\alpha q}{p^{4}}\left(1+4 q+q^{2}+4 \alpha q(1+q)+6 \alpha^{2} q^{2}+3 \alpha^{3} q^{3}\right)
\end{aligned}
$$

More generally factorial moments are easily found using the probability generating function

$$
E(r(r-1) \cdots(r-k+1))=\left.\frac{d^{k}}{d z^{k}} G(z)\right|_{z=1}=-(n-1)!\alpha \frac{q^{k}}{p^{k}}
$$

From these moments ordinary algebraic and central moments may be found by straightforward but somewhat tedious algebra.

### 22.3 Probability Generating Function

The probability generating function is given by

$$
G(z)=E\left(z^{r}\right)=\sum_{r=0}^{\infty}-\frac{z^{r}(1-p)^{r}}{r \ln p}=-\frac{1}{\ln p} \sum_{r=0}^{\infty} \frac{(z q)^{r}}{r}=\frac{\ln (1-z q)}{\ln (1-q)}
$$

where $q=1-p$ and since

$$
\ln (1-x)=-\left(x+\frac{x^{2}}{2}+\frac{x^{3}}{3}+\frac{x^{4}}{4}+\ldots\right) \quad \text { for } \quad-1 \leq x<1
$$

### 22.4 Random Number Generation

The most straightforward way to obtain random numbers from a logarithmic distribution is to use the cumulative technique. If $p$ is fixed the most efficient way is to prepare a cumulative vector starting with $p(1)=-\alpha q$ and subsequent elements by the recursive formula $p(i)=p(i-1) q / i$. The cumulative vector may, however, become very long for small values of $p$. Ideally it should extend until the cumulative vector element is exactly one due to computer precision. It $p$ is not fixed the same procedure has to be made at each generation.

## 23 Logistic Distribution

### 23.1 Introduction

The Logistic distribution is given by

$$
f(x ; a, k)=\frac{e^{z}}{k\left(1+e^{z}\right)^{2}} \quad \text { with } \quad z=\frac{x-a}{k}
$$

where the variable $x$ is a real quantity, the parameter $a$ a real location parameter (the mode, median, and mean) and $k$ a positive real scale parameter (related to the standard deviation). In figure 15 the logistic distribution with parameters $a=0$ and $k=1$ (i.e. $z=x$ ) is shown.


Figure 15: Graph of logistic distribution for $a=0$ and $k=1$

### 23.2 Cumulative Distribution

The distribution function is given by

$$
F(x)=1-\frac{1}{1+e^{z}}=\frac{1}{1+e^{-z}}=\frac{1}{1+e^{-\frac{x-a}{k}}}
$$

The inverse function is found by solving $F(x)=\alpha$ giving

$$
x=F^{-1}(\alpha)=a-k \ln \left(\frac{1-\alpha}{\alpha}\right)
$$

from which we may find e.g. the median as $\mathcal{M}=a$. Similarly the lower and upper quartiles are given by $\mathcal{Q}_{1,2}=a \mp k \ln 3$.

### 23.3 Characteristic Function

The characteristic function is given by

$$
\begin{aligned}
\phi(t) & =E\left(e^{\imath t x}\right)=\int_{-\infty}^{\infty} e^{\imath t x} \frac{e^{\frac{x-a}{k}}}{k\left(1+e^{\frac{x-a}{k}}\right)^{2}} d x=e^{\imath t a} \int_{-\infty}^{\infty} \frac{e^{\imath t z k} e^{z}}{k\left(1+e^{z}\right)^{2}} k d z= \\
& =e^{\imath t a} \int_{0}^{\infty} \frac{y^{\imath t k} y}{(1+y)^{2}} \cdot \frac{d y}{y}=e^{\imath t a} B(1+\imath t k, 1-\imath t k)= \\
& =e^{\imath t a} \frac{\Gamma(1+\imath t k) \Gamma(1-\imath t k)}{\Gamma(2)}=e^{\imath t a} \imath t k \Gamma(\imath t k) \Gamma(1-\imath t k)=e^{\imath t a} \frac{\imath t k \pi}{\sin \pi \imath t k}
\end{aligned}
$$

where we have used the transformations $z=(x-a) / k$ and $y=e^{z}$ in simplifying the integral, at the end identifying the beta function, and using relation of this in terms of Gamma functions and their properties (see appendix A in section 42).

### 23.4 Moments

The characteristic function is slightly awkward to use in determining the algebraic moments by taking partial derivatives in $t$. However, using

$$
\ln \phi(t)=\imath t a+\ln \Gamma(1+\imath t k)+\ln \Gamma(1-\imath t k)
$$

we may determine the cumulants of the distributions. In the process we take derivatives of $\ln \phi(t)$ which involves polygamma functions (see section 42.4) but all of them with argument 1 when inserting $t=0$ a case which may be explicitly written in terms of Riemann's zetafunctions with even real argument (see page 59). It is quite easily found that all cumulants of odd order except $\kappa_{1}=a$ vanish and that for even orders

$$
\kappa_{2 n}=2 k^{2 n} \psi^{(2 n-1)}(1)=2(2 n-1)!k^{2 n} \zeta(2 n)=2(2 n-1)!k^{2 n} \frac{2^{2 n-1} \pi^{2 n}\left|B_{2 n}\right|}{(2 n)!}
$$

for $n=1,2, \ldots$ and where $B_{2 n}$ are the Bernoulli numbers (see table 4 on page 174).
Using this formula lower order moments and the coefficients of skewness and kurtosis is found to be

$$
\begin{aligned}
\mu_{1}^{\prime} & =E(x)=\kappa_{1}=a \\
\mu_{2} & =V(x)=\kappa_{2}=k^{2} \pi^{2} / 3 \\
\mu_{3} & =0 \\
\mu_{4} & =\kappa_{4}+3 \kappa_{2}^{2}=\frac{2 k^{4} \pi^{4}}{15}+\frac{k^{4} \pi^{4}}{3}=\frac{7 k^{4} \pi^{4}}{15} \\
\mu_{5} & =0 \\
\mu_{6} & =\kappa_{6}+15 \kappa_{4} \kappa_{2}+10 \kappa_{3}^{2}+15 \kappa_{2}^{3}= \\
& =\frac{16 k^{6} \pi^{6}}{63}+\frac{2 k^{6} \pi^{6}}{3}+\frac{15 k^{6} \pi^{6}}{27}=\frac{31 k^{6} \pi^{6}}{11} \\
\gamma_{1} & =0 \\
\gamma_{2} & =1.2 \quad \text { (exact) }
\end{aligned}
$$

### 23.5 Random numbers

Using the inverse cumulative function one easily obtains a random number from a logistic distribution by

$$
x=a+k \ln \left(\frac{\xi}{1-\xi}\right)
$$

with $\xi$ a uniform random number between zero and one (limits not included).

## 24 Log-normal Distribution

### 24.1 Introduction

The log-normal distribution or is given by

$$
f(x ; \mu, \sigma)=\frac{1}{x \sigma \sqrt{2 \pi}} e^{-\frac{1}{2}\left(\frac{\ln x-\mu}{\sigma}\right)^{2}}
$$

where the variable $x>0$ and the parameters $\mu$ and $\sigma>0$ all are real numbers. It is sometimes denoted $\Lambda\left(\mu, \sigma^{2}\right)$ in the same spirit as we often denote a normally distributed variable by $N\left(\mu, \sigma^{2}\right)$.

If $u$ is distributed as $N\left(\mu, \sigma^{2}\right)$ and $u=\ln x$ then $x$ is distributed according to the log-normal distribution.

Note also that if $x$ has the distribution $\Lambda\left(\mu, \sigma^{2}\right)$ then $y=e^{a} x^{b}$ is distributed as $\Lambda(a+$ $b \mu, b^{2} \sigma^{2}$ ).

In figure 16 we show the log-normal distribution for the basic form, with $\mu=0$ and $\sigma=1$.


Figure 16: Log-normal distribution

The log-normal distribution is sometimes used as a first approximation to the Landau distribution describing the energy loss by ionization of a heavy charged particle ( $c f$ also the Moyal distribution in section 26).

### 24.2 Moments

The expectation value and the variance of the distribution are given by

$$
E(x)=e^{\mu+\frac{\sigma^{2}}{2}} \quad \text { and } \quad V(x)=e^{2 \mu+\sigma^{2}}\left(e^{\sigma^{2}}-1\right)
$$

and the coefficients of skewness and kurtosis becomes

$$
\gamma_{1}=\sqrt{e^{\sigma^{2}}-1}\left(e^{\sigma^{2}}+2\right) \quad \text { and } \quad \gamma_{2}=\left(e^{\sigma^{2}}-1\right)\left(e^{3 \sigma^{2}}+3 e^{2 \sigma^{2}}+6 e^{\sigma^{2}}+6\right)
$$

More generally algebraic moments of the log-normal distribution are given by

$$
\begin{aligned}
\mu_{k}^{\prime}=E\left(x^{k}\right) & =\frac{1}{\sigma \sqrt{2 \pi}} \int_{0}^{\infty} x^{k-1} e^{-\frac{1}{2}\left(\frac{\ln x-\mu}{\sigma}\right)^{2}} d x=\frac{1}{\sigma \sqrt{2 \pi}} \int_{-\infty}^{\infty} e^{y k} e^{-\frac{1}{2}\left(\frac{y-\mu}{\sigma}\right)^{2}} d y= \\
& =\frac{1}{\sigma \sqrt{2 \pi}} e^{k \mu+\frac{k^{2} \sigma^{2}}{2}} \int_{-\infty}^{\infty} e^{-\frac{1}{2}\left(\frac{y-\mu-k \sigma^{2}}{\sigma}\right)^{2}} d y=e^{k \mu+\frac{k^{2} \sigma^{2}}{2}}
\end{aligned}
$$

where we have used the transformation $y=\ln x$ in simplifying the integral.

### 24.3 Cumulative Distribution

The cumulative distribution, or distribution function, for the log-normal distribution is given by

$$
\begin{aligned}
F(x) & =\frac{1}{\sigma \sqrt{2 \pi}} \int_{0}^{x} \frac{1}{t} e^{-\frac{1}{2}\left(\frac{\ln t-\mu}{\sigma}\right)^{2}} d t=\frac{1}{\sigma \sqrt{2 \pi}} \int_{0}^{\ln x} e^{-\frac{1}{2}\left(\frac{y-\mu}{\sigma}\right)^{2}} d y= \\
& =\frac{1}{2} \pm \frac{1}{2} P\left(\frac{1}{2}, \frac{z^{2}}{2}\right)
\end{aligned}
$$

where we have put $z=(\ln x-\mu) / \sigma$ and the positive sign is valid for $z \geq 0$ and the negative $\operatorname{sign}$ for $z<0$.

### 24.4 Random Number Generation

The most straightforward way of achieving random numbers from a log-normal distribution is to generate a random number $u$ from a normal distribution with mean $\mu$ and standard deviation $\sigma$ and construct $r=e^{u}$.

## 25 Maxwell Distribution

### 25.1 Introduction

The Maxwell distribution is given by

$$
f(x ; \alpha)=\frac{1}{\alpha^{3}} \sqrt{\frac{2}{\pi}} x^{2} e^{-\frac{x^{2}}{2 \alpha^{2}}}
$$

where the variable $x$ with $x \geq 0$ and the parameter $\alpha$ with $\alpha>0$ are real quantities. It is named after the famous scottish physicist James Clerk Maxwell (1831-1879).

The parameter $\alpha$ is simply a scale factor and the variable $y=x / \alpha$ has the simplified distribution

$$
g(y)=\sqrt{\frac{2}{\pi}} y^{2} e^{-\frac{y^{2}}{2}}
$$



Figure 17: The Maxwell distribution
The distribution, shown in figure 17, has a mode at $x=\alpha$ and is positively skewed.

### 25.2 Moments

Algebraic moments are given by

$$
E\left(x^{n}\right)=\int_{0}^{\infty} x^{n} f(x) d x=\frac{1}{2 \alpha^{3}} \sqrt{\frac{2}{\pi}} \int_{-\infty}^{\infty}|x|^{n+2} e^{-x^{2} / 2 \alpha^{2}}
$$

i.e. we have a connection to the absolute moments of the Gauss distribution. Using these (see section on the normal distribution) the result is

$$
E\left(x^{n}\right)= \begin{cases}\sqrt{\frac{2}{\pi}} 2^{k} k!\alpha^{2 k-1} & \text { for } n=2 k-1 \\ (n+1)!!\alpha^{n} & \text { for } n \text { even }\end{cases}
$$

Specifically we note that the expectation value, variance, and the third and fourth central moments are given by

$$
E(x)=2 \alpha \sqrt{\frac{2}{\pi}}, \quad V(x)=\alpha^{2}\left(3-\frac{8}{\pi}\right), \quad \mu_{3}=2 \alpha^{3}\left(\frac{16}{\pi}-5\right) \sqrt{\frac{2}{\pi}}, \quad \text { and } \quad \mu_{4}=\alpha^{4}\left(15-\frac{8}{\pi}\right)
$$

The coefficients of skewness and kurtosis is thus

$$
\gamma_{1}=\frac{2\left(\frac{16}{\pi}-5\right) \sqrt{\frac{2}{\pi}}}{\left(3-\frac{8}{\pi}\right)^{\frac{3}{2}}} \approx 0.48569 \quad \text { and } \quad \gamma_{2}=\frac{15-\frac{8}{\pi}}{\left(3-\frac{8}{\pi}\right)^{2}}-3 \approx 0.10818
$$

### 25.3 Cumulative Distribution

The cumulative distribution, or the distribution function, is given by

$$
F(x)=\int_{0}^{x} f(y) d y=\frac{1}{a^{3}} \sqrt{\frac{2}{\pi}} \int_{0}^{x} y^{2} e^{-\frac{y^{2}}{2 \alpha^{2}}} d y=\frac{2}{\sqrt{\pi}} \int_{0}^{\frac{x^{2}}{2 \alpha^{2}}} \sqrt{z} e^{-z} d z=\frac{\gamma\left(\frac{3}{2}, \frac{x^{2}}{2 \alpha^{2}}\right)}{\Gamma\left(\frac{3}{2}\right)}=P\left(\frac{3}{2}, \frac{x^{2}}{2 \alpha^{2}}\right)
$$

where we have made the substitution $z=\frac{y^{2}}{2 \alpha^{2}}$ in order to simplify the integration. Here $P(a, x)$ is the incomplete Gamma function.

Using the above relation we may estimate the median $\mathcal{M}$ and the lower and upper quartile, $\mathcal{Q}_{1}$ and $\mathcal{Q}_{3}$, as

$$
\begin{aligned}
& \mathcal{Q}_{1}=\alpha \sqrt{P^{-1}\left(\frac{3}{2}, \frac{1}{2}\right)} \approx 1.10115 \alpha \\
& \mathcal{M}=\alpha \sqrt{P^{-1}\left(\frac{3}{2}, \frac{1}{2}\right)} \approx 1.53817 \alpha \\
& \mathcal{Q}_{3}=\alpha \sqrt{P^{-1}\left(\frac{3}{2}, \frac{1}{2}\right)} \approx 2.02691 \alpha
\end{aligned}
$$

where $P^{-1}(a, p)$ denotes the inverse of the incomplete Gamma function i.e. the value $x$ for which $P(a, x)=p$.

### 25.4 Kinetic Theory

The following is taken from kinetic theory, see e.g. [34]. Let $v=\left(v_{x}, v_{y}, v_{z}\right)$ be the velocity vector of a particle where each component is distributed independently according to normal distributions with zero mean and the same variance $\sigma^{2}$.

First construct

$$
w=\frac{v^{2}}{\sigma^{2}}=\frac{v_{x}^{2}}{\sigma^{2}}+\frac{v_{y}^{2}}{\sigma^{2}}+\frac{v_{z}^{2}}{\sigma^{2}}
$$

Since $v_{x} / \sigma, v_{y} / \sigma$, and $v_{z} / \sigma$ are distributed as standard normal variables the sum of their squares has the chi-squared distribution with 3 degrees of freedom i.e. $g(w)=\sqrt{\frac{w}{2 \pi}} e^{-w / 2}$ which leads to

$$
f(v)=g(w)\left|\frac{d w}{d v}\right|=g\left(\frac{v^{2}}{\sigma^{2}}\right) \frac{2 v}{\sigma^{2}}=\frac{1}{\sigma^{3}} \sqrt{\frac{2}{\pi}} v^{2} e^{-\frac{v^{2}}{2 \sigma^{2}}}
$$

which we recognize as a Maxwell distribution with $\alpha=\sigma$.

In kinetic theory $\sigma=k T / m$, where $k$ is Boltzmann's constant, $T$ the temperature, and $m$ the mass of the particles, and we thus have

$$
f(v)=\sqrt{\frac{2 m^{3}}{\pi k^{3} T^{3}}} v^{2} e^{-\frac{m v^{2}}{2 k T}}
$$

The distribution in kinetic energy $E=m v^{2} / 2$ becomes

$$
g(E)=\sqrt{\frac{4 E}{\pi k^{3} T^{3}}} e^{-\frac{E}{k T}}
$$

which is a Gamma distribution with parameters $a=1 / k T$ and $b=\frac{3}{2}$.

### 25.5 Random Number Generation

To obtain random numbers from the Maxwell distribution we first make the transformation $y=x^{2} / 2 \alpha^{2}$ a variable which follow the Gamma distribution $g(y)=\sqrt{y} e^{-y} / \Gamma\left(\frac{3}{2}\right)$.

A random number from this distribution may be obtained using the so called Johnk's algorithm which in this particular case becomes (denoting independent pseudorandom numbers from a uniform distribution from zero to one by $\xi_{i}$ )
i Put $r=-\ln \xi_{1}$ i.e. a random number from an exponential distribution.
ii Calculate $w_{1}=\xi_{2}^{2}$ and $w_{2}=\xi_{3}^{2}$ (with new uniform random numbers $\xi_{2}$ and $\xi_{3}$ each iteration, of course).
iii If $w=w_{1}+w_{2}>1$ then go back to ii above.
iv Put $r=r-\frac{w_{1}}{w} \ln \xi_{4}$
v Finally construct $a \sqrt{2 r}$ as a random number from the Maxwell distribution with parameter $r$.

Following the examples given above we may also use three independent random numbers from a standard normal distribution, $z_{1}, z_{2}$, and $z_{3}$, and construct

$$
r=\frac{1}{\alpha} \sqrt{z_{1}^{2}+z_{2}^{2}+z_{3}^{2}}
$$

However, this technique is not as efficient as the one outlined above.
As a third alternative we could also use the cumulative distribution putting

$$
F(x)=\xi \quad \Rightarrow \quad P\left(\frac{3}{2}, \frac{x^{2}}{2 \alpha^{2}}\right)=\xi \quad \Rightarrow \quad x=\alpha \sqrt{2 P^{-1}\left(\frac{3}{2}, \xi\right)}
$$

where $P^{-1}(a, p)$, as above, denotes the value $x$ where $P(a, x)=p$. This technique is, however, much slower than the alternatives given above.

The first technique described above is not very fast but still the best alternative presented here. Also it is less dependent on numerical algorithms (such as those to find the inverse of the incomplete Gamma function) which may affect the precision of the method.

## 26 Moyal Distribution

### 26.1 Introduction

The Moyal distribution is given by

$$
f(z)=\frac{1}{\sqrt{2 \pi}} \exp \left\{-\frac{1}{2}\left(z+e^{-z}\right)\right\}
$$

for real values of $z$. A scale shift and a scale factor is introduced by making the standardized variable $z=(x-\mu) / \sigma$ and hence the distribution in the variable $x$ is given by

$$
g(x)=\frac{1}{\sigma} f\left(\frac{x-\mu}{\sigma}\right)
$$

Without loss of generality we treat the Moyal distribution in its simpler form, $f(z)$, in this document. Properties for $g(x)$ are easily obtained from these results which is sometimes indicated.

The Moyal distribution is a universal form for
(a) the energy loss by ionization for a fast charged particle and
(b) the number of ion pairs produced in this process.

It was proposed by J. E. Moyal [35] as a good approximation to the Landau distribution. It was also shown that it remains valid taking into account quantum resonance effects and details of atomic structure of the absorber.


Figure 18: The Moyal distribution

The distribution, shown in figure 18 , has a mode at $z=0$ and is positively skewed. This implies that the mode of the $x$-distribution, $g(x)$, is equal to the parameter $\mu$.

### 26.2 Normalization

Making the transformation $x=e^{-z}$ we find that

$$
\begin{aligned}
\int_{-\infty}^{\infty} f(z) d z & =\int_{0}^{\infty} \frac{1}{\sqrt{2 \pi}} \exp \left\{-\frac{1}{2}(-\ln x+x)\right\} \frac{d x}{x}=\frac{1}{\sqrt{2 \pi}} \int_{0}^{\infty} \frac{e^{-\frac{x}{2}}}{\sqrt{x}} d x= \\
& =\frac{1}{\sqrt{2 \pi}} \int_{0}^{\infty} \frac{e^{-y}}{\sqrt{2 y}} 2 d y=\frac{1}{\sqrt{\pi}} \int_{0}^{\infty} \frac{e^{-y}}{\sqrt{y}} d y=\frac{1}{\sqrt{\pi}} \Gamma\left(\frac{1}{2}\right)=1
\end{aligned}
$$

where we have made the simple substitution $y=x / 2$ in order to clearly recognize the Gamma function at the end. The distribution is thus properly normalized.

### 26.3 Characteristic Function

The characteristic function for the Moyal distribution becomes

$$
\begin{aligned}
\phi(t) & =E\left(e^{\imath t z}\right)=\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} e^{\imath t z} e^{-\frac{1}{2}\left(z+e^{-z}\right)} d z=\frac{1}{\sqrt{2 \pi}} \int_{0}^{\infty}(2 x)^{\frac{1}{2}(1-2 t t)} e^{-x} \frac{d x}{x}= \\
& =\frac{2^{\frac{1}{2}(1-2 t t)}}{\sqrt{2 \pi}} \int_{0}^{\infty} x^{-\frac{1}{2}(1+2 \imath t)} e^{-x} d x=\frac{2^{-\imath t}}{\sqrt{\pi}} \Gamma\left(\frac{1}{2}-\imath t\right)
\end{aligned}
$$

where we made the substitution $x=e^{-z} / 2$ in simplifying the integral. The last relation to the Gamma function with complex argument is valid when the real part of the argument is positive which indeed is true in the case at hand.

### 26.4 Moments

As in some other cases the most convenient way to find the moments of the distribution is via its cumulants (see section 2.5). We find that

$$
\begin{aligned}
& \kappa_{1}=-\ln 2-\psi\left(\frac{1}{2}\right)=\ln 2+\gamma \\
& \kappa_{n}=(-1)^{n} \psi^{(n-1)}\left(\frac{1}{2}\right)=(n-1)!\left(2^{n}-1\right) \zeta_{n} \quad \text { for } \quad n \geq 2
\end{aligned}
$$

with $\gamma \approx 0.5772156649$ Euler's constant, $\psi^{(n)}$ polygamma functions (see section 42.4) and $\zeta$ Riemann's zeta-function (see page 59). Using the cumulants we find the lower order moments and the coefficients of skewness and kurtosis to be

$$
\begin{aligned}
& \mu_{1}^{\prime}=E(z)=\kappa_{1}=\ln 2+\gamma \approx 1.27036 \\
& \mu_{2}=V(z)=\kappa_{2}=\psi^{(1)}\left(\frac{1}{2}\right)=\frac{\pi^{2}}{2} \approx 4.93480 \\
& \mu_{3}=\kappa_{3}=-\psi^{(2)}\left(\frac{1}{2}\right)=14 \zeta_{3} \\
& \mu_{4}=\kappa_{4}+3 \kappa_{2}^{2}=\psi^{(3)}\left(\frac{1}{2}\right)+3 \psi^{(1)}\left(\frac{1}{2}\right)^{2}=\frac{7 \pi^{4}}{4} \\
& \gamma_{1}=\frac{28 \sqrt{2} \zeta_{3}}{\pi^{3}} \approx 1.53514 \\
& \gamma_{2}=4
\end{aligned}
$$

For the distribution $g(x)$ we have $E(x)=\sigma E(z)+\mu, V(x)=\sigma^{2} V(z)$ or more generally central moments are obtained by $\mu_{n}(x)=\sigma^{n} \mu_{n}(z)$ for $n \geq 2$ while $\gamma_{1}$ and $\gamma_{2}$ are identical.

### 26.5 Cumulative Distribution

Using the same transformations as was used above in evaluating the normalization of the distribution we write the cumulative (or distribution) function as

$$
\begin{aligned}
F(Z) & =\int_{-\infty}^{Z} f(z) d z=\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{Z} \exp \left\{-\frac{1}{2}\left(z+e^{-z}\right)\right\}=\frac{1}{\sqrt{2 \pi}} \int_{e^{-Z}}^{\infty} \frac{e^{-\frac{x}{2}}}{\sqrt{x}} d x= \\
& =\frac{1}{\sqrt{\pi}} \int_{e^{-Z / 2}}^{\infty} \frac{e^{-y}}{\sqrt{y}} d y=\frac{1}{\sqrt{\pi}} \Gamma\left(\frac{1}{2}, \frac{e^{-Z}}{2}\right)=\frac{\Gamma\left(\frac{1}{2}, \frac{e^{-Z}}{2}\right)}{\Gamma\left(\frac{1}{2}\right)}=1-P\left(\frac{1}{2}, \frac{e^{-Z}}{2}\right)
\end{aligned}
$$

where $P$ is the incomplete Gamma function.
Using the inverse of the cumulative function we find the median $\mathcal{M} \approx 0.78760$ and the lower and upper quartiles $\mathcal{Q}_{1} \approx-0.28013$ and $\mathcal{Q}_{3} \approx 2.28739$.

### 26.6 Random Number Generation

To obtain random numbers from the Moyal distribution we may either make use of the inverse to the incomplete Gamma function such that given a pseudorandom number $\xi$ we get a random number by solving the equation

$$
1-P\left(\frac{1}{2}, \frac{e^{-z}}{2}\right)=\xi
$$

for $z$. If $P^{-1}(a, p)$ denotes the value $x$ where $P(a, x)=p$ then

$$
z=-\ln \left\{2 P^{-1}\left(\frac{1}{2}, 1-\xi\right)\right\}
$$

is a random number from a Moyal distribution.
This is, however, a very slow method and one may instead use a straightforward rejectaccept (or hit-miss) method. To do this we prefer to transform the distribution to get it into a finite interval. For this purpose we make the transformation $\tan y=x$ giving

$$
h(y)=f(\tan y) \frac{1}{\cos ^{2} y}=\frac{1}{\sqrt{2 \pi}} \frac{1}{\cos ^{2} y} \exp \left\{-\frac{1}{2}\left(\tan y+e^{-\tan y}\right)\right\}
$$

This distribution, shown in figure 19, has a maximum of about 0.911 and is limited to the interval $-\frac{\pi}{2} \leq y \leq \frac{\pi}{2}$.

A simple algorithm to get random numbers from a Moyal distribution, either $f(z)$ or $g(x)$, using the reject-accept technique is as follows:
a Get into $\xi_{1}$ and $\xi_{2}$ two uniform random numbers uniformly distributed between zero and one using a good basic pseudorandom number generator.


Figure 19: Transformed Moyal distribution
b Calculate uniformly distributed variables along the horizontal and vertical direction by $y=\pi \xi_{1}-\frac{\pi}{2}$ and $h=\xi_{2} h_{\max }$ where $h_{\max }=0.912$ is chosen slightly larger than the maximum value of the function.
c Calculate $z=\tan y$ and the function value $h(y)$.
d If $h \leq h(y)$ then accept $z$ as a random number from the Moyal distribution $f(z)$ else go back to point a above.
e If required then scale and shift the result by $x=z \sigma+\mu$ in order to obtain a random number from $g(x)$.

This method is easily improved e.g. by making a more tight envelope to the distribution than a uniform distribution. The efficiency of the reject-accept technique outlined here is only $1 / 0.912 \pi \approx 0.35$ (the ratio between the area of the curve and the uniform distribution). The method seems, however, fast enough for most applications.

## 27 Multinomial Distribution

### 27.1 Introduction

The Multinomial distribution is given by

$$
p(\underline{r} ; N, k, \underline{p})=\frac{N!}{r_{1}!r_{2}!\cdots r_{k}!} p_{1}^{r_{1}} p_{2}^{r_{2}} \cdots p_{k}^{r_{k}}=N!\prod_{i=1}^{k} \frac{p_{i}^{r_{i}}}{r_{i}!}
$$

where the variable $\bar{r}$ is a vector with $k$ integer elements for which $0 \leq r_{i} \leq N$ and $\sum r_{i}=N$. The parameters $N>0$ and $k>2$ are integers and $\bar{p}$ is a vector with elements $0 \leq p_{i} \leq 1$ with the constraint that $\sum p_{i}=1$.

The distribution is a generalization of the Binomial distribution $(k=2)$ to many dimensions where, instead of two groups, the $N$ elements are divided into $k$ groups each with a probability $p_{i}$ with $i$ ranging from 1 to $k$. A common example is a histogram with $N$ entries in $k$ bins.

### 27.2 Histogram

The histogram example is valid when the total number of events $N$ is regarded as a fixed number. The variance in each bin then becomes, see also below, $V\left(r_{i}\right)=N p_{i}\left(1-p_{1}\right) \approx r_{i}$ if $p_{i} \ll 1$ which normally is the case for a histogram with many bins.

If, however, we may regard the total number of events $N$ as a random variable distributed according to the Poisson distribution we find: Given a multinomial distribution, here denoted $M(\underline{r} ; N, \underline{p})$, for the distribution of events into bins for fixed $N$ and a Poisson distribution, denoted $\bar{P}(N ; \nu)$, for the distribution of $N$ we write the joint distribution

$$
\begin{aligned}
\mathcal{P}(\underline{r}, N) & =M(\underline{r} ; N, \underline{p}) P(N ; \nu)=\left(\frac{N!}{r_{1}!r_{2}!\ldots r_{k}!} p_{1}^{r_{1}} p_{2}^{r_{2}} \ldots p_{k}^{r_{k}}\right)\left(\frac{\nu^{N} e^{-\nu}}{N!}\right)= \\
& =\left(\frac{1}{r_{1}!}\left(\nu p_{1}\right)^{r_{1}} e^{-\nu p_{1}}\right)\left(\frac{1}{r_{2}!}\left(\nu p_{2}\right)^{r_{2}} e^{-\nu p_{2}}\right) \ldots\left(\frac{1}{r_{k}!}\left(\nu p_{k}\right)^{r_{k}} e^{-\nu p_{k}}\right)
\end{aligned}
$$

where we have used that

$$
\sum_{i=1}^{k} p_{i}=1 \quad \text { and } \quad \sum_{i=1}^{k} r_{i}=N
$$

i.e. we get a product of independent Poisson distributions with means $\nu p_{i}$ for each individual bin.

As seen, in both cases, we find justification for the normal rule of thumb to assign the square root of the bin contents as the error in a certain bin. Note, however, that in principle we should insert the true value of $r_{i}$ for this error. Since this normally is unknown we use the observed number of events in accordance with the law of large numbers. This means that caution must be taken in bins with few entries.

### 27.3 Moments

For each specific $r_{i}$ we may obtain moments using the Binomial distribution with $q_{i}=1-p_{i}$

$$
E\left(r_{i}\right)=N p_{i} \quad \text { and } \quad V\left(r_{i}\right)=N p_{i}\left(1-p_{i}\right)=N p_{i} q_{i}
$$

The covariance between two groups are given by

$$
\operatorname{Cov}\left(r_{i}, r_{j}\right)=-N p_{i} p_{j} \quad \text { for } \quad i \neq j
$$

### 27.4 Probability Generating Function

The probability generating function for the multinomial distribution is given by

$$
G(\underline{z})=\left(\sum_{i=1}^{k} p_{i} z_{i}\right)^{N}
$$

### 27.5 Random Number Generation

The straightforward but time consuming way to generate random numbers from a multinomial distribution is to follow the definition and generate $N$ uniform random numbers which are assigned to specific bins according to the cumulative value of the $p$-vector.

### 27.6 Significance Levels

To determine a significance level for a certain outcome from a multinomial distribution one may add all outcomes which are as likely or less likely than the probability of the observed outcome. This may be a non-trivial calculation for large values of $N$ since the number of possible outcomes grows very fast. An alternative, although quite clumsy, is to generate a number of multinomial random numbers and evaluate how often these outcomes are as likely or less likely than the observed one.

If we as an example observe the outcome $r=(4,1,0,0,0,0)$ for a case with 5 observations in 6 groups ( $N=5$ and $k=6$ ) and the probability for all groups are the same $p_{i}=1 / k=1 / 6$ we obtain a probability of $p \approx 0.02$. This includes all orderings of the same outcome since these are all equally probable but also all less likely outcomes of the type $p=(5,0,0,0,0,0)$.

If a probability calculated in this manner is too small one may conclude that the null hypothesis that all probabilities are equal is wrong. Thus if our confidence level is preset to $95 \%$ this conclusion would be drawn in the above example. Of course, the conclusion would be wrong in $2 \%$ of all cases.

### 27.7 Equal Group Probabilities

A common case or null hypothesis for a multinomial distribution is that the probability of the $k$ groups is the same i.e. $p=1 / k$. In this case the multinomial distribution is simplified and since ordering become insignificant much fewer unique outcomes are possible.

Take as an example a game where five dices are thrown. The probabilities for different outcomes may quite readily be evaluated from basic probability theory properly accounting for the $6^{5}=7776$ possible outcomes. But one may also use the multinomial distribution with $k=6$ and $N=5$ to find probabilities for different outcomes. If we properly take care of combinatorial coefficients for each outcome we obtain (with zeros for empty groups suppressed)

| name | outcome | \# combinations | probability |
| :--- | :--- | :---: | :---: |
| one doublet | $2,1,1,1$ | 3600 | 0.46296 |
| two doublets | $2,2,1$ | 1800 | 0.23148 |
| triplets | $3,1,1$ | 1200 | 0.15432 |
| nothing | $1,1,1,1,1$ | 720 | 0.09259 |
| full house | 3,2 | 300 | 0.03858 |
| quadruplets | 4,1 | 150 | 0.01929 |
| quintuplets | 5 | 6 | 0.00077 |
| total |  | 7776 | 1.00000 |

The experienced dice player may note that the "nothing" group includes 240 combinations giving straights ( 1 to 5 or 2 to 6 ). From this table we may verify the statement from the previous subsection that the probability to get an outcome with quadruplets or less likely outcomes is given by 0.02006 .

Generally we have for $N<k$ that the two extremes of either all observations in separate groups $p_{\text {sep }}$ or all observations in one group $p_{\text {all }}$

$$
\begin{aligned}
p_{\text {sep }} & =\frac{k!}{k^{N}(k-N)!}=\frac{k}{k} \cdot \frac{k-1}{k} \cdots \frac{k-N+1}{k} \\
p_{\text {all }} & =\frac{1}{k^{N-1}}
\end{aligned}
$$

which we could have concluded directly from a quite simple probability calculation.
The first case is the formula which shows the quite well known fact that if 23 people or more are gathered the probability that at least two have the same birthday, i.e. $1-p_{\text {sep }}$, is greater than $50 \%$ (using $N=23$ and $k=365$ and not bothering about leap-years or possible deviations from the hypothesis of equal probabilities for each day). This somewhat nonintuitive result becomes even more pronounced for higher values of $k$ and the level above which $p_{\text {sep }}<0.5$ is approximately given by

$$
N \approx 1.2 \sqrt{k}
$$

For higher significance levels we may note that in the case with $k=365$ the probability $1-p_{\text {sep }}$ becomes greater than $90 \%$ at $N=41$, greater than $99 \%$ at $N=57$ and greater than $99.9 \%$ at $N=70$ i.e. already for $N \ll k$ a bet would be almost certain.

In Fig. 20 we show, in linear scale to the left and logarithmic scale to the right, the lower limit on $N$ for which the probability to have $1-p_{\text {sep }}$ above $50 \%, 90 \%, 99 \%$ and $99.9 . \%$ for $k$-values ranging up to 1000 . By use of the gamma function the problem has been generalized to real numbers. Note that the curves start at certain values where $k=N$ since for $N>k$ it is impossible to have all events in separate groups ${ }^{5}$.

[^4]

Figure 20: Limits for $N$ at several confidence levels as a function of $k$ (linear scale to the left and logarithmic scale to the right).

## 28 Multinormal Distribution

### 28.1 Introduction

As a generalization of the normal or Gauss distribution to many dimensions we define the multinormal distribution.

A multinormal distribution in $x=\left\{x_{1}, x_{2}, \ldots, x_{n}\right\}$ with parameters $\mu$ (mean vector) and $V$ (variance matrix) is given by

$$
f(x \mid \mu, V)=\frac{e^{-\frac{1}{2}(x-\mu) V^{-1}(x-\mu)^{T}}}{(2 \pi)^{\frac{n}{2}} \sqrt{|V|}}
$$

The variance matrix $V$ has to be a positive semi-definite matrix in order for $f$ to be a proper probability density function (necessary in order that the normalization integral $\int f(x) d x$ should converge).

If $x$ is normal and $V$ non-singular then $(x-\mu) V^{-1}(x-\mu)^{T}$ is called the covariance form of $x$ and has a $\chi^{2}$-distribution with $n$ degrees of freedom. Note that the distribution has constant probability density for constant values of the covariance form.

The characteristic function is given by

$$
\phi(t)=e^{\imath t \mu-\frac{1}{2} t^{T} V t}
$$

where $t$ is a vector of length $n$.

### 28.2 Conditional Probability Density

The conditional density for a fixed value of any $x_{i}$ is given by a multinormal density with $n-1$ dimensions where the new variance matrix is obtained by deleting the $i$ :th row and column of $V^{-1}$ and inverting the resulting matrix.

This may be compared to the case where we instead just want to neglect one of the variables $x_{i}$. In this case the remaining variables has a multinormal distribution with $n-1$ dimensions with a variance matrix obtained by deleting the $i$ :th row and column of $V$.

### 28.3 Probability Content

As discussed in section 6.6 on the binormal distribution the joint probability content of a multidimensional normal distribution is different, and smaller, than the corresponding well known figures for the one-dimensional normal distribution. In the case of the binormal distribution the ellipse (see figure 2 on page 20) corresponding to one standard deviation has a joint probability content of $39.3 \%$.

The same is even more true for the probability content within the hyperellipsoid in the case of a multinormal distribution. In the table below we show, for different dimensions $n$, the probability content for the one (denoted $z=1$ ), two and three standard deviation contours. We also give $z$-values $z_{1}, z_{2}$, and $z_{3}$ adjusted to give a probability content within the hyperellipsoid corresponding to the one-dimensional one, two, and three standard deviation contents ( $68.3 \%, 95.5 \%$, and $99.7 \%$ ). Finally $z$-value corresponding to joint probability
contents of $90 \%, 95 \%$ and $99 \%$ in $z_{90}, z_{95}$, and $z_{99}$, respectively, are given. Note that these probability contents are independent of the variance matrix which only has the effect to change the shape of the hyperellipsoid from a perfect hypersphere with radius $z$ when all variables are uncorrelated to e.g. cigar shapes when correlations are large.

Note that this has implications on errors estimated from a chi-square or a maximum likelihood fit. If a multiparameter confidence limit is requested and the chi-square minimum is at $\chi_{\text {min }}^{2}$ or the logarithmic likelihood maximum at $\ln \mathcal{L}_{\text {max }}$, one should look for the error contour at $\chi_{\text {min }}^{2}+z^{2}$ or $\ln \mathcal{L}_{\text {max }}-z^{2} / 2$ using a $z$-value from the right-hand side of the table below. The probability content for a $n$-dimensional multinormal distribution as given below may be expressed in terms of the incomplete Gamma function by

$$
p=P\left(\frac{n}{2}, \frac{z^{2}}{2}\right)
$$

as may be deduced by integrating a standard multinormal distribution out to a radius $z$. Special formulæ for the incomplete Gamma function $P(a, x)$ for integer and half-integer $a$ are given in section 42.5.3.

|  | Probability content in $\%$ |  |  |  |  |  |  |  | Adjusted $z$-values |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $n$ | $z=1$ | $z=2$ | $z=3$ | $z_{1}$ | $z_{2}$ | $z_{3}$ | $z_{90}$ | $z_{95}$ | $z_{99}$ |  |  |  |  |  |
| 1 | 68.27 | 95.45 | 99.73 | 1.000 | 2.000 | 3.000 | 1.645 | 1.960 | 2.576 |  |  |  |  |  |
| 2 | 39.35 | 86.47 | 98.89 | 1.515 | 2.486 | 3.439 | 2.146 | 2.448 | 3.035 |  |  |  |  |  |
| 3 | 19.87 | 73.85 | 97.07 | 1.878 | 2.833 | 3.763 | 2.500 | 2.795 | 3.368 |  |  |  |  |  |
| 4 | 9.020 | 59.40 | 93.89 | 2.172 | 3.117 | 4.031 | 2.789 | 3.080 | 3.644 |  |  |  |  |  |
| 5 | 3.743 | 45.06 | 89.09 | 2.426 | 3.364 | 4.267 | 3.039 | 3.327 | 3.884 |  |  |  |  |  |
| 6 | 1.439 | 32.33 | 82.64 | 2.653 | 3.585 | 4.479 | 3.263 | 3.548 | 4.100 |  |  |  |  |  |
| 7 | 0.517 | 22.02 | 74.73 | 2.859 | 3.786 | 4.674 | 3.467 | 3.751 | 4.298 |  |  |  |  |  |
| 8 | 0.175 | 14.29 | 65.77 | 3.050 | 3.974 | 4.855 | 3.655 | 3.938 | 4.482 |  |  |  |  |  |
| 9 | 0.0562 | 8.859 | 56.27 | 3.229 | 4.149 | 5.026 | 3.832 | 4.113 | 4.655 |  |  |  |  |  |
| 10 | 0.0172 | 5.265 | 46.79 | 3.396 | 4.314 | 5.187 | 3.998 | 4.279 | 4.818 |  |  |  |  |  |
| 11 | 0.00504 | 3.008 | 37.81 | 3.556 | 4.471 | 5.340 | 4.156 | 4.436 | 4.972 |  |  |  |  |  |
| 12 | 0.00142 | 1.656 | 29.71 | 3.707 | 4.620 | 5.486 | 4.307 | 4.585 | 5.120 |  |  |  |  |  |
| 13 | 0.00038 | 0.881 | 22.71 | 3.853 | 4.764 | 5.626 | 4.451 | 4.729 | 5.262 |  |  |  |  |  |
| 14 | 0.00010 | 0.453 | 16.89 | 3.992 | 4.902 | 5.762 | 4.590 | 4.867 | 5.398 |  |  |  |  |  |
| 15 | 0.00003 | 0.226 | 12.25 | 4.126 | 5.034 | 5.892 | 4.723 | 5.000 | 5.530 |  |  |  |  |  |
| 16 | 0.00001 | 0.1097 | 8.659 | 4.256 | 5.163 | 6.018 | 4.852 | 5.128 | 5.657 |  |  |  |  |  |
| 17 | $\approx 0$ | 0.0517 | 5.974 | 4.382 | 5.287 | 6.140 | 4.977 | 5.252 | 5.780 |  |  |  |  |  |
| 18 | $\approx 0$ | 0.0237 | 4.026 | 4.503 | 5.408 | 6.259 | 5.098 | 5.373 | 5.900 |  |  |  |  |  |
| 19 | $\approx 0$ | 0.0106 | 2.652 | 4.622 | 5.525 | 6.374 | 5.216 | 5.490 | 6.016 |  |  |  |  |  |
| 20 | $\approx 0$ | 0.00465 | 1.709 | 4.737 | 5.639 | 6.487 | 5.330 | 5.605 | 6.129 |  |  |  |  |  |
| 25 | $\approx 0$ | 0.00005 | 0.1404 | 5.272 | 6.170 | 7.012 | 5.864 | 6.136 | 6.657 |  |  |  |  |  |
| 30 | $\approx 0$ | $\approx 0$ | 0.0074 | 5.755 | 6.650 | 7.486 | 6.345 | 6.616 | 7.134 |  |  |  |  |  |

### 28.4 Random Number Generation

In order to obtain random numbers from a multinormal distribution we proceed as follows:

- If $x=\left\{x_{1}, x_{2}, \ldots, x_{n}\right\}$ is distributed multinormally with mean 0 (zero vector) and variance matrix $I$ (unity matrix) then each $x_{i}(i=1,2, \ldots, n)$ can be found independently from a standard normal distribution.
- If $x$ is multinormally distributed with mean $\mu$ and variance matrix $V$ then any linear combination $y=S x$ is also multinormally distributed with mean $S \mu$ and variance matrix $S V S^{T}$,
- If we want to generate vectors, $y$, from a multinormal distribution with mean $\mu$ and variance matrix $V$ we may make a so called Cholesky decomposition of $V$, i.e. we find a triangular matrix $S$ such that $V=S S^{T}$. We then calculate $y=S x+\mu$ with the components of $x$ generated independently from a standard normal distribution.

Thus we have found a quite nice way of generating multinormally distributed random numbers which is important in many simulations where correlations between variables may not be ignored. If many random numbers are to be generated for multinormal variables from the same distribution it is beneficial to make the Cholesky decomposition once and store the matrix $S$ for further usage.

## 29 Negative Binomial Distribution

### 29.1 Introduction

The Negative Binomial distribution is given by

$$
p(r ; k, p)=\binom{r-1}{k-1} p^{k}(1-p)^{r-k}
$$

where the variable $r \geq k$ and the parameter $k>0$ are integers and the parameter $p$ $(0 \leq p \leq 1)$ is a real number.

The distribution expresses the probability of having to wait exactly $r$ trials until $k$ successes have occurred if the probability of a success in a single trial is $p$ (probability of failure $q=1-p$ ).

The above form of the Negative Binomial distribution is often referred to as the Pascal distribution after the french mathematician, physicist and philosopher Blaise Pascal (16231662).

The distribution is sometimes expressed in terms of the number of failures occurring while waiting for $k$ successes, $n=r-k$, in which case we write

$$
p(n ; k, p)=\binom{n+k-1}{n} p^{k}(1-p)^{n}
$$

where the new variable $n \geq 0$.
Changing variables, for this last form, to $\bar{n}$ and $k$ instead of $p$ and $k$ we sometimes use

$$
p(n ; \bar{n}, k)=\binom{n+k-1}{n} \frac{\bar{n}^{n} k^{k}}{(\bar{n}+k)^{n+k}}=\binom{n+k-1}{n}\left(\frac{\bar{n}}{\bar{n}+k}\right)^{n}\left(\frac{k}{\bar{n}+k}\right)^{k}
$$

The distribution may also be generalized to real values of $k$, although this may seem obscure from the above probability view-point ("fractional success"), writing the binomial coefficient as $(n+k-1)(n+k-2) \cdots(k+1) k / n!$.

### 29.2 Moments

In the first form given above the expectation value, variance, third and fourth central moments of the distribution are

$$
E(r)=\frac{k}{p}, \quad V(r)=\frac{k q}{p^{2}}, \quad \mu_{3}=\frac{k q(2-p)}{p^{3}}, \quad \text { and } \quad \mu_{4}=\frac{k q\left(p^{2}-6 p+6+3 k q\right)}{p^{4}}
$$

The coefficients of skewness and kurtosis are

$$
\gamma_{1}=\frac{2-p}{\sqrt{k q}} \quad \text { and } \quad \gamma_{2}=\frac{p^{2}-6 p+6}{k q}
$$

In the second formulation above, $p(n)$, the only difference is that the expectation value becomes

$$
E(n)=E(r)-k=\frac{k(1-p)}{p}=\frac{k q}{p}
$$

while higher moments remain unchanged as they should since we have only shifted the scale by a fixed amount.

In the last form given, using the parameters $\bar{n}$ and $k$, the expectation value and the variance are

$$
E(n)=\bar{n} \quad \text { and } \quad V(n)=\bar{n}+\frac{\bar{n}^{2}}{k}
$$

### 29.3 Probability Generating Function

The probability generating function is given by

$$
G(z)=\left(\frac{p z}{1-z q}\right)^{k}
$$

in the first case $(p(r))$ and

$$
G(z)=\left(\frac{p}{1-z q}\right)^{k}=\left(\frac{1}{1+(1-z) \frac{\bar{n}}{k}}\right)^{k}
$$

in the second case $(p(n))$ for the two different parameterizations.

### 29.4 Relations to Other Distributions

There are several interesting connections between the Negative Binomial distribution and other standard statistical distributions. In the following subsections we briefly address some of these connections.

### 29.4.1 Poisson Distribution

Regard the negative binomial distribution in the form

$$
p(n ; \bar{n}, k)=\binom{n+k-1}{n}\left(\frac{1}{1+\bar{n} / k}\right)^{k}\left(\frac{\bar{n} / k}{1+\bar{n} / k}\right)^{n}
$$

where $n \geq 0, k>0$ and $\bar{n}>0$.
As $k \rightarrow \infty$ the three terms become

$$
\begin{gathered}
\binom{n+k-1}{n}=\frac{(n+k-1)(n+k-2) \ldots k}{n!} \rightarrow \frac{k^{n}}{n!} \\
\left(\frac{1}{1+\bar{n} / k}\right)^{k}=1-k \frac{\bar{n}}{k}+\frac{k(k+1)}{2}\left(\frac{\bar{n}}{k}\right)^{2}-\frac{k(k+1)(k+2)}{6}\left(\frac{\bar{n}}{k}\right)^{3}+\ldots \rightarrow e^{-\bar{n}} \text { and } \\
k^{n}\left(\frac{\bar{n} / k}{1+\bar{n} / k}\right)^{n} \rightarrow \bar{n}^{n}
\end{gathered}
$$

where, for the last term we have incorporated the factor $k^{n}$ from the first term.
Thus we have shown that

$$
\lim _{k \rightarrow \infty} p(n ; \bar{n}, k)=\frac{\bar{n}^{n} e^{-\bar{n}}}{n!}
$$

i.e. a Poisson distribution.

This "proof" could perhaps better be made using the probability generating function of the negative binomial distribution

$$
G(z)=\left(\frac{p}{1-z q}\right)^{k}=\left(\frac{1}{1-(z-1) \bar{n} / k}\right)^{k}
$$

Making a Taylor expansion of this for $(z-1) \bar{n} / k \ll 1$ we get

$$
G(z)=1+(z-1) \bar{n}+\frac{k+1}{k} \frac{(z-1)^{2} \bar{n}^{2}}{2}+\frac{(k+1)(k+2)}{k^{2}} \frac{(z-1)^{3} \bar{n}^{3}}{6}+\ldots \rightarrow e^{(z-1) \bar{n}}
$$

as $k \rightarrow \infty$. This result we recognize as the probability generating function of the Poisson distribution.

### 29.4.2 Gamma Distribution

Regard the negative binomial distribution in the form

$$
p(n ; k, p)=\binom{n+k-1}{n} p^{k} q^{n}
$$

where $n \geq 0, k>0$ and $0 \leq p \leq 1$ and where we have introduced $q=1-p$. If we change parameters from $k$ and $p$ to $k$ and $\bar{n}=k q / p$ this may be written

$$
p(n ; \bar{n}, k)=\binom{n+k-1}{n}\left(\frac{1}{1+\bar{n} / k}\right)^{k}\left(\frac{\bar{n} / k}{1+\bar{n} / k}\right)^{n}
$$

Changing variable from $n$ to $z=n / \bar{n}$ we get ( $d n / d z=\bar{n}$ )

$$
\begin{aligned}
p(z ; \bar{n}, k) & =p(n ; \bar{n}, k) \frac{d n}{d z}=\bar{n}\binom{z \bar{n}+k-1}{z \bar{n}}\left(\frac{1}{1+\bar{n} / k}\right)^{k}\left(\frac{\bar{n} / k}{1+\bar{n} / k}\right)^{z \bar{n}}= \\
& =\bar{n} \frac{(z \bar{n}+k-1)(z \bar{n}+k-2) \ldots(z \bar{n}+1)}{\Gamma(k)} k^{k}\left(\frac{1}{k+\bar{n}}\right)^{k}\left(\frac{1}{k / \bar{n}+1}\right)^{z \bar{n}} \rightarrow \\
& \rightarrow \bar{n} k^{k} \frac{(z \bar{n})^{k-1}}{\Gamma(k)}\left(\frac{1}{k+\bar{n}}\right)^{k}\left(\frac{1}{k / \bar{n}+1}\right)^{z \bar{n}}= \\
& =\frac{z^{k-1} k^{k}}{\Gamma(k)}\left(\frac{\bar{n}}{k+\bar{n}}\right)^{k}\left(\frac{1}{k / \bar{n}+1}\right)^{z \bar{n}} \rightarrow \frac{z^{k-1} k^{k} e^{-k z}}{\Gamma(k)}
\end{aligned}
$$

where we have used that for $k \ll \bar{n} \rightarrow \infty$

$$
\begin{aligned}
&\left(\frac{\bar{n}}{k+\bar{n}}\right)^{k} \rightarrow 1 \text { and } \\
&\left(\frac{1}{k / \bar{n}+1}\right)^{z \bar{n}}=1-z \bar{n} \frac{k}{\bar{n}}+\frac{z \bar{n}(z \bar{n}+1)}{2}\left(\frac{k}{\bar{n}}\right)^{2}-\frac{z \bar{n}(z \bar{n}+1)(z \bar{n}+2)}{6}\left(\frac{k}{\bar{n}}\right)^{3}+\ldots \rightarrow \\
& \rightarrow 1-z k+\frac{z^{2} k^{2}}{2}-\frac{z^{3} k^{3}}{6}+\ldots=e^{-k z}
\end{aligned}
$$

as $\bar{n} \rightarrow \infty$.
Thus we have "shown" that as $\bar{n} \rightarrow \infty$ and $\bar{n} \gg k$ we obtain a gamma distribution in the variable $z=n / \bar{n}$.

### 29.4.3 Logarithmic Distribution

Regard the negative binomial distribution in the form

$$
p(n ; k, p)=\binom{n+k-1}{n} p^{k} q^{n}
$$

where $n \geq 0, k>0$ and $0 \leq p \leq 1$ and where we have introduced $q=1-p$.
The probabilities for $n=0,1,2,3 \ldots$ are given by

$$
\{p(0), p(1), p(2), p(3), \ldots\}=p^{k}\left\{1, k q, \frac{k(k+1)}{2!} q^{2}, \frac{k(k+1)(k+2)}{3!} q^{3}, \ldots\right\}
$$

if we omit the zero class ( $\mathrm{n}=0$ ) and renormalize we get

$$
\frac{k p^{k}}{1-p^{k}}\left\{0, q, \frac{k+1}{2!} q^{2}, \frac{(k+1)(k+2)}{3!} q^{3}, \ldots\right\}
$$

and if we let $k \rightarrow 0$ we finally obtain

$$
-\frac{1}{\ln p}\left\{0, q, \frac{q^{2}}{2}, \frac{q^{3}}{3}, \ldots\right\}
$$

where we have used that

$$
\lim _{k \rightarrow 0} \frac{k}{p^{-k}-1}=-\frac{1}{\ln p}
$$

which is easily realized expanding $p^{-k}=e^{-k \ln p}$ into a power series.
This we recognize as the logarithmic distribution

$$
p(n ; p)=-\frac{1}{\ln p} \frac{(1-p)^{n}}{n}
$$

thus we have shown that omitting the zero class and letting $k \rightarrow 0$ the negative binomial distribution becomes the logarithmic distribution.

### 29.4.4 Branching Process

In a process where a branching occurs from a Poisson to a logarithmic distribution the most elegant way to determine the resulting distribution is by use of the probability generating function. The probability generating functions for a Poisson distribution with parameter (mean) $\mu$ and for a logarithmic distribution with parameter $p(q=1-p)$ are given by

$$
G_{P}(z)=e^{\mu(z-1)} \quad \text { and } \quad G_{L}(z)=\ln (1-z q) / \ln (1-q)=\alpha \ln (1-z q)
$$

where $\mu>0,0 \leq q \leq 1$ and $\alpha=1 / \ln p$.
For a branching process in $n$ steps

$$
G(z)=G_{1}\left(G_{2}\left(\ldots G_{n-1}\left(G_{n}(z)\right) \ldots\right)\right)
$$

where $G_{k}(z)$ is the probability generating function in the $k$ :th step. In the above case this gives

$$
\begin{aligned}
G(z) & =G_{P}\left(G_{L}(z)\right)=\exp \{\mu(\alpha \ln (1-z q)-1)\}= \\
& =\exp \{\alpha \mu \ln (1-z q)-\mu\}=(1-z q)^{\alpha \mu} e^{-\mu}= \\
& =(1-z q)^{-k}(1-q)^{k}=p^{k} /(1-z q)^{k}
\end{aligned}
$$

where we have put $k=-\alpha \mu$. This we recognize as the probability generating function of a negative binomial distribution with parameters $k$ and $p$.

We have thus shown that a Poisson distribution with mean $\mu$ branching into a logarithmic distribution with parameter $p$ gives rise to a negative binomial distribution with parameters $k=-\alpha \mu=-\mu / \ln p$ and $p($ or $\bar{n}=k q / p)$.

Conversely a negative binomial distribution with parameters $k$ and $p$ or $\bar{n}$ could arise from the combination of a Poisson distribution with parameter $\mu=-k \ln p=k \ln \left(1+\frac{\bar{n}}{k}\right)$ and a logarithmic distribution with parameter $p$ and mean $\bar{n} / \mu$.

A particle physics example would be a charged multiplicity distribution arising from the production of independent clusters subsequently decaying into charged particles according to a logarithmic distribution. The UA5 experiment [36] found on the $S \bar{p} p S$ collider at CERN that at a centre of mass energy of 540 GeV a negative binomial distribution with $\bar{n}=28.3$ and $k=3.69$ fitted the data well. With the above scenario this would correspond to $\approx 8$ clusters being independently produced (Poisson distribution with $\mu=7.97$ ) each one decaying, according to a logarithmic distribution, into 3.55 charged particles on average.

### 29.4.5 Poisson and Gamma Distributions

If a Poisson distribution with mean $\mu>0$

$$
p(n ; \mu)=\frac{e^{-\mu} \mu^{n}}{n!} \quad \text { for } \quad n \geq 0
$$

is weighted by a gamma distribution with parameters $a>0$ and $b>0$

$$
f(x ; a, b)=\frac{a(a x)^{b-1} e^{-a x}}{\Gamma(b)} \quad \text { for } \quad x>0
$$

we obtain

$$
\begin{aligned}
\mathcal{P}(n) & =\int_{0}^{\infty} p(n ; \mu) f(\mu ; a, b) d \mu=\int_{0}^{\infty} \frac{e^{-\mu} \mu^{n}}{n!} \frac{a(a \mu)^{b-1} e^{-a \mu}}{\Gamma(b)} d \mu= \\
& =\frac{a^{b}}{n!\Gamma(b)} \int_{0}^{\infty} \mu^{n+b-1} e^{-\mu(a+1)} d \mu=\frac{a^{b}}{n!(b-1)!}(n+b-1)!(a+1)^{-(n+b)}= \\
& =\binom{n+b-1}{n}\left(\frac{a}{a+1}\right)^{b}\left(\frac{1}{a+1}\right)^{n}
\end{aligned}
$$

which is a negative binomial distribution with parameters $p=\frac{a}{a+1}$, i.e. $q=1-p=\frac{1}{a+1}$, and $k=b$. If we aim at a negative binomial distribution with parameters $\bar{n}$ and $k$ we should
thus weight a Poisson distribution with a gamma distribution with parameters $a=k / \bar{n}$ and $b=k$. This is the same as superimposing Poisson distributions with means coming from a gamma distribution with mean $\bar{n}$.

In the calculation above we have made use of integral tables for the integral

$$
\int_{0}^{\infty} x^{n} e^{-\alpha x} d x=n!\alpha^{-(n+1)}
$$

### 29.5 Random Number Generation

In order to obtain random numbers from a Negative Binomial distribution we may use the recursive formula

$$
p(r+1)=p(r) \frac{q r}{r+1-k} \quad \text { or } \quad p(n+1)=p(n) \frac{q(k+n)}{n+1}
$$

for $r=k, k+1, \ldots$ and $n=0,1, \ldots$ in the two cases starting with the first term $(p(k)$ or $p(0))$ being equal to $p^{k}$. This technique may be speeded up considerably, if $p$ and $k$ are constants, by preparing a cumulative vector once for all.

One may also use some of the relations described above such as the branching of a Poisson to a Logarithmic distribution ${ }^{6}$ or a Poisson distribution weighted by a Gamma distribution ${ }^{7}$. This, however, will always be less efficient than the straightforward cumulative technique.

[^5]
## 30 Non-central Beta-distribution

### 30.1 Introduction

The non-central Beta-distribution is given by

$$
f(x ; p, q)=\sum_{r=0}^{\infty} e^{-\frac{\lambda}{2}} \frac{\left(\frac{\lambda}{2}\right)^{r}}{r!} \frac{x^{p+r-1}(1-x)^{q-1}}{B(p+r, q)}
$$

where $p$ and $q$ are positive real quantities and the non-centrality parameter $\lambda \geq 0$.
In figure 21 we show examples of a non-central Beta distribution with $p=\frac{\overline{3}}{2}$ and $q=3$ varying the non-central parameter $\lambda$ from zero (an ordinary Beta distribution) to ten in steps of two.


Figure 21: Graph of non-central Beta-distribution for $p=\frac{3}{2}, q=3$ and some values of $\lambda$

### 30.2 Derivation of distribution

If $y_{m}$ and $y_{n}$ are two independent variables distributed according to the chi-squared distribution with $m$ and $n$ degrees of freedom, respectively, then the ratio $y_{m} /\left(y_{m}+y_{n}\right)$ follows a Beta distribution with parameters $p=\frac{m}{2}$ and $q=\frac{n}{2}$. If instead $y_{m}$ follows a non-central chi-square distribution we may proceed in a similar way as was done for the derivation of the Beta-distribution (see section 4.2).

We make a change of variables to $x=y_{m} /\left(y_{m}+y_{n}\right)$ and $y=y_{m}+y_{n}$ which implies that $y_{m}=x y$ and $y_{n}=y(1-x)$ obtaining

$$
\begin{aligned}
& f(x, y)=\left\|\begin{array}{ll}
\frac{\partial y_{m}}{\partial x} & \frac{\partial y_{m}}{\partial y} \\
\frac{\partial y_{n}}{\partial x} & \frac{\partial y_{n}}{\partial y}
\end{array}\right\| f\left(y_{m}, y_{n}\right)= \\
& =\left|\begin{array}{cc}
y & x \\
-y & 1-x
\end{array}\right|\left\{\sum_{r=0}^{\infty} e^{-\frac{\lambda}{2}} \frac{\left(\frac{\lambda}{2}\right)^{r}}{r!} \frac{\left(\frac{y_{m}}{2}\right)^{\frac{m}{2}+r-1} e^{-\frac{y_{m}}{2}}}{2 \Gamma\left(\frac{m}{2}+r\right)}\right\}\left\{\frac{\left(\frac{y_{n}}{2}\right)^{\frac{n}{2}-1} e^{-\frac{y_{n}}{2}}}{2 \Gamma\left(\frac{n}{2}\right)}\right\}=
\end{aligned}
$$

$$
\begin{aligned}
& =y\left\{\sum_{r=0}^{\infty} e^{-\frac{\lambda}{2}} \frac{\left(\frac{\lambda}{2}\right)^{r}}{r!} \frac{\left(\frac{x y}{2}\right)^{\frac{m}{2}+r-1} e^{-\frac{x y}{2}}}{2 \Gamma\left(\frac{m}{2}+r\right)}\right\}\left\{\frac{\left(\frac{y(1-x)}{2}\right)^{\frac{n}{2}-1} e^{-\frac{y(1-x)}{2}}}{2 \Gamma\left(\frac{n}{2}\right)}\right\}= \\
& =\sum_{r=0}^{\infty} e^{-\frac{\lambda}{2}} \frac{\left(\frac{\lambda}{2}\right)^{r}}{r!} \frac{x^{\frac{m}{2}+r-1}(1-x)^{\frac{n}{2}-1}}{B\left(\frac{m}{2}+r, \frac{n}{2}\right)}\left\{\frac{\left(\frac{y}{2}\right)^{\frac{m+n}{2}+r-1} e^{-\frac{y}{2}}}{2 \Gamma\left(\frac{m+n}{2}+r\right)}\right\}
\end{aligned}
$$

In the last braces we see a chi-square distribution in $y$ with $m+n+2 r$ degrees of freedom and integrating $f(x, y)$ over $y$ in order to get the marginal distribution in $x$ gives us the non-central Beta-distribution as given above with $p=m / 2$ and $q=n / 2$.

If instead $y_{n}$ were distributed as a non-central chi-square distribution we would get a very similar expression (not amazing since $y_{m} /\left(y_{m}+y_{n}\right)=1-y_{n} /\left(y_{m}+y_{n}\right)$ ) but it's the form obtained when $y_{m}$ is non-central, that is normally referred to as the non-central Beta-distribution.

### 30.3 Moments

Algebraic moments of the non-central Beta-distribution are given in terms of the hypergeometric function ${ }_{2} F_{2}$ as

$$
\begin{aligned}
E\left(x^{k}\right) & =\int_{0}^{1} x^{k} f(x ; p, q) d x=\int_{0}^{1} \sum_{r=0}^{\infty} e^{-\frac{\lambda}{2}} \frac{\left(\frac{\lambda}{2}\right)^{r}}{r!} \frac{x^{p+r+k-1}(1-x)^{q-1}}{B(p+r, q)} d x= \\
& =\sum_{r=0}^{\infty} e^{-\frac{\lambda}{2}} \frac{\left(\frac{\lambda}{2}\right)^{r}}{r!} \frac{B(p+r+k, q)}{B(p+r, q)}=\sum_{r=0}^{\infty} e^{-\frac{\lambda}{2}} \frac{\left(\frac{\lambda}{2}\right)^{r}}{r!} \frac{\Gamma(p+r+k)}{\Gamma(p+r)} \frac{\Gamma(p+r+q)}{\Gamma(p+r+q+k)}= \\
& =\sum_{r=0}^{\infty} e^{-\frac{\lambda}{2}} \frac{\left(\frac{\lambda}{2}\right)^{r}}{r!} \frac{(p+r+k-1) \cdots(p+r+1)(p+r)}{(p+q+r+k-1) \cdots(p+q+r+1)(p+q+r)}= \\
& =e^{-\frac{\lambda}{2}} \cdot \frac{\Gamma(p+k)}{\Gamma(p)} \cdot \frac{\Gamma(p+q)}{\Gamma(p+q+k)} \cdot{ }_{2} F_{2}\left(p+q, p+k ; p, p+q+k ; \frac{\lambda}{2}\right)
\end{aligned}
$$

However, to evaluate the hypergeometric function involves a summation so it is more efficient to directly use the penultimate expression above.

### 30.4 Cumulative distribution

The cumulative distribution is found by straightforward integration

$$
F(x)=\int_{0}^{x} \sum_{r=0}^{\infty} e^{-\frac{\lambda}{2}} \frac{\left(\frac{\lambda}{2}\right)^{r}}{r!} \frac{u^{p+r-1}(1-u)^{q-1}}{B(p+r, q)} d u=\sum_{r=0}^{\infty} e^{-\frac{\lambda}{2}} \frac{\left(\frac{\lambda}{2}\right)^{r}}{r!} I_{x}(p+r, q)
$$

### 30.5 Random Number Generation

Random numbers from a non-central Beta-distribution with integer or half-integer $p-$ and $q$-values is easily obtained using the definition above i.e. by using a random number from a non-central chi-square distribution and another from a (central) chi-square distribution.

## 31 Non-central Chi-square Distribution

### 31.1 Introduction

If we instead of adding squares of $n$ independent standard normal, $N(0,1)$, variables, giving rise to the chi-square distribution with $n$ degrees of freedom, add squares of $N\left(\mu_{i}, 1\right)$ variables we obtain the non-central chi-square distribution

$$
f(x ; n, \lambda)=\sum_{r=0}^{\infty} e^{-\frac{\lambda}{2}} \frac{\left(\frac{\lambda}{2}\right)^{r}}{r!} f(x ; n+2 r)=\frac{1}{2^{\frac{n}{2}} \Gamma\left(\frac{1}{2}\right)} x^{\frac{n}{2}-1} e^{-\frac{1}{2}(x+\lambda)} \sum_{r=0}^{\infty} \frac{(\lambda x)^{r}}{(2 r)!} \frac{\Gamma\left(\frac{1}{2}+r\right)}{\Gamma\left(\frac{n}{2}+r\right)}
$$

where $\lambda=\sum \mu_{i}^{2}$ is the non-central parameter and $f(x ; n)$ the ordinary chi-square distribution. As for the latter the variable $x \geq 0$ and the parameter $n$ a positive integer. The additional parameter $\lambda \geq 0$ and in the limit $\lambda=0$ we retain the ordinary chi-square distribution. According to [2] pp 227-229 the non-central chi-square distribution was first introduced by R. A. Fisher in 1928. In figure 22 we show the distribution for $n=5$ and non-central parameter $\lambda=0,1,2,3,4,5$ (zero corresponding to the ordinary chi-squared distribution).


Figure 22: Graph of non-central chi-square distribution for $n=5$ and some values of $\lambda$

### 31.2 Characteristic Function

The characteristic function for the non-central chi-square distribution is given by

$$
\phi(t)=\frac{\exp \left(\frac{\imath t \lambda}{1-2 t t}\right)}{(1-2 \imath t)^{\frac{n}{2}}}
$$

but even more useful in determining moments is

$$
\ln \phi(t)=\frac{\imath t \lambda}{1-2 \imath t}-\frac{n}{2} \ln (1-2 \imath t)
$$

from which cumulants may be determined in a similar manner as we normally obtain algebraic moments from $\phi(t)$ (see below).

By looking at the characteristic function one sees that the sum of two non-central chisquare variates has the same distribution with degrees of freedoms as well as non-central parameters being the sum of the corresponding parameters for the individual distributions.

### 31.3 Moments

To use the characteristic function to obtain algebraic moments is not trivial but the cumulants (see section 2.5) are easily found to be given by the formula

$$
\kappa_{r}=2^{r-1}(r-1)!(n+r \lambda) \quad \text { for } \quad r \geq 1
$$

from which we may find the lower order algebraic and central moments (with $a=n+\lambda$ and $b=\lambda / a$ ) as

$$
\begin{aligned}
\mu_{1}^{\prime} & =\kappa_{1}=a=n+\lambda \\
\mu_{2} & =\kappa_{2}=2 a(1+b)=2(n+2 \lambda) \\
\mu_{3} & =\kappa_{3}=8 a(1+2 b)=8(n+3 \lambda) \\
\mu_{4} & =\kappa_{4}+3 \kappa_{2}^{2}=48(n+4 \lambda)+12(n+2 \lambda)^{2} \\
\mu_{5} & =\kappa_{5}+10 \kappa_{3} \kappa_{2}=384(n+5 \lambda)+160(n+2 \lambda)(n+3 \lambda) \\
\mu_{6} & =\kappa_{6}+15 \kappa_{4} \kappa_{2}+10 \kappa_{3}^{2}+15 \kappa_{2}^{3}= \\
& =3840(n+6 \lambda)+1440(n+2 \lambda)(n+4 \lambda)+640(n+3 \lambda)^{2}+120(n+2 \lambda)^{3} \\
\gamma_{1} & =\left(\frac{2}{1+b}\right)^{\frac{3}{2}} \cdot \frac{1+2 b}{\sqrt{a}}=\frac{8(n+3 \lambda)}{[2(n+2 \lambda)]^{\frac{3}{2}}} \\
\gamma_{2} & =\frac{12}{a} \cdot \frac{1+3 b}{(1+b)^{2}}=\frac{12(n+4 \lambda)}{(n+2 \lambda)^{2}}
\end{aligned}
$$

### 31.4 Cumulative Distribution

The cumulative, or distribution, function may be found by

$$
\begin{aligned}
F(x) & =\frac{1}{2^{\frac{n}{2}} \Gamma\left(\frac{1}{2}\right)} e^{-\frac{\lambda}{2}} \sum_{r=0}^{\infty} \frac{\lambda^{r}}{(2 r)!} \frac{\Gamma\left(\frac{1}{2}+r\right)}{\Gamma\left(\frac{n}{2}+r\right)} \int_{0}^{x} u^{\frac{n}{2}+r-1} e^{-\frac{u}{2}} d u= \\
& =\frac{1}{2^{\frac{n}{2}} \Gamma\left(\frac{1}{2}\right)} e^{-\frac{\lambda}{2}} \sum_{r=0}^{\infty} \frac{\lambda^{r}}{(2 r)!} \frac{\Gamma\left(\frac{1}{2}+r\right)}{\Gamma\left(\frac{n}{2}+r\right)} 2^{\frac{n}{2}+r} \gamma\left(\frac{n}{2}+r, \frac{x}{2}\right)= \\
& =e^{-\frac{\lambda}{2}} \sum_{r=0}^{\infty} \frac{\left(\frac{\lambda}{2}\right)^{r}}{r!} P\left(\frac{n}{2}+r, \frac{x}{2}\right)
\end{aligned}
$$

### 31.5 Approximations

An approximation to a chi-square distribution is found by equating the first two cumulants of a non-central chi-square distribution with those of $\rho$ times a chi-square distribution.

Here $\rho$ is a constant to be determined. The result is that with

$$
\rho=\frac{n+2 \lambda}{n+\lambda}=1+\frac{\lambda}{n+\lambda} \quad \text { and } \quad n^{*}=\frac{(n+\lambda)^{2}}{n+2 \lambda}=n+\frac{\lambda^{2}}{n+2 \lambda}
$$

we may approximate a non-central chi-square distribution $f(x ; n, \lambda)$ with a (central) chisquare distribution in $x / \rho$ with $n^{*}$ degrees of freedom ( $n^{*}$ in general being fractional).

Approximations to the standard normal distribution are given using

$$
z=\sqrt{\frac{2 x}{1+b}}-\sqrt{\frac{2 a}{1+b}-1} \quad \text { or } \quad z=\frac{\left(\frac{x}{a}\right)^{\frac{1}{3}}-\left[1-\frac{2}{9} \cdot \frac{1+b}{a}\right]}{\sqrt{\frac{2}{9} \cdot \frac{1+b}{a}}}
$$

### 31.6 Random Number Generation

Random numbers from a non-central chi-square distribution is easily obtained using the definition above by e.g.

- Put $\mu=\sqrt{\lambda / n}$
- Sum $n$ random numbers from a normal distribution with mean $\mu$ and variance unity. Note that this is not a unique choice. The only requirement is that $\lambda=\sum \mu_{i}^{2}$.
- Return the sum as a random number from a non-central chi-square distribution with $n$ degrees of freedom and non-central parameter $\lambda$.

This ought to be sufficient for most applications but if needed more efficient techniques may easily be developed e.g. using more general techniques.

## 32 Non-central $F$-Distribution

### 32.1 Introduction

If $x_{1}$ is distributed according to a non-central chi-square distribution with $m$ degrees of freedom and non-central parameter $\lambda$ and $x_{2}$ according to a (central) chi-square distribution with $n$ degrees of freedom then, provided $x_{1}$ and $x_{2}$ are independent, the variable

$$
F^{\prime}=\frac{x_{1} / m}{x_{2} / n}
$$

is said to have a non-central $F$-distribution with $m, n$ degrees of freedom (positive integers) and non-central parameter $\lambda \geq 0$. As the non-central chi-square distribution it was first discussed by R. A. Fisher in 1928.

This distribution in $F^{\prime}$ may be written

$$
f\left(F^{\prime} ; m, n, \lambda\right)=e^{-\frac{\lambda}{2}} \sum_{r=0}^{\infty} \frac{1}{r!}\left(\frac{\lambda}{2}\right)^{r} \frac{\Gamma\left(\frac{m+n}{2}+r\right)}{\Gamma\left(\frac{m}{2}+r\right) \Gamma\left(\frac{n}{2}\right)}\left(\frac{m}{n}\right)^{\frac{m}{2}+r} \frac{\left(F^{\prime}\right)^{\frac{m}{2}}-1+r}{\left(1+\frac{m F^{\prime}}{n}\right)^{\frac{1}{2}(m+n)+r}}
$$

In figure 23 we show the non-central $F$-distribution for the case with $m=10$ and $n=5$ varying $\lambda$ from zero (an ordinary, central, $F$-distribution) to five.


Figure 23: Graph of non-central $F$-distribution for $m=10, n=5$ and some values of $\lambda$
When $m=1$ the non-central $F$-distribution reduces to a non-central $t^{2}$-distribution with $\delta^{2}=\lambda$. As $n \rightarrow \infty$ then $n F^{\prime}$ approaches a non-central chi-square distribution with $m$ degrees of freedom and non-central parameter $\lambda$.

### 32.2 Moments

Algebraic moments of the non-central $F$-distribution may be achieved by straightforward, but somewhat tedious, algebra as

$$
\begin{aligned}
E\left(F^{\prime k}\right) & =\int_{0}^{\infty} x^{k} f(x ; m, n, \lambda) d x= \\
& =e^{-\frac{\lambda}{2}}\left(\frac{n}{m}\right)^{k} \frac{\Gamma\left(\frac{n}{2}-k\right)}{\Gamma\left(\frac{n}{2}\right)} \sum_{r=0}^{\infty} \frac{1}{r!}\left(\frac{\lambda}{2}\right)^{r} \frac{\Gamma\left(\frac{m}{2}+r+k\right)}{\Gamma\left(\frac{m}{2}+r\right)}
\end{aligned}
$$

an expression which may be used to find lower order moments (defined for $n>2 k$ )

$$
\begin{aligned}
& E\left(F^{\prime}\right)=\frac{n}{m} \cdot \frac{m+\lambda}{n-2} \\
& E\left(F^{\prime 2}\right)=\left(\frac{n}{m}\right)^{2} \frac{1}{(n-2)(n-4)}\left\{\lambda^{2}+(2 \lambda+m)(m+2)\right\} \\
& E\left(F^{\prime 3}\right)=\left(\frac{n}{m}\right)^{3} \frac{1}{(n-2)(n-4)(n-6)} \cdot\left\{\lambda^{3}+3(m+4) \lambda^{2}+(3 \lambda+m)(m+4)(m+2)\right\} \\
& E\left(F^{\prime 4}\right)=\left(\frac{n}{m}\right)^{4} \frac{1}{(n-2)(n-4)(n-6)(n-8)} \cdot\left\{\lambda^{4}+4(m+6) \lambda^{3}+6(m+6)(m+4) \lambda^{2}+\right. \\
&+(4 \lambda+m)(m+6)(m+4)(m+2)\} \\
& V\left(F^{\prime}\right)=\left(\frac{n}{m}\right)^{2} \frac{2}{(n-2)(n-4)}\left\{\frac{(\lambda+m)^{2}}{n-2}+2 \lambda+m\right\}
\end{aligned}
$$

### 32.3 Cumulative Distribution

The cumulative, or distribution, function may be found by

$$
\begin{aligned}
F(x) & =\int_{0}^{x} u^{k} f(u ; m, n, \lambda) d u= \\
& =e^{-\frac{\lambda}{2}} \sum_{r=0}^{\infty} \frac{1}{r!}\left(\frac{\lambda}{2}\right)^{r} \frac{\Gamma\left(\frac{m+n}{2}+r\right)}{\Gamma\left(\frac{m}{2}+r\right) \Gamma\left(\frac{n}{2}\right)}\left(\frac{m}{n}\right)^{\frac{m}{2}+r} \int_{0}^{x} \frac{u^{\frac{m}{2}-1+r+k}}{\left(1+\frac{m u}{n}\right)^{\frac{m+n}{2}+r}} d u= \\
& =e^{-\frac{\lambda}{2}} \sum_{r=0}^{\infty} \frac{1}{r!}\left(\frac{\lambda}{2}\right)^{r} \frac{B_{q}\left(\frac{m}{2}+r, \frac{n}{2}\right)}{B\left(\frac{m}{2}+r, \frac{n}{2}\right)}=e^{-\frac{\lambda}{2}} \sum_{r=0}^{\infty} \frac{\left(\frac{\lambda}{2}\right)^{r}}{r!} I_{q}\left(\frac{m}{2}+r, \frac{n}{2}\right)
\end{aligned}
$$

with

$$
q=\frac{\frac{m x}{n}}{1+\frac{m x}{n}}
$$

### 32.4 Approximations

Using the approximation of a non-central chi-square distribution to a (central) chi-square distribution given in the previous section we see that

$$
\frac{m}{m+\lambda} F^{\prime}
$$

is approximately distributed according to a (central) $F$-distribution with $m^{*}=m+\frac{\lambda^{2}}{m+2 \lambda}$ and $n$ degrees of freedom.

Approximations to the standard normal distribution is achieved with

$$
\begin{aligned}
z_{1} & =\frac{F^{\prime}-E\left(F^{\prime}\right)}{\sqrt{V\left(F^{\prime}\right)}}=\frac{F^{\prime}-\frac{n(m+\lambda)}{m(n-2)}}{\frac{n}{m}\left[\frac{2}{(n-2)(n-4)}\left\{\frac{(m+\lambda)^{2}}{n-2}+m+2 \lambda\right\}\right]^{\frac{1}{2}}} \\
\text { or } \quad z_{2} & =\frac{\left(\frac{m F^{\prime}}{m+\lambda}\right)^{\frac{1}{3}}\left(1-\frac{2}{9 n}\right)-\left(1-\frac{2}{9} \cdot \frac{m+2 \lambda}{(m+\lambda)^{2}}\right)}{\left[\frac{2}{9} \cdot \frac{m+2 \lambda}{(m+\lambda)^{2}}+\frac{2}{9 n} \cdot\left(\frac{m F^{\prime}}{m+\lambda}\right)^{\frac{2}{3}}\right]^{\frac{1}{2}}}
\end{aligned}
$$

### 32.5 Random Number Generation

Random numbers from a non-central chi-square distribution is easily obtained using the definition above i.e. by using a random number from a non-central chi-square distribution and another from a (central) chi-square distribution.

## 33 Non-central $t$-Distribution

### 33.1 Introduction

If $x$ is distributed according to a normal distribution with mean $\delta$ and variance 1 and $y$ according to a chi-square distribution with $n$ degrees of freedom (independent of $x$ ) then

$$
t^{\prime}=\frac{x}{\sqrt{y / n}}
$$

has a non-central t-distribution with $n$ degrees of freedom (positive integer) and non-central parameter $\delta$ (real).

We may also write

$$
t^{\prime}=\frac{z+\delta}{\sqrt{w / n}}
$$

where $z$ is a standard normal variate and $w$ is distributed as a chi-square variable with $n$ degrees of freedom.

The distribution is given by (see comments on derivation in section below)

$$
f\left(t^{\prime} ; n, \delta\right)=\frac{e^{-\frac{\delta^{2}}{2}}}{\sqrt{n \pi} \Gamma\left(\frac{n}{2}\right)} \sum_{r=0}^{\infty} \frac{\left(t^{\prime} \delta\right)^{r}}{r!n^{\frac{r}{2}}}\left(1+\frac{t^{\prime 2}}{n}\right)^{-\frac{n+r+1}{2}} 2^{\frac{r}{2}} \Gamma\left(\frac{n+r+1}{2}\right)
$$

In figure 24 we show the non-central $t$-distribution for the case with $n=10$ varying $\delta$ from zero (an ordinary $t$-distribution) to five.


Figure 24: Graph of non-central $t$-distribution for $n=10$ and some values of $\delta$
This distribution is of importance in hypotheses testing if we are interested in the probability of committing a Type II error implying that we would accept an hypothesis although it was wrong, see discussion in section 38.11 on page 146.

### 33.2 Derivation of distribution

Not many text-books include a formula for the non-central $t$-distribution and some turns out to give erroneous expressions. A non-central $F$-distribution with $m=1$ becomes a
non-central $t^{2}$-distribution which then may be transformed to a non-central $t$-distribution. However, with this approach one easily gets into trouble for $t^{\prime}<0$. Instead we adopt a technique very similar to what is used in section 38.6 to obtain the normal (central) $t$-distribution from a $t$-ratio.

The difference in the non-central case is the presence of the $\delta$-parameter which introduces two new exponential terms in the equations to be solved. One is simply $\exp \left(-\delta^{2} / 2\right)$ but another factor we treat by a serial expansion leading to the p.d.f. above. This may not be the 'best' possible expression but empirically it works quite well.

### 33.3 Moments

With some effort the p.d.f. above may be used to calculate algebraic moments of the distribution yielding

$$
E\left(t^{\prime k}\right)=\frac{e^{-\frac{\delta^{2}}{2}}}{\sqrt{\pi} \Gamma\left(\frac{n}{2}\right)} \Gamma\left(\frac{n-k}{2}\right) n^{\frac{k}{2}} \sum_{r=0}^{\infty} \frac{\delta^{r} 2^{\frac{r}{2}}}{r!} \Gamma\left(\frac{r+k+1}{2}\right)
$$

where the sum should be made for odd (even) values of $r$ if $k$ is odd (even). This gives for low orders

$$
\begin{aligned}
& \mu_{1}^{\prime}=\sqrt{\frac{n}{2}} \frac{\Gamma\left(\frac{n-1}{2}\right)}{\Gamma\left(\frac{n}{2}\right)} \delta \\
& \mu_{2}^{\prime}=\frac{n \Gamma\left(\frac{n-2}{2}\right)}{2 \Gamma\left(\frac{n}{2}\right)}\left(1+\delta^{2}\right)=\frac{n}{n-2}\left(1+\delta^{2}\right) \\
& \mu_{3}^{\prime}=\frac{n^{\frac{3}{2}} \sqrt{2} \Gamma\left(\frac{n-3}{2}\right)}{4 \Gamma\left(\frac{n}{2}\right)} \delta\left(3+\delta^{2}\right) \\
& \mu_{4}^{\prime}=\frac{n^{2} \Gamma\left(\frac{n-4}{2}\right)}{4 \Gamma\left(\frac{n}{2}\right)}\left(\delta^{4}+6 \delta^{2}+3\right)=\frac{n^{2}}{(n-2)(n-4)}\left(\delta^{4}+6 \delta^{2}+3\right)
\end{aligned}
$$

from which expressions for central moments may be found e.g. the variance

$$
\mu_{2}=V\left(t^{\prime}\right)=\frac{n}{2}\left\{\left(1+\delta^{2}\right) \frac{2}{n-2}-\delta^{2}\left(\frac{\Gamma\left(\frac{n-1}{2}\right)}{\Gamma\left(\frac{n}{2}\right)}\right)^{2}\right\}
$$

### 33.4 Cumulative Distribution

The cumulative, or distribution, function may be found by

$$
\begin{aligned}
F(t) & =\frac{e^{-\frac{\delta^{2}}{2}}}{\sqrt{n \pi} \Gamma\left(\frac{n}{2}\right)} \sum_{r=0}^{\infty} \frac{\delta^{r}}{r!n^{\frac{r}{2}}} 2^{\frac{r}{2}} \Gamma\left(\frac{n+r+1}{2}\right) \int_{-\infty}^{t} \frac{u^{r}}{\left(1+\frac{u^{2}}{n}\right)^{\frac{n+r+1}{2}}} d u= \\
& =\frac{e^{-\frac{\delta^{2}}{2}}}{\sqrt{\pi} \Gamma\left(\frac{n}{2}\right)} \sum_{r=0}^{\infty} \frac{\delta^{r}}{r!} 2^{\frac{r}{2}-1} \Gamma\left(\frac{n+r+1}{2}\right)\left\{s_{1} B\left(\frac{r+1}{2}, \frac{n}{2}\right)+s_{2} B_{q}\left(\frac{r+1}{2}, \frac{n}{2}\right)\right\}=
\end{aligned}
$$

$$
=\frac{e^{-\frac{\delta^{2}}{2}}}{\sqrt{\pi}} \sum_{r=0}^{\infty} \frac{\delta^{r}}{r!} 2^{\frac{r}{2}-1} \Gamma\left(\frac{r+1}{2}\right)\left\{s_{1}+s_{2} I_{q}\left(\frac{r+1}{2}, \frac{n}{2}\right)\right\}
$$

where $s_{1}$ are and $s_{2}$ are signs differing between cases with positive or negative $t$ as well as odd or even $r$ in the summation. The sign $s_{1}$ is -1 if $r$ is odd and +1 if it is even while $s_{2}$ is +1 unless $t<0$ and $r$ is even in which case it is -1 .

### 33.5 Approximation

An approximation is given by

$$
z=\frac{t^{\prime}\left(1-\frac{1}{4 n}\right)-\delta}{\sqrt{1+\frac{t^{\prime 2}}{2 n}}}
$$

which is asymptotically distributed as a standard normal variable.

### 33.6 Random Number Generation

Random numbers from a non-central $t$-distribution is easily obtained using the definition above i.e. by using a random number from a normal distribution and another from a chisquare distribution. This ought to be sufficient for most applications but if needed more efficient techniques may easily be developed e.g. using more general techniques.

## 34 Normal Distribution

### 34.1 Introduction

The normal distribution or, as it is often called, the Gauss distribution is the most important distribution in statistics. The distribution is given by

$$
f\left(x ; \mu, \sigma^{2}\right)=\frac{1}{\sigma \sqrt{2 \pi}} e^{-\frac{1}{2}\left(\frac{x-\mu}{\sigma}\right)^{2}}
$$

where $\mu$ is a location parameter, equal to the mean, and $\sigma$ the standard deviation. For $\mu=0$ and $\sigma=1$ we refer to this distribution as the standard normal distribution. In many connections it is sufficient to use this simpler form since $\mu$ and $\sigma$ simply may be regarded as a shift and scale parameter, respectively. In figure 25 we show the standard normal distribution.


Figure 25: Standard normal distribution
Below we give some useful information in connection with the normal distribution. Note, however, that this is only a minor collection since there is no limit on important and interesting statistical connections to this distribution.

### 34.2 Moments

The expectation value of the distribution is $E(x)=\mu$ and the variance $V(x)=\sigma^{2}$.
Generally odd central moments vanish due to the symmetry of the distribution and even central moments are given by

$$
\mu_{2 r}=\frac{(2 r)!}{2^{r} r!} \sigma^{2 r}=(2 r-1)!!\sigma^{2 r}
$$

for $r \geq 1$.

It is sometimes also useful to evaluate absolute moments $E\left(|x|^{n}\right)$ for the normal distribution. To do this we make use of the integral

$$
\int_{-\infty}^{\infty} e^{-a x^{2}} d x=\sqrt{\frac{\pi}{a}}
$$

which if differentiated $k$ times with respect to $a$ yields

$$
\int_{-\infty}^{\infty} x^{2 k} e^{-a x^{2}} d x=\frac{(2 k-1)!!}{2^{k}} \sqrt{\frac{\pi}{a^{2 k+1}}}
$$

In our case $a=1 / 2 \sigma^{2}$ and since even absolute moments are identical to the algebraic moments it is enough to evaluate odd absolute moments for which we get

$$
E\left(|x|^{2 k+1}\right)=\frac{2}{\sigma \sqrt{2 \pi}} \int_{0}^{\infty} x^{2 k+1} e^{-\frac{x^{2}}{2 \sigma^{2}}} d x=\sqrt{\frac{2}{\pi}} \frac{\left(2 \sigma^{2}\right)^{k+1}}{2 \sigma} \int_{0}^{\infty} y^{k} e^{-y} d y
$$

The last integral we recognize as being equal to $k$ ! and we finally obtain the absolute moments of the normal distribution as

$$
E\left(|x|^{n}\right)= \begin{cases}(n-1)!!\sigma^{n} & \text { for } n=2 k \\ \sqrt{\frac{2}{\pi}} 2^{k} k!\sigma^{2 k+1} & \text { for } n=2 k+1\end{cases}
$$

The half-width at half-height of the normal distribution is given by $\sqrt{2 \ln 2} \sigma \approx 1.177 \sigma$ which may be useful to remember when estimating $\sigma$ using a ruler.

### 34.3 Cumulative Function

The distribution function, or cumulative function, may be expressed in term of the incomplete gamma function $P$ as

$$
F(z)= \begin{cases}\frac{1}{2}+\frac{1}{2} P\left(\frac{1}{2}, \frac{z^{2}}{2}\right) & \text { if } z \geq 0 \\ \frac{1}{2}-\frac{1}{2} P\left(\frac{1}{2}, \frac{z^{2}}{2}\right) & \text { if } z<0\end{cases}
$$

or we may use the error function $\operatorname{erf}(z / \sqrt{2})$ in place of the incomplete gamma function.

### 34.4 Characteristic Function

The characteristic function for the normal distribution is easily found from the general definition

$$
\phi(t)=E\left(e^{\imath t x}\right)=\exp \left\{\mu \imath t-\frac{1}{2} \sigma^{2} t^{2}\right\}
$$

### 34.5 Addition Theorem

The so called Addition theorem for normally distributed variables states that any linear combination of independent normally distributed random variables $x_{i}(i=1,2, \ldots, n)$ is also distributed according to the normal distribution.

If each $x_{i}$ is drawn from a normal distribution with mean $\mu_{i}$ and variance $\sigma_{i}^{2}$ then regard the linear combination

$$
\mathcal{S}=\sum_{i=1}^{n} a_{i} x_{i}
$$

where $a_{i}$ are real coefficients. Each term $a_{i} x_{i}$ has characteristic function

$$
\phi_{a_{i} x_{i}}(t)=\exp \left\{\left(a_{i} \mu_{i}\right) \imath t-\frac{1}{2}\left(a_{i}^{2} \sigma_{i}^{2}\right) t^{2}\right\}
$$

and thus $\mathcal{S}$ has characteristic function

$$
\phi_{\mathcal{S}}(t)=\prod_{i=1}^{n} \phi_{a_{i} x_{i}}(t)=\exp \left\{\left(\sum_{i=1}^{n} a_{i} \mu_{i}\right) \imath t-\frac{1}{2}\left(\sum_{i=1}^{n} a_{i}^{2} \sigma_{i}^{2}\right) t^{2}\right\}
$$

which is seen to be a normal distribution with mean $\sum a_{i} \mu_{i}$ and variance $\sum a_{i}^{2} \sigma_{i}^{2}$.

### 34.6 Independence of $\bar{x}$ and $s^{2}$

A unique property of the normal distribution is the independence of the sample statistics $\bar{x}$ and $s^{2}$, estimates of the mean and variance of the distribution. Recall that the definition of these quantities are

$$
\bar{x}=\frac{1}{n} \sum_{i=1}^{n} x_{i} \quad \text { and } \quad s^{2}=\frac{1}{n-1} \sum_{i=1}^{n}\left(x_{i}-\bar{x}\right)^{2}
$$

where $\bar{x}$ is an estimator of the true mean $\mu$ and $s^{2}$ is the usual unbiased estimator for the true variance $\sigma^{2}$.

For a population of $n$ events from a normal distribution $\bar{x}$ has the distribution $N\left(\mu, \sigma^{2} / n\right)$ and $(n-1) s^{2} / \sigma^{2}$ is distributed according to a chi-square distribution with $n-1$ degrees of freedom. Using the relation

$$
\sum_{i=1}^{n}\left(\frac{x_{i}-\mu}{\sigma}\right)^{2}=\frac{(n-1) s^{2}}{\sigma^{2}}+\left(\frac{\bar{x}-\mu}{\sigma / \sqrt{n}}\right)^{2}
$$

and creating the joint characteristic function for the variables $(n-1) s^{2} / \sigma^{2}$ and $(\sqrt{n}(\bar{x}-$ $\left.\mu) / \sigma^{2}\right)^{2}$ one may show that this function factorizes thus implying independence of these quantities and thus also of $\bar{x}$ and $s^{2}$.

In summary the "independence theorem" states that given $n$ independent random variables with identical normal distributions the two statistics $\bar{x}$ and $s^{2}$ are independent. Also conversely it holds that if the mean $\bar{x}$ and the variance $s^{2}$ of a random sample are independent then the population is normal.

### 34.7 Probability Content

The probability content of the normal distribution is often referred to in statistics. When the term one standard deviation is mentioned one immediately thinks in terms of a probability content of $68.3 \%$ within the symmetric interval from the value given.

Without loss of generality we may treat the standard normal distribution only since the transformation from a more general case is straightforward putting $z=(x-\mu) / \sigma$. In different situation one may want to find

- the probability content, two-side or one-sided, to exceed a certain number of standard deviations, or
- the number of standard deviations corresponding to a certain probability content.

In calculating this we need to evaluate integrals like

$$
\alpha=\int \frac{1}{\sqrt{2 \pi}} e^{-\frac{t^{2}}{2}} d t
$$

There are no explicit solution to this integral but it is related to the error function (see section 13) as well as the incomplete gamma function (see section 42).

$$
\frac{1}{\sqrt{2 \pi}} \int_{-z}^{z} e^{-\frac{z^{2}}{2}}=\operatorname{erf}\left(\frac{z}{\sqrt{2}}\right)=P\left(\frac{1}{2}, \frac{z^{2}}{2}\right)
$$

These relations may be used to calculate the probability content. Especially the error function is often available as a system function on different computers. Beware, however, that it seems to be implemented such that $\operatorname{erf}(z)$ is the symmetric integral from $-z$ to $z$ and thus the $\sqrt{2}$ factor should not be supplied. Besides from the above relations there are also excellent approximations to the integral which may be used.

In the tables below we give the probability content for exact $z$-values (left-hand table) as well as $z$-values for exact probability contents (right-hand table).


| 0.0 | 0.50000 | 0.00000 | 0.50000 |
| :--- | :--- | :--- | :---: |
| 0.5 | 0.69146 | 0.38292 | 0.30854 |
| 1.0 | 0.84134 | 0.68269 | 0.15866 |
| 1.5 | 0.93319 | 0.86639 | 0.06681 |
| 2.0 | 0.97725 | 0.95450 | 0.02275 |
| 2.5 | 0.99379 | 0.98758 | $6.210 \cdot 10^{-3}$ |
| 3.0 | 0.99865 | 0.99730 | $1.350 \cdot 10^{-3}$ |
| 3.5 | 0.99977 | 0.99953 | $2.326 \cdot 10^{-4}$ |
| 4.0 | 0.99997 | 0.99994 | $3.167 \cdot 10^{-5}$ |
| 4.5 | 1.00000 | 0.99999 | $3.398 \cdot 10^{-6}$ |
| 5.0 | 1.00000 | 1.00000 | $2.867 \cdot 10^{-7}$ |
| 6.0 | 1.00000 | 1.00000 | $9.866 \cdot 10^{-10}$ |
| 7.0 | 1.00000 | 1.00000 | $1.280 \cdot 10^{-12}$ |
| 8.0 | 1.00000 | 1.00000 | $6.221 \cdot 10^{-16}$ |


| 0.00000 | 0.5 | 0.0 | 0.5 |
| :--- | :--- | :--- | :--- |
| 0.25335 | 0.6 | 0.2 | 0.4 |
| 0.67449 | 0.75 | 0.5 | 0.25 |
| 0.84162 | 0.8 | 0.6 | 0.2 |
| 1.28155 | 0.9 | 0.8 | 0.1 |
| 1.64485 | 0.95 | 0.9 | 0.05 |
| 1.95996 | 0.975 | 0.95 | 0.025 |
| 2.32635 | 0.99 | 0.98 | 0.01 |
| 2.57583 | 0.995 | 0.99 | 0.005 |
| 3.09023 | 0.999 | 0.998 | 0.001 |
| 3.29053 | 0.9995 | 0.999 | 0.0005 |
| 3.71902 | 0.9999 | 0.9998 | 0.0001 |
| 3.89059 | 0.99995 | 0.9999 | 0.00005 |
| 4.26489 | 0.99999 | 0.99998 | 0.00001 |

It is sometimes of interest to scrutinize extreme significance levels which implies integrating the far tails of a normal distribution. In the table below we give the number of standard deviations, $z$, required in order to achieve a one-tailed probability content of $10^{-n}$.

$$
z \text {-values for which } \frac{1}{\sqrt{2 \pi}} \int_{z}^{\infty} e^{-z^{2} / 2} d z=10^{-n} \text { for } n=1,2, \ldots, 23
$$

| $n$ | $z$ | $n$ | $z$ | $n$ | $z$ | $n$ | $z$ | $n$ | $z$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 1.28155 | 6 | 4.75342 | 11 | 6.70602 | 16 | 8.22208 | 21 | 9.50502 |
| 2 | 2.32635 | 7 | 5.19934 | 12 | 7.03448 | 17 | 8.49379 | 22 | 9.74179 |
| 3 | 3.09023 | 8 | 5.61200 | 13 | 7.34880 | 18 | 8.75729 | 23 | 9.97305 |
| 4 | 3.71902 | 9 | 5.99781 | 14 | 7.65063 | 19 | 9.01327 |  |  |
| 5 | 4.26489 | 10 | 6.36134 | 15 | 7.94135 | 20 | 9.26234 |  |  |

Below are also given the one-tailed probability content for a standard normal distribution in the region from $z$ to $\infty$ (or $-\infty$ to $-z$ ). The information in the previous as well as this table is taken from [26].

$$
\text { Probability content } Q(z)=\frac{1}{\sqrt{2 \pi}} \int_{z}^{\infty} e^{-z^{2} / 2} d z \text { for } z=1,2, \ldots, 50,60, \ldots, 100,150, \ldots, 500
$$

| $z$ | $-\log Q(z)$ | $z$ | $-\log Q(z)$ | $z$ | $-\log Q(z)$ | $z$ | $-\log Q(z)$ | $z$ | $-\log Q(z)$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 0.79955 | 14 | 44.10827 | 27 | 160.13139 | 40 | 349.43701 | 80 | 1392.04459 |
| 2 | 1.64302 | 15 | 50.43522 | 28 | 172.09024 | 41 | 367.03664 | 90 | 1761.24604 |
| 3 | 2.86970 | 16 | 57.19458 | 29 | 184.48283 | 42 | 385.07032 | 100 | 2173.87154 |
| 4 | 4.49934 | 17 | 64.38658 | 30 | 197.30921 | 43 | 403.53804 | 150 | 4888.38812 |
| 5 | 6.54265 | 18 | 72.01140 | 31 | 210.56940 | 44 | 422.43983 | 200 | 8688.58977 |
| 6 | 9.00586 | 19 | 80.06919 | 32 | 224.26344 | 45 | 441.77568 | 250 | 13574.49960 |
| 7 | 11.89285 | 20 | 88.56010 | 33 | 238.39135 | 46 | 461.54561 | 300 | 19546.12790 |
| 8 | 15.20614 | 21 | 97.48422 | 34 | 252.95315 | 47 | 481.74964 | 350 | 26603.48018 |
| 9 | 18.94746 | 22 | 106.84167 | 35 | 267.94888 | 48 | 502.38776 | 400 | 34746.55970 |
| 10 | 23.11805 | 23 | 116.63253 | 36 | 283.37855 | 49 | 523.45999 | 450 | 43975.36860 |
| 11 | 27.71882 | 24 | 126.85686 | 37 | 299.24218 | 50 | 544.96634 | 500 | 54289.90830 |
| 12 | 32.75044 | 25 | 137.51475 | 38 | 315.53979 | 60 | 783.90743 |  |  |
| 13 | 38.21345 | 26 | 148.60624 | 39 | 332.27139 | 70 | 1066.26576 |  |  |

Beware, however, that extreme significance levels are purely theoretical and that one seldom or never should trust experimental limits at these levels. In an experimental situations one rarely fulfills the statistical laws to such detail and any bias or background may heavily affect statements on extremely small probabilities.

Although one normally would use a routine to find the probability content for a normal distribution it is sometimes convenient to have a "classical" table available. In table 6 on page 176 we give probability contents for a symmetric region from $-z$ to $z$ for $z$-values ranging from 0.00 to 3.99 in steps of 0.01 . Conversely we give in table 7 on page 177 the $z$-values corresponding to specific probability contents from 0.000 to 0.998 in steps of 0.002 .

### 34.8 Random Number Generation

There are many different methods to obtain random numbers from a normal distribution some of which are reviewed below. It is enough to consider the case of a standard normal distribution since given such a random number $z$ we may easily obtain one from a general normal distribution by making the transformation $x=\mu+\sigma z$.

Below $f(x)$ denotes the standard normal distribution and if not explicitly stated all variables denoted by $\xi$ are uniform random numbers in the range from zero to one.

### 34.8.1 Central Limit Theory Approach

The sum of $n$ independent random numbers from a uniform distribution between zero and one, $R_{n}$, has expectation value $E\left(R_{n}\right)=n / 2$ and variance $V\left(R_{n}\right)=n / 12$. By the central limit theorem the quantity

$$
z_{n}=\frac{R_{n}-E\left(R_{n}\right)}{\sqrt{V\left(R_{n}\right)}}=\frac{R_{n}-\frac{n}{2}}{\sqrt{\frac{n}{12}}}
$$

approaches the standard normal distribution as $n \rightarrow \infty$. A practical choice is $n=12$ since this expression simplifies to $z_{12}=R_{12}-6$ which could be taken as a random number from a standard normal distribution. Note, however, that this method is neither accurate nor fast.

### 34.8.2 Exact Transformation

The Box-Muller transformation used to find random numbers from the binormal distribution (see section 6.5 on page 22), using two uniform random numbers between zero and one in $\xi_{1}$ and $\xi_{2}$,

$$
\begin{aligned}
& z_{1}=\sqrt{-2 \ln \xi_{1}} \sin 2 \pi \xi_{2} \\
& z_{2}=\sqrt{-2 \ln \xi_{1}} \cos 2 \pi \xi_{2}
\end{aligned}
$$

may be used to obtain two independent random numbers from a standard normal distribution.

### 34.8.3 Polar Method

The above method may be altered in order to avoid the cosine and sine by
i Generate $u$ and $v$ as two uniformly distributed random numbers in the range from -1 to 1 by $u=2 \xi_{1}-1$ and $v=2 \xi_{2}-1$.
ii Calculate $w=u^{2}+v^{2}$ and if $w>1$ then go back to $\mathbf{i}$.
iii Return $x=u z$ and $y=v z$ with $z=\sqrt{-2 \ln w / w}$ as two independent random numbers from a standard normal distribution.

This method is often faster than the previous since it eliminates the sine and cosine at the slight expense of $1-\pi / 4 \approx 21 \%$ rejection in step iii and a few more arithmetic operations. As is easily seen $u / \sqrt{w}$ and $v / \sqrt{w}$ plays the role of the cosine and the sine in the previous method.

### 34.8.4 Trapezoidal Method

The maximum trapezoid that may be inscribed under the standard normal curve covers an area of $91.95 \%$ of the total area. Random numbers from a trapezoid is easily obtained by a linear combination of two uniform random numbers. In the remaining cases a tail-technique and accept-reject techniques, as described in figure 26, are used.


Figure 26: Trapezoidal method
Below we describe, in some detail, a slightly modified version of what is presented in [28]. For more exact values of the constants used see this reference.
i Generate two uniform random numbers between zero and one $\xi$ and $\xi_{0}$
ii If $\xi<0.9195$ generate a random number from the trapezoid by $x=2.404 \xi_{0}+1.984 \xi-$ 2.114 and exit
iii Else if $\xi<0.9541$ ( $3.45 \%$ of all cases) generate a random number from the tail $x>2.114$
a Generate two uniform random numbers $\xi_{1}$ and $\xi_{2}$
b Put $x=2.114^{2}-2 \ln \xi_{1}$ and if $x \xi_{2}^{2}>2.114^{2}$ then go back to a
c Put $x=\sqrt{x}$ and go to vii
iv Else if $\xi<0.9782(2.41 \%$ of all cases) generate a random number from the region $0.290<x<1.840$ between the normal curve and the trapezoid
a Generate two uniform random numbers $\xi_{1}$ and $\xi_{2}$
b Put $x=0.290+1.551 \xi_{1}$ and if $f(x)-0.443+0.210 x<0.016 \xi_{2}$ then go to a
c Go to vii
$\mathbf{v}$ Else if $\xi<0.9937$ ( $1.55 \%$ of all cases) generate a random number from the region $1.840<\xi<2.114$ between the normal curve and the trapezoid
a Generate two uniform random numbers $\xi_{1}$ and $\xi_{2}$
b Put $x=1.840+0.274 \xi_{1}$ and if $f(x)-0.443+0.210 x<0.043 \xi_{2}$ then go to a
c Go to vii
vi Else, in $0.63 \%$ of all cases, generate a random number from the region $0<x<0.290$ between the normal curve and the trapezoid by
a Generate two uniform random numbers $\xi_{1}$ and $\xi_{2}$
b Put $x=0.290 \xi_{1}$ and if $f(x)-0.383<0.016 \xi_{2}$ then go back to a
vii Assign a minus sign to $x$ if $\xi_{0} \geq \frac{1}{2}$

### 34.8.5 Center-tail method

Ahrens and Dieter [28] also proposes a so called center-tail method. In their article they treat the tails outside $|z|>\sqrt{2}$ with a special tail method which avoids the logarithm. However, it turns out that using the same tail method as in the previous method is even faster. The method is as follows:
i Generate a uniform random number $\xi$ and use the first bit after the decimal point as a sign bit $s$ i.e. for $\xi \leq \frac{1}{2}$ put $\xi=2 \xi$ and $s=-1$ and for $\xi>\frac{1}{2}$ put $\xi=2 \xi-1$ and $s=1$
ii If $\xi>0.842700792949715$ (the area for $-\sqrt{2}<z<\sqrt{2}$ ) go to vi.
iii Center method: Generate $\xi_{0}$ and set $\nu=\xi+0$
iv Generate $\xi_{1}$ and $\xi_{2}$ and set $\nu^{*}=\max \left(\xi_{1}, \xi_{2}\right)$.
If $\nu<\nu^{*}$ calculate $y=\xi_{0} \sqrt{2}$ and go to viii
$\mathbf{v}$ Generate $\xi_{1}$ and $\xi_{2}$ and set $\nu=\max \left(\xi_{1}, \xi_{2}\right)$
If $\nu<\nu^{*}$ go to iv else go to iii
vi Tail method: Generate $\xi_{1}$ and set $y=1-\ln \xi_{1}$
vii Generate $\xi_{2}$ and if $y \xi_{2}^{2}>1$ go to vi else put $y=\sqrt{y}$
viii Set $x=s y \sqrt{2}$.

### 34.8.6 Composition-rejection Methods

In reference [21] two methods using the composition-rejection method is proposed. The first one, attributed to Butcher [23] and Kahn, uses only one term in the sum and has $\alpha=\sqrt{2 e / \pi}, f(x)=\exp \{-x\}$ and $g(x)=\exp \left\{-(x-1)^{2} / 2\right\}$. The algorithm is as follows:
i Generate $\xi_{1}$ and $\xi_{2}$
ii Determine $x=-\ln \xi_{1}$, i.e. a random number from $f(x)$
iii Determine $g(x)=\exp \left\{-(x-1)^{2} / 2\right\}$
iv If $\xi_{2}>g(x)$ then go to $\mathbf{i}$
$\mathbf{v}$ Decide a sign either by generating a new random number, or by using $\xi_{2}$ for which $0<\xi_{2} \leq g(x)$ here, and exit with $x$ with this sign.

The second method is originally proposed by J. C. Butcher [23] and uses two terms

$$
\begin{array}{lll}
\alpha_{1}=\sqrt{\frac{2}{\pi}} & f_{1}(x)=1 & g_{1}(x)=e^{-\frac{x^{2}}{2}} \\
\alpha_{2}=1 / \sqrt{2 \pi} & f_{2}(x)=2 e^{-2(x-1)} & g_{2}(x)=e^{-\frac{(x-2)^{2}}{2}} \\
\text { for } x>1
\end{array}
$$

i Generate $\xi_{1}$ and $\xi_{2}$
ii If $\xi_{1}-\frac{2}{3}>0$ then determine $x=1-\frac{1}{2} \ln \left(3 \xi_{1}-2\right)$ and $z=\frac{1}{2}(x-2)^{2}$ else determine $x=3 \xi_{1} / 2$ and $z=x^{2} / 2$
iii Determine $g=e^{-z}$
iv If $\xi_{2}>g$ the go to $\mathbf{i}$
$\mathbf{v}$ Determine the sign of $\xi_{2}-g / 2$ and exit with $x$ with this sign.

### 34.8.7 Method by Marsaglia

A nice method proposed by G. Marsaglia is based on inscribing a spline function beneath the standard normal curve and subsequently a triangular distribution beneath the remaining difference. See figure 27 for a graphical presentation of the method. The algorithm used is described below.


Figure 27: Marsaglia method

- The sum of three uniform random numbers $\xi_{1}, \xi_{2}$, and $\xi_{3}$ follow a parabolic spline function. Using $x=2\left(\xi_{1}+\xi_{2}+\xi_{3}-\frac{3}{2}\right)$ we obtain a distribution

$$
f_{1}(x)=\left\{\begin{array}{lll}
\left(3-x^{2}\right) / 8 & \text { if } & |x| \leq 1 \\
(3-|x|)^{2} / 16 & \text { if } & 1<|x| \leq 3 \\
0 & \text { if } & |x|>3
\end{array}\right.
$$

Maximizing $\alpha_{1}$ with the constraint $f(x)-\alpha_{1} f_{1}(x) \geq 0$ in the full interval $|x| \leq 3$ gives $\alpha_{1}=16 e^{-2} / \sqrt{2 \pi} \approx 0.8638554$ i.e. in about $86 \%$ of all cases such a combination is made.

- Moreover, a triangular distribution given by making the combination $x=\frac{3}{2}\left(\xi_{1}+\xi_{2}-1\right)$ leading to a function

$$
f_{2}(x)=\frac{4}{9}\left(\frac{3}{2}-|x|\right) \quad \text { for } \quad|x|<\frac{3}{2}
$$

and zero elsewhere. This function may be inscribed under the remaining curve $f(x)-f_{1}(x)$ maximizing $\alpha_{2}$ such that $f_{3}(x)=f(x)-\alpha_{1} f_{1}(x)-\alpha_{2} f_{2}(x) \geq 0$ in the interval $|x| \leq \frac{3}{2}$. This leads to a value $\alpha_{2} \approx 0.1108$ i.e. in about $11 \%$ of all cases this combination is used

- The maximum value of $f_{3}(x)$ in the region $|x| \leq 3$ is 0.0081 and here we use a straightforward reject-accept technique. This is done in about $2.26 \%$ of all cases.
- Finally, the tails outside $|x|>3$, covering about $0.27 \%$ of the total area is dealt with with a standard tail-method where
a Put $x=9-2 \ln \xi_{1}$
b If $x \xi_{2}^{2}>9$ then go to a
c Else generate a $\operatorname{sign} s=+1$ or $s=-1$ with equal probability and exit with $x=s \sqrt{x}$


### 34.8.8 Histogram Technique

Yet another method due to G. Marsaglia and collaborators [37] is one where a histogram with $k$ bins and bin-width $c$ is inscribed under the (folded) normal curve. The difference between the normal curve and the histogram is treated with a combination of triangular distributions and accept-reject techniques as well as the usual technique for the tails. Trying to optimize fast generation we found $k=9$ and $c=\frac{1}{3}$ to be a fair choice. This may, however, not be true on all computers. See figure 28 for a graphical presentation of the method.


Figure 28: Histogram method
The algorithm is as follows:
$\mathbf{i}$ Generate $\xi_{1}$ and chose which region to generate from. This is done e.g. with a sequential search in a cumulative vector where the areas of the regions have been
sorted in descending order. The number of elements in this vector is $2 k+\frac{1}{c}+1$ which for the parameters mentioned above becomes 22 .
ii If a histogram bin $i(i=1,2, \ldots, k)$ is selected then determine $x=\left(\xi_{2}+i-1\right) c$ and go to vii.
iii If an inscribed triangle $i\left(i=1,2, \ldots, \frac{1}{c}\right)$ then determine $x=\left(\min \left(\xi_{2}, \xi_{3}\right)+1-i\right) c$ and go to vii.
iv If subscribed triangle $i\left(i=\frac{1}{c}+1, \ldots, k\right)$ then determine $x=\left(\min \left(\xi_{2}, \xi_{3}\right)+i-1\right) c$ and accept this value with a probability equal to the ratio between the normal curve and the triangle at this $x$-value (histogram subtracted in both cases) else iterate. When a value is accepted then go to vii.
$\mathbf{v}$ For the remaining $\frac{1}{c}$ regions between the inscribed triangles and the normal curve for $x<1$ use a standard reject accept method in each bin and then go to vii.
vi If the tail region is selected then use a standard technique e.g. (a) $x=(k c)^{2}-2 \ln \xi_{2}$, (b) if $x \xi_{3}^{2}>(k c)^{2}$ then go to a else use $x=\sqrt{x}$.
vii Attach a random sign to $x$ and exit. This is done by either generating a new uniform random number or by saving the first bit of $\xi_{1}$ in step $\mathbf{i}$. The latter is faster and the degradation in precision is negligible.

### 34.8.9 Ratio of Uniform Deviates

A technique using the ratio of two uniform deviates was propose by A. J. Kinderman and J. F. Monahan in 1977 [38]. It is based on selecting an acceptance region such that the ratio of two uniform pseudorandom numbers follow the standard normal distribution. With $u$ and $v$ uniform random numbers, $u$ between 0 and 1 and $v$ between $-\sqrt{2 / e}$ and $\sqrt{2 / e}$, such a region is defined by

$$
v^{2}<-4 u^{2} \ln u
$$

as is shown in the left-hand side of figure 29 .
Note that it is enough to consider the upper part $(v>0)$ of the acceptance limit due to the symmetry of the problem. In order to avoid taking the logarithm, which may slow the algorithm down, simpler boundary curves were designed. An improvement to the original proposal was made by Joseph L. Leva in 1992 [39,40] choosing the same quadratic form for both the lower and the upper boundary namely

$$
Q(u, v)=(u-s)^{2}-b(u-s)(v-t)+(a-v)^{2}
$$

Here $(s, t)=(0.449871,-0.386595)$ is the center of the ellipses and $a=0.196$ and $b=0.25472$ are suitable constants to obtain tight boundaries. In the right-hand side of figure 29 we show the value of the quadratic form at the acceptance limit $Q(u, 2 u \sqrt{-\ln u})$ as a function of $u$. It may be deduced that only in the interval $r_{1}<Q<r_{2}$ with $r_{1}=0.27597$ and $r_{2}=0.27846$ we still have to evaluate the logarithm.

The algorithm is as follows:


Figure 29: Method using ratio between two uniform deviates
i Generate uniform random numbers $u=\xi_{1}$ and $v=2 \sqrt{2 / e}\left(\xi_{2}-\frac{1}{2}\right)$.
ii Evaluate the quadratic form $Q=x^{2}+y(a y-b x)$ with $x=u-s$ and $y=|v|-t$.
iii Accept if inside inner boundary, i.e. if $Q<r_{1}$, then go to vi.
iv Reject if outside upper boundary, i.e. if $Q>r_{2}$, then go to $\mathbf{i}$.
$\mathbf{v}$ Reject if outside acceptance region, i.e. if $v^{2}>-4 u^{2} \ln u$, then go to $\mathbf{i}$.
vi Return the ratio $v / u$ as a pseudorandom number from a standard normal distribution.
On average 2.738 uniform random numbers are consumed and 0.012 logarithms are computed per each standard normal random number obtained by this algorithm. As a comparison the number of logarithmic evaluations without cutting on the boundaries, skipping steps ii through iv above, would be 1.369. The penalty when using logarithms on modern computers is not as severe as it used to be but still some efficiency is gained by using the proposed algorithm.

### 34.8.10 Comparison of random number generators

Above we described several methods to achieve pseudorandom numbers from a standard normal distribution. Which one is the most efficient may vary depending on the actual implementation and the computer it is used at. To give a rough idea we found the following times per random number ${ }^{8}$ (in the table are also given the average number of uniform

[^6]pseudorandom numbers consumed per random number in our implementations)

| Method | section | $\mu \mathrm{s} /$ r.n. | $N_{\xi} /$ r.n. | comment |
| :--- | :---: | :---: | :---: | :--- |
| Trapezoidal method | 34.8 .4 | 0.39 | 2.246 |  |
| Polar method | 34.8 .3 | 0.41 | 1.273 | pair |
| Histogram method | 34.8 .8 | 0.42 | 2.121 |  |
| Box-Muller transformation | 34.8 .2 | 0.44 | 1.000 | pair |
| Spline functions | 34.8 .7 | 0.46 | 3.055 |  |
| Ratio of two uniform deviates | 34.8 .9 | 0.55 | 2.738 |  |
| Composition-rejection, two terms | 34.8 .6 | 0.68 | 2.394 |  |
| Center-tail method | 34.8 .5 | 0.88 | 5.844 |  |
| Composition-rejection, one term | 34.8 .6 | 0.90 | 2.631 |  |
| Central limit method approach | 34.8 .1 | 1.16 | 12.000 | inaccurate |

The trapezoidal method is thus fastest but the difference is not great as compared to some of the others. The central limit theorem method is slow as well as inaccurate although it might be the easiest to remember. The other methods are all exact except for possible numerical problems. "Pair" indicates that these generators give two random numbers at a time which may implies that either one is not used or one is left pending for the next call (as is the case in our implementations).

### 34.9 Tests on Parameters of a Normal Distribution

For observations from a normal sample different statistical distributions are applicable in different situations when it comes to estimating one or both of the parameters $\mu$ and $\sigma$. In the table below we try to summarize this in a condensed form.

| TESTS OF MEAN AND VARIANCE OF NORMAL DISTRIBUTION |  |  |  |
| :---: | :---: | :---: | :---: |
| $H_{0}$ | Condition | Statistic | Distribution |
| $\mu=\mu_{0}$ | $\sigma^{2}$ known | $\frac{\bar{x}-\mu_{0}}{\sigma / \sqrt{n}}$ | $N(0,1)$ |
|  | $\sigma^{2}$ unknown | $\frac{\bar{x}-\mu_{0}}{s / \sqrt{n}}$ | $t_{n-1}$ |
| $\sigma^{2}=\sigma_{0}^{2}$ | $\mu$ known | $\frac{(n-1) s^{2}}{\sigma_{0}^{2}}=\sum_{i=1}^{n} \frac{\left(x_{i}-\mu\right)^{2}}{\sigma_{0}^{2}}$ | $\chi_{n}^{2}$ |
|  | $\mu$ unknown | $\frac{(n-1) s^{2}}{\sigma_{0}^{2}}=\sum_{i=1}^{n} \frac{\left(x_{i}-\bar{x}\right)^{2}}{\sigma_{0}^{2}}$ | $\chi_{n-1}^{2}$ |
| $\mu_{1}=\mu_{2}=\mu$ | $\sigma_{1}^{2}=\sigma_{2}^{2}=\sigma^{2}$ known | $\frac{\bar{x}-\bar{y}}{\sigma \sqrt{\frac{1}{n}+\frac{1}{m}}}$ | $N(0,1)$ |
|  | $\sigma_{1}^{2} \neq \sigma_{2}^{2}$ known | $\frac{\bar{x}-\bar{y}}{\sqrt{\frac{\sigma_{1}^{2}}{}+\frac{\sigma_{2}^{2}}{}}}$ | $N(0,1)$ |
|  | $\sigma_{1}^{2}=\sigma_{2}^{2}=\sigma^{2}$ unknown | $s=\frac{\frac{\sqrt{\frac{1}{n}}+\frac{2}{m}}{s \sqrt{\frac{1}{n}+\frac{1}{m}}}}{\frac{(n-1) s_{1}^{2}+(m-1) s_{2}^{2}}{n+m-2}}$ | $t_{n+m-2}$ |
| $\sigma_{1}^{2}=\sigma_{2}^{2}$ | $\sigma_{1}^{2} \neq \sigma_{2}^{2}$ unknown | $\frac{\bar{x}-\bar{y}}{\sqrt{\frac{s_{1}^{2}}{n}+\frac{s_{2}^{2}}{m}}}$ | $\approx N(0,1)$ |
|  | $\mu_{1} \neq \mu_{2}$ known | $\frac{s_{1}^{2}}{s_{2}^{2}}=\frac{\frac{1}{n-1} \sum_{i=1}^{n}\left(x_{i}-\mu_{1}\right)^{2}}{\frac{1}{m-1} \sum_{i=1}^{m}\left(y_{i}-\mu_{2}\right)^{2}}$ | $F_{n, m}$ |
|  | $\mu_{1} \neq \mu_{2}$ unknown | $\frac{s_{1}^{2}}{s_{2}^{2}}=\frac{\frac{1}{n-1} \sum_{i=1}^{n}\left(x_{i}-\bar{x}\right)^{2}}{\frac{1}{m-1} \sum_{j=1}^{m}\left(y_{i}-\bar{y}\right)^{2}}$ | $F_{n-1, m-1}$ |

## 35 Pareto Distribution

### 35.1 Introduction

The Pareto distribution is given by

$$
f(x ; \alpha, k)=\alpha k^{\alpha} / x^{\alpha+1}
$$

where the variable $x \geq k$ and the parameter $\alpha>0$ are real numbers. As is seen $k$ is only a scale factor.

The distribution has its name after its inventor the italian Vilfredo Pareto (1848-1923) who worked in the fields of national economy and sociology (professor in Lausanne, Switzerland). It was introduced in order to explain the distribution of wages in society.

### 35.2 Cumulative Distribution

The cumulative distribution is given by

$$
F(x)=\int_{k}^{x} f(u) d u=1-\left(\frac{k}{x}\right)^{\alpha}
$$

### 35.3 Moments

Algebraic moments are given by

$$
E\left(x^{n}\right)=\int_{k}^{\infty} x^{n} f(x)=\int_{k}^{\infty} x^{n} \frac{\alpha k^{\alpha}}{x^{\alpha+1}}=\left[-\frac{\alpha k^{\alpha}}{x^{\alpha-n+1}}\right]_{k}^{\infty}=\frac{\alpha k^{\alpha}}{\alpha-n}
$$

which is defined for $\alpha>n$.
Especially the expectation value and variance are given by

$$
\begin{aligned}
E(x) & =\frac{\alpha k}{\alpha-1} \text { for } \alpha>1 \\
V(x) & =\frac{\alpha k^{2}}{(\alpha-2)(\alpha-1)^{2}} \text { for } \alpha>2
\end{aligned}
$$

### 35.4 Random Numbers

To obtain a random number from a Pareto distribution we use the straightforward way of solving the equation $F(x)=\xi$ with $\xi$ a random number uniformly distributed between zero and one. This gives

$$
F(x)=1-\left(\frac{k}{x}\right)^{\alpha}=\xi \quad \Rightarrow \quad x=\frac{k}{(1-\xi)^{\frac{1}{\alpha}}}
$$

## 36 Poisson Distribution

### 36.1 Introduction

The Poisson distribution is given by

$$
p(r ; \mu)=\frac{\mu^{r} e^{-\mu}}{r!}
$$

where the variable $r$ is an integer $(r \geq 0)$ and the parameter $\mu$ is a real positive quantity. It is named after the french mathematician Siméon Denis Poisson (1781-1840) who was the first to present this distribution in 1837 (implicitly the distribution was known already in the beginning of the 18th century).

As is easily seen by comparing two subsequent $r$-values the distribution increases up to $r+1<\mu$ and then declines to zero. For low values of $\mu$ it is very skewed (for $\mu<1$ it is J-shaped).

The Poisson distribution describes the probability to find exactly $r$ events in a given length of time if the events occur independently at a constant rate $\mu$. An unbiased and efficient estimator of the Poisson parameter $\mu$ for a sample with $n$ observations $x_{i}$ is $\hat{\mu}=\bar{x}$, the sample mean, with variance $V(\hat{\mu})=\mu / n$.

For $\mu \rightarrow \infty$ the distribution tends to a normal distribution with mean $\mu$ and variance $\mu$.

The Poisson distribution is one of the most important distributions in statistics with many applications. Along with the properties of the distribution we give a few examples here but for a more thorough description we refer to standard text-books.

### 36.2 Moments

The expectation value, variance, third and fourth central moments of the Poisson distribution are

$$
\begin{aligned}
E(r) & =\mu \\
V(r) & =\mu \\
\mu_{3} & =\mu \\
\mu_{4} & =\mu(1+3 \mu)
\end{aligned}
$$

The coefficients of skewness and kurtosis are $\gamma_{1}=1 / \sqrt{\mu}$ and $\gamma_{2}=1 / \mu$ respectively, i.e. they tend to zero as $\mu \rightarrow \infty$ in accordance with the distribution becoming approximately normally distributed for large values of $\mu$.

Algebraic moments may be found by the recursive formula

$$
\mu_{k+1}^{\prime}=\mu\left\{\mu_{k}^{\prime}+\frac{d \mu_{k}^{\prime}}{d \mu}\right\}
$$

and central moments by a similar formula

$$
\mu_{k+1}=\mu\left\{k \mu_{k-1}+\frac{d \mu_{k}}{d \mu}\right\}
$$

For a Poisson distribution one may note that factorial moments $g_{k}$ (cf page 6) and cumulants $\kappa_{k}$ (see section 2.5) become especially simple

$$
\begin{aligned}
& g_{k}=E(r(r-1) \cdots(r-k+1))=\mu^{k} \\
& \kappa_{r}=\mu \text { for all } r \geq 1
\end{aligned}
$$

### 36.3 Probability Generating Function

The probability generating function is given by

$$
G(z)=E\left(z^{r}\right)=\sum_{r=0}^{\infty} z^{r} \frac{\mu^{r} e^{-\mu}}{r!}=e^{-\mu} \sum_{r=0}^{\infty} \frac{(\mu z)}{r!}=e^{\mu(z-1)}
$$

Although we mostly use the probability generating function in the case of a discrete distribution we may also define the characteristic function

$$
\phi(t)=E\left(e^{\imath t r}\right)=e^{-\mu} \sum_{r=0}^{\infty} e^{\imath t r} \frac{\mu^{r}}{r!}=\exp \left\{\mu\left(e^{\imath t}-1\right)\right\}
$$

a result which could have been given directly since $\phi(t)=G\left(e^{\imath t}\right)$.

### 36.4 Cumulative Distribution

When calculating the probability content of a Poisson distribution we need the cumulative, or distribution, function. This is easily obtained by finding the individual probabilities e.g. by the recursive formula $p(r)=p(r-1) \frac{\mu}{r}$ starting with $p(0)=e^{-\mu}$.

There is, however, also an interesting connection to the incomplete Gamma function [10]

$$
P(r)=\sum_{k=0}^{r} \frac{\mu^{k} e^{-\mu}}{k!}=1-P(r+1, \mu)
$$

with $P(a, x)$ the incomplete Gamma function not to be confused with $P(r)$.
Since the cumulative chi-square distribution also has a relation to the incomplete Gamma function one may obtain a relation between these cumulative distributions namely

$$
P(r)=\sum_{k=0}^{r} \frac{\mu^{k} e^{-\mu}}{k!}=1-\int_{0}^{2 \mu} f(x ; \nu=2 r+2) d x
$$

where $f(x ; \nu=2 r+2)$ denotes the chi-square distribution with $\nu$ degrees of freedom.

### 36.5 Addition Theorem

The so called addition theorem states that the sum of any number of independent Poissondistributed variables is also distributed according to a Poisson distribution.

For $n$ variables each distributed according to the Poisson distribution with parameters (means) $\mu_{i}$ we find characteristic function

$$
\phi\left(r_{1}+r_{2}+\ldots+r_{n}\right)=\prod_{i=1}^{n} \exp \left\{\mu_{i}\left(e^{\imath t}-1\right)\right\}=\exp \left\{\sum_{i=1}^{n} \mu_{i}\left(e^{\imath t}-1\right)\right\}
$$

which is the characteristic function for a Poisson variable with parameter $\mu=\sum \mu_{i}$.

### 36.6 Derivation of the Poisson Distribution

For a binomial distribution the rate of "success" $p$ may be very small but in a long series of trials the total number of successes may still be a considerable number. In the limit $p \rightarrow 0$ and $N \rightarrow \infty$ but with $N p=\mu$ a finite constant we find

$$
\begin{aligned}
p(r) & =\binom{N}{r} p^{r}(1-p)^{N-r} \approx \frac{1}{r!} \frac{\sqrt{2 \pi N} N^{N} e^{-N}}{\sqrt{2 \pi(N-r)}(N-r)^{N-r} e^{-(N-r)}}\left(\frac{\mu}{N}\right)^{r}\left(1-\frac{\mu}{N}\right)^{N-r}= \\
& =\frac{1}{r!} \sqrt{\frac{N}{N-r}} \frac{1}{\left(1-\frac{r}{N}\right)^{N}} e^{-r} \mu^{r}\left(1-\frac{\mu}{N}\right)^{N-r} \rightarrow \frac{\mu^{r} e^{-\mu}}{r!}
\end{aligned}
$$

as $N \rightarrow \infty$ and where we have used that $\lim _{n \rightarrow \infty}\left(1-\frac{x}{n}\right)^{n}=e^{-x}$ and Stirling's formula (se section 42.2) for the factorial of a large number $n!\approx \sqrt{2 \pi n} n^{n} e^{-n}$.

It was this approximation to the binomial distribution which S. D. Poisson presented in his book in 1837.

### 36.7 Histogram

In a histogram of events we would regard the distribution of the bin contents as multinomially distributed if the total number of events $N$ were regarded as a fixed number. If, however, we would regard the total number of events not as fixed but as distributed according to a Poisson distribution with mean $\nu$ we obtain (with $k$ bins in the histogram and the multinomial probabilities for each bin in the vector $p$ )

Given a multinomial distribution, denoted $M(\underline{r} ; N, \underline{p})$, for the distribution of events into bins for fixed $N$ and a Poisson distribution, denoted $P(N ; \nu)$, for the distribution of $N$ we write the joint distribution

$$
\begin{aligned}
\mathcal{P}(\underline{r}, N) & =M(\underline{r} ; N, \underline{p}) P(N ; \nu)=\left(\frac{N!}{r_{1}!r_{2}!\ldots r_{k}!} p_{1}^{r_{1}} p_{2}^{r_{2}} \ldots p_{k}^{r_{k}}\right)\left(\frac{\nu^{N} e^{-\nu}}{N!}\right)= \\
& =\left(\frac{1}{r_{1}!}\left(\nu p_{1}\right)^{r_{1}} e^{-\nu p_{1}}\right)\left(\frac{1}{r_{2}!}\left(\nu p_{2}\right)^{r_{2}} e^{-\nu p_{2}}\right) \ldots\left(\frac{1}{r_{k}!}\left(\nu p_{k}\right)^{r_{k}} e^{-\nu p_{k}}\right)
\end{aligned}
$$

where we have used that

$$
\sum_{i=1}^{k} p_{i}=1 \quad \text { and } \quad \sum_{i=1}^{k} r_{i}=N
$$

i.e. we get a product of independent Poisson distributions with means $\nu p_{i}$ for each individual bin. A simpler case leading to the same result would be the classification into only two groups using a binomial and a Poisson distribution.

The assumption of independent Poisson distributions for the number events in each bin is behind the usual rule of using $\sqrt{N}$ as the standard deviation in a bin with $N$ entries and neglecting correlations between bins in a histogram.

### 36.8 Random Number Generation

By use of the cumulative technique e.g. forming the cumulative distribution by starting with $P(0)=e^{-\mu}$ and using the recursive formula

$$
P(r)=P(r-1) \frac{\mu}{r}
$$

a random number from a Poisson distribution is easily obtained using one uniform random number between zero and one. If $\mu$ is a constant the by far fastest generation is obtained if the cumulative vector is prepared once for all.

An alternative is to obtain, in $\rho$, a random number from a Poisson distribution by multiplying independent uniform random numbers $\xi_{i}$ until

$$
\prod_{i=0}^{\rho} \xi_{i} \leq e^{-\mu}
$$

For large values of $\mu$ use the normal approximation but beware of the fact that the Poisson distribution is a function in a discrete variable.

## 37 Rayleigh Distribution

### 37.1 Introduction

The Rayleigh distribution is given by

$$
f(x ; \alpha)=\frac{x}{\alpha^{2}} e^{-\frac{x^{2}}{2 \alpha^{2}}}
$$

for real positive values of the variable $x$ and a real positive parameter $\alpha$. It is named after the british physicist Lord Rayleigh (1842-1919), also known as Baron John William Strutt Rayleigh of Terling Place and Nobel prize winner in physics 1904.

Note that the parameter $\alpha$ is simply a scale factor and that the variable $y=x / \alpha$ has the simplified distribution $g(y)=y e^{-y^{2} / 2}$.


Figure 30: The Rayleigh distribution
The distribution, shown in figure 30 , has a mode at $x=\alpha$ and is positively skewed.

### 37.2 Moments

Algebraic moments are given by

$$
E\left(x^{n}\right)=\int_{0}^{\infty} x^{n} f(x) d x=\frac{1}{2 \alpha^{2}} \int_{-\infty}^{\infty}|x|^{n+1} e^{-x^{2} / 2 \alpha^{2}}
$$

i.e. we have a connection to the absolute moments of the Gauss distribution. Using these (see section 34 on the normal distribution) the result is

$$
E\left(x^{n}\right)= \begin{cases}\sqrt{\frac{\pi}{2}} n!!\alpha^{n} & \text { for } n \text { odd } \\ 2^{k} k!\alpha^{2 k} & \text { for } n=2 k\end{cases}
$$

Specifically we note that the expectation value, variance, and the third and fourth central moments are given by

$$
E(x)=\alpha \sqrt{\frac{\pi}{2}}, \quad V(x)=\alpha^{2}\left(2-\frac{\pi}{2}\right), \quad \mu_{3}=\alpha^{3}(\pi-3) \sqrt{\frac{\pi}{2}}, \quad \text { and } \mu_{4}=\alpha^{4}\left(8-\frac{3 \pi^{2}}{4}\right)
$$

The coefficients of skewness and kurtosis is thus

$$
\gamma_{1}=\frac{(\pi-3) \sqrt{\frac{\pi}{2}}}{\left(2-\frac{\pi}{2}\right)^{\frac{3}{2}}} \approx 0.63111 \quad \text { and } \quad \gamma_{2}=\frac{8-\frac{3 \pi^{2}}{4}}{\left(2-\frac{\pi}{2}\right)^{2}}-3 \approx 0.24509
$$

### 37.3 Cumulative Distribution

The cumulative distribution, or the distribution function, is given by

$$
F(x)=\int_{0}^{x} f(y) d y=\frac{1}{a^{2}} \int_{0}^{x} y e^{-\frac{y^{2}}{2 \alpha^{2}}} d y=\int_{0}^{\frac{x^{2}}{2 \alpha^{2}}} e^{-z} d z=1-e^{-\frac{x^{2}}{2 \alpha^{2}}}
$$

where we have made the substitution $z=\frac{y^{2}}{2 \alpha^{2}}$ in order to simplify the integration. As it should we see that $F(0)=0$ and $F(\infty)=1$.

Using this we may estimate the median $\mathcal{M}$ by

$$
F(\mathcal{M})=\frac{1}{2} \quad \Rightarrow \mathcal{M}=\alpha \sqrt{2 \ln 2} \approx 1.17741 \alpha
$$

and the lower and upper quartiles becomes

$$
\mathcal{Q}_{1}=\alpha \sqrt{-2 \ln \frac{3}{4}} \approx 0.75853 \alpha \quad \text { and } \quad \mathcal{Q}_{3}=\alpha \sqrt{2 \ln 4} \approx 1.66511 \alpha
$$

and the same technique is useful when generating random numbers from the Rayleigh distribution as is described below.

### 37.4 Two-dimensional Kinetic Theory

Given two independent coordinates $x$ and $y$ from normal distributions with zero mean and the same variance $\sigma^{2}$ the distance $z=\sqrt{x^{2}+y^{2}}$ is distributed according to the Rayleigh distribution. The $x$ and $y$ may e.g. be regarded as the velocity components of a particle moving in a plane.

To realize this we first write

$$
w=\frac{z^{2}}{\sigma^{2}}=\frac{x^{2}}{\sigma^{2}}+\frac{y^{2}}{\sigma^{2}}
$$

Since $x / \sigma$ and $y / \sigma$ are distributed as standard normal variables the sum of their squares has the chi-squared distribution with 2 degrees of freedom i.e. $g(w)=e^{-w / 2} / 2$ from which we find

$$
f(z)=g(w)\left|\frac{d w}{d z}\right|=g\left(\frac{z^{2}}{\sigma^{2}}\right) \frac{2 z}{\sigma^{2}}=\frac{z}{\sigma^{2}} e^{-\frac{z^{2}}{2 \sigma^{2}}}
$$

which we recognize as the Rayleigh distribution. This may be compared to the threedimensional case where we end up with the Maxwell distribution.

### 37.5 Random Number Generation

To obtain random numbers from the Rayleigh distribution in an efficient way we make the transformation $y=x^{2} / 2 \alpha^{2}$ a variable which follow the exponential distribution $g(y)=e^{-y}$. A random number from this distribution is easily obtained by taking minus the natural logarithm of a uniform random number. We may thus find a random number $r$ from a Rayleigh distribution by the expression

$$
r=\alpha \sqrt{-2 \ln \xi}
$$

where $\xi$ is a random number uniformly distributed between zero and one.
This could have been found at once using the cumulative distribution putting

$$
F(x)=\xi \quad \Rightarrow \quad 1-e^{-\frac{x^{2}}{2 \alpha^{2}}}=\xi \quad \Rightarrow \quad x=\alpha \sqrt{-2 \ln (1-\xi)}
$$

a result which is identical since if $\xi$ is uniformly distributed between zero and one so is $1-\xi$.

Following the examples given above we may also have used two independent random numbers from a standard normal distribution, $z_{1}$ and $z_{2}$, and construct

$$
r=\frac{1}{\alpha} \sqrt{z_{1}^{2}+z_{2}^{2}}
$$

However, this technique is not as efficient as the one outlined above.

## 38 Student's $t$-distribution

### 38.1 Introduction

The Student's $t$-distribution is given by

$$
f(t ; n)=\frac{\Gamma\left(\frac{n+1}{2}\right)}{\sqrt{n \pi} \Gamma\left(\frac{n}{2}\right)}\left(1+\frac{t^{2}}{n}\right)^{-\frac{n+1}{2}}=\frac{\left(1+\frac{t^{2}}{n}\right)^{-\frac{n+1}{2}}}{\sqrt{n} B\left(\frac{1}{2}, \frac{n}{2}\right)}
$$

where the parameter $n$ is a positive integer and the variable $t$ is a real number. The functions $\Gamma$ and $B$ are the usual Gamma and Beta functions. In figure 31 we show the $t$-distribution for $n$ values of 1 (lowest maxima), 2,5 and $\infty$ (fully drawn and identical to the standard normal distribution).


Figure 31: Graph of $t$-distribution for some values of $n$
If we change variable to $x=t / \sqrt{n}$ and put $m=\frac{n+1}{2}$ the Student's $t$-distribution becomes

$$
f(x ; m)=\frac{k}{\left(1+x^{2}\right)^{m}} \quad \text { with } \quad k=\frac{\Gamma(m)}{\Gamma\left(\frac{1}{2}\right) \Gamma\left(m-\frac{1}{2}\right)}=\frac{1}{B\left(\frac{1}{2}, m-\frac{1}{2}\right)}
$$

where $k$ is simply a normalization constant and $m$ is a positive half-integer.

### 38.2 History

A brief history behind this distribution and its name is the following. William Sealy Gosset (1876-1937) had a degree in mathematics and chemistry from Oxford when he in 1899 began working for Messrs. Guinness brewery in Dublin. In his work at the brewery he developed a small-sample theory of statistics which he needed in making small-scale experiments.

Due to company policy it was forbidden for employees to publish scientific papers and his work on the $t$-ratio was published under the pseudonym "Student". It is a very important contribution to statistical theory.

### 38.3 Moments

The Student's $t$-distribution is symmetrical around $t=0$ and thus all odd central moments vanish. In calculating even moments (note that algebraic and central moments are equal) we make use of the somewhat simpler $f(x ; m)$ form given above with $x=\frac{t}{\sqrt{n}}$ which implies the following relation between expectation values $E\left(t^{2 r}\right)=n^{r} E\left(x^{2 r}\right)$. Central moments of even order are given by, with $r$ an integer $\geq 0$,

$$
\mu_{2 r}(x)=\int_{-\infty}^{\infty} f(x ; m) d x=k \int_{-\infty}^{\infty} \frac{x^{2 r}}{\left(1+x^{2}\right)^{m}} d x=2 k \int_{0}^{\infty} \frac{x^{2 r}}{\left(1+x^{2}\right)^{m}} d x
$$

If we make the substitution $y=\frac{x^{2}}{1+x^{2}}$ implying $\frac{1}{1+x^{2}}=1-y$ and $x=\sqrt{\frac{y}{1-y}}$ then $d y=$ $\frac{2 x}{\left(1+x^{2}\right)^{2}} d x$ and we obtain

$$
\begin{aligned}
\mu_{2 r}(x) & =2 k \int_{0}^{1} \frac{x^{2 r}}{\left(1+x^{2}\right)^{m}} \cdot \frac{\left(1+x^{2}\right)^{2}}{2 x} d y=k \int_{0}^{1} \frac{x^{2 r-1}}{\left(1+x^{2}\right)^{m-2}} d y= \\
& =k \int_{0}^{1}(1-y)^{m-2}\left(\sqrt{\frac{y}{1-y}}\right)^{2 r-1} d y=k \int_{0}^{1}(1-y)^{m-r-\frac{3}{2}} y^{r-\frac{1}{2}} d y= \\
& =k B\left(r+\frac{1}{2}, m-r-\frac{1}{2}\right)=\frac{B\left(r+\frac{1}{2}, m-r-\frac{1}{2}\right)}{B\left(\frac{1}{2}, m-\frac{1}{2}\right)}
\end{aligned}
$$

The normalization constant $k$ was given above and we may now verify this expression by looking at $\mu_{0}=1$ giving $k=1 / B\left(\frac{1}{2}, m-\frac{1}{2}\right)$ and thus finally, including the $n^{r}$ factor giving moments in $t$ we have

$$
\mu_{2 r}(t)=n^{r} \mu_{2 r}(x)=n^{r} \frac{B\left(r+\frac{1}{2}, m-r-\frac{1}{2}\right)}{B\left(\frac{1}{2}, m-\frac{1}{2}\right)}=n^{r} \frac{B\left(r+\frac{1}{2}, \frac{n}{2}-r\right)}{B\left(\frac{1}{2}, \frac{n}{2}\right)}
$$

As can be seen from this expression we get into problems for $2 r \geq n$ and indeed those moments are undefined or divergent ${ }^{9}$. The formula is thus valid only for $2 r<n$. A recursive formula to obtain even algebraic moments of the $t$-distribution is

$$
\mu_{2 r}^{\prime}=\mu_{2 r-2}^{\prime} \cdot n \cdot \frac{r-\frac{1}{2}}{\frac{n}{2}-r}
$$

starting with $\mu_{0}^{\prime}=1$.
Especially we note that, when $n$ is big enough so that these moments are defined, the second central moment (i.e. the variance) is $\mu_{2}=V(t)=\frac{n}{n-2}$ and the fourth central moment is given by $\mu_{4}=\frac{3 n^{2}}{(n-2)(n-4)}$. The coefficients of skewness and kurtosis are given by $\gamma_{1}=0$ and $\gamma_{2}=\frac{6}{n-4}$, respectively.

[^7]
### 38.4 Cumulative Function

In calculating the cumulative function for the $t$-distribution it turns out to be simplifying to first estimate the integral for a symmetric region

$$
\begin{aligned}
\int_{-t}^{t} f(u) d u & =\frac{1}{\sqrt{n} B\left(\frac{1}{2}, \frac{n}{2}\right)} \int_{-t}^{t}\left(1+\frac{u^{2}}{n}\right)^{-\frac{n+1}{2}} d u= \\
& =\frac{2}{\sqrt{n} B\left(\frac{1}{2}, \frac{n}{2}\right)} \int_{0}^{t}\left(1+\frac{u^{2}}{n}\right)^{-\frac{n+1}{2}} d u= \\
& =\frac{-2}{\sqrt{n} B\left(\frac{1}{2}, \frac{n}{2}\right)} \int_{1}^{\frac{n}{n+t^{2}}} \frac{x^{\frac{n+1}{2}} n \sqrt{x}}{2 x^{2} \sqrt{n} \sqrt{1-x}} d x= \\
& =\frac{1}{B\left(\frac{1}{2}, \frac{n}{2}\right)} \int_{\frac{n}{n+t^{2}}}^{1}(1-x)^{-\frac{1}{2}} x^{\frac{n}{2}-1} d x= \\
& =\frac{1}{B\left(\frac{1}{2}, \frac{n}{2}\right)}\left(B\left(\frac{n}{2}, \frac{1}{2}\right)-B_{\frac{n}{n+t^{2}}}\left(\frac{n}{2}, \frac{1}{2}\right)\right)= \\
& \left.=1-I_{\frac{n}{n+t^{2}}}^{2} \frac{n}{2}, \frac{1}{2}\right)=I_{\frac{t^{2}}{n+t^{2}}}\left(\frac{1}{2}, \frac{n}{2}\right)
\end{aligned}
$$

where we have made the substitution $x=n /\left(n+u^{2}\right)$ in order to simplify the integration. From this we find the cumulative function as

$$
F(t)= \begin{cases}\frac{1}{2}-\frac{1}{2} I_{\frac{t^{2}}{}}\left(\frac{1}{2}, \frac{n}{2}\right) & \text { for }-\infty<x<0 \\ \frac{1}{2}+\frac{1}{2} I_{\frac{x^{2}}{2}}^{n+x^{2}}\left(\frac{1}{2}, \frac{n}{2}\right) & \text { for } 0 \leq x<\infty\end{cases}
$$

### 38.5 Relations to Other Distributions

The distribution in $F=t^{2}$ is given by

$$
f(F)=\left|\frac{d t}{d F}\right| f(t)=\frac{1}{2 \sqrt{F}} \cdot \frac{\left(1+\frac{F}{n}\right)^{-\frac{n+1}{2}}}{\sqrt{n} B\left(\frac{1}{2}, \frac{n}{2}\right)}=\frac{n^{\frac{n}{2}} F^{-\frac{1}{2}}}{B\left(\frac{1}{2}, \frac{n}{2}\right)(F+n)^{\frac{n+1}{2}}}
$$

which we recognize as a $F$-distribution with 1 and $n$ degrees of freedom.
As $n \rightarrow \infty$ the Student's $t$-distribution approaches the standard normal distribution. However, a better approximation than to create a simpleminded standardized variable, dividing by the square root of the variance, is to use

$$
z=\frac{t\left(1-\frac{1}{4 n}\right)}{\sqrt{1+\frac{t^{2}}{2 n}}}
$$

which is more closely distributed according to the standard normal distribution.

## $38.6 \quad t$-ratio

Regard $t=x / \sqrt{\frac{y}{n}}$ where $x$ and $y$ are independent variables distributed according to the standard normal and the chi-square distribution with $n$ degrees of freedom, respectively. The independence implies that the joint probability function in $x$ and $y$ is given by

$$
f(x, y ; n)=\left(\frac{1}{\sqrt{2 \pi}} e^{-\frac{x^{2}}{2}}\right)\left(\frac{y^{\frac{n}{2}-1} e^{-\frac{y}{2}}}{2^{\frac{n}{2}} \Gamma\left(\frac{n}{2}\right)}\right)
$$

where $-\infty<x<\infty$ and $y>0$. If we change variables to $t=x / \sqrt{\frac{y}{n}}$ and $u=y$ the distribution in $t$ and $u$, with $-\infty<t<\infty$ and $u>0$, becomes

$$
f(t, u ; n)=\left\|\frac{\partial(x, y)}{\partial(t, u)}\right\| f(x, y ; n)
$$

The determinant is $\sqrt{\frac{u}{n}}$ and thus we have

$$
f(t, u ; n)=\sqrt{\frac{u}{n}}\left(\frac{1}{\sqrt{2 \pi}} e^{-\frac{u t^{2}}{2 n}}\right)\left(\frac{u^{\frac{n}{2}-1} e^{-\frac{u}{2}}}{2^{\frac{n}{2}} \Gamma\left(\frac{n}{2}\right)}\right)=\frac{u^{\frac{1}{2}(n+1)-1} e^{-\frac{u}{2}\left(1+\frac{t^{2}}{n}\right)}}{\sqrt{n \pi} \Gamma\left(\frac{n}{2}\right) 2^{\frac{n+1}{2}}}
$$

Finally, since we are interested in the marginal distribution in $t$ we integrate over $u$

$$
\begin{aligned}
f(t ; n) & =\int_{0}^{\infty} f(t, u ; n) d u=\frac{1}{\sqrt{n \pi} \Gamma\left(\frac{n}{2}\right) 2^{\frac{n+1}{2}}} \int_{0}^{\infty} u^{\frac{n+1}{2}-1} e^{-\frac{u}{2}\left(1+\frac{t^{2}}{n}\right)} d u= \\
& =\frac{1}{\sqrt{n \pi} \Gamma\left(\frac{n}{2}\right) 2^{\frac{n+1}{2}}} \int_{0}^{\infty}\left(\frac{2 v}{1+\frac{t^{2}}{n}}\right)^{\frac{n+1}{2}-1} e^{-v} \frac{d v}{\frac{1}{2}\left(1+\frac{t^{2}}{n}\right)}= \\
& =\frac{\left(1+\frac{t^{2}}{n}\right)^{-\frac{n+1}{2}}}{\sqrt{n \pi} \Gamma\left(\frac{n}{2}\right)} \int_{0}^{\infty} v^{\frac{n+1}{2}-1} e^{-v} d v=\frac{\left(1+\frac{t^{2}}{n}\right)^{-\frac{n+1}{2}}}{\sqrt{n \pi} \Gamma\left(\frac{n}{2}\right)} \Gamma\left(\frac{n+1}{2}\right)=\frac{\left(1+\frac{t^{2}}{n}\right)^{-\frac{n+1}{2}}}{\sqrt{n} B\left(\frac{1}{2}, \frac{n}{2}\right)}
\end{aligned}
$$

where we made the substitution $v=\frac{u}{2}\left(1+\frac{t^{2}}{n}\right)$ in order to simplify the integral which in the last step is recognized as being equal to $\Gamma\left(\frac{n+1}{2}\right)$.

### 38.7 One Normal Sample

Regard a sample from a normal population $N\left(\mu, \sigma^{2}\right)$ where the mean value $\bar{x}$ is distributed as $N\left(\mu, \frac{\sigma^{2}}{n}\right)$ and $\frac{(n-1) s^{2}}{\sigma^{2}}$ is distributed according to the chi-square distribution with $n-1$ degrees of freedom. Here $s^{2}$ is the usual unbiased variance estimator $s^{2}=\frac{1}{n-1} \sum_{i=1}^{n}\left(x_{i}-\bar{x}\right)^{2}$ which in the case of a normal distribution is independent of $\bar{x}$. This implies that

$$
t=\frac{\frac{\bar{x}-\mu}{\sigma / \sqrt{n}}}{\sqrt{\frac{(n-1) s^{2}}{\sigma^{2}} /(n-1)}}=\frac{\bar{x}-\mu}{s / \sqrt{n}}
$$

is distributed according to Student's $t$-distribution with $n-1$ degrees of freedom. We may thus use Student's $t$-distribution to test the hypothesis that $\bar{x}=\mu$ (see below).

### 38.8 Two Normal Samples

Regard two samples $\left\{x_{1}, x_{2}, \ldots, x_{m}\right\}$ and $\left\{y_{1}, y_{2}, \ldots, y_{n}\right\}$ from normal distributions having the same variance $\sigma^{2}$ but possibly different means $\mu_{x}$ and $\mu_{y}$, respectively. Then the quantity $(\bar{x}-\bar{y})-\left(\mu_{x}-\mu_{y}\right)$ has a normal distribution with zero mean and variance equal to $\sigma^{2}\left(\frac{1}{m}+\frac{1}{n}\right)$. Furthermore the pooled variance estimate

$$
s^{2}=\frac{(m-1) s_{x}^{2}+(n-1) s_{y}^{2}}{m+n-2}=\frac{\sum_{i=1}^{m}\left(x_{i}-\bar{x}\right)^{2}+\sum_{i=1}^{n}\left(y_{i}-\bar{y}\right)^{2}}{m+n-2}
$$

is a normal theory estimate of $\sigma^{2}$ with $m+n-2$ degrees of freedom ${ }^{10}$.
Since $s^{2}$ is independent of $\bar{x}$ for normal populations the variable

$$
t=\frac{(\bar{x}-\bar{y})-\left(\mu_{x}-\mu_{y}\right)}{s \sqrt{\frac{1}{m}+\frac{1}{n}}}
$$

has the $t$-distribution with $m+n-2$ degrees of freedom. We may thus use Student's $t$-distribution to test the hypotheses that $\bar{x}-\bar{y}$ is consistent with $\delta=\mu_{x}-\mu_{y}$. In particular we may test if $\delta=0$ i.e. if the two samples originate from population having the same means as well as variances.

### 38.9 Paired Data

If observations are made in pairs $\left(x_{i}, y_{i}\right)$ for $i=1,2, \ldots, n$ the appropriate test statistic is

$$
t=\frac{\bar{d}}{s_{\bar{d}}}=\frac{\bar{d}}{s_{d} / \sqrt{n}}=\frac{\bar{d}}{\sqrt{\frac{\sum_{i=1}^{n}\left(d_{i}-\bar{d}\right)^{2}}{n(n-1)}}}
$$

where $d_{i}=x_{i}-y_{i}$ and $\bar{d}=\bar{x}-\bar{y}$. This quantity has a $t$-distribution with $n-1$ degrees of freedom. We may also write this $t$-ratio as

$$
t=\frac{\sqrt{n} \cdot \bar{d}}{\sqrt{s_{x}^{2}+s_{y}^{2}-2 C_{x y}}}
$$

where $s_{x}^{2}$ and $s_{y}^{2}$ are the estimated variances of $x$ and $y$ and $C_{x y}$ is the covariance between them. If we would not pair the data the covariance term would be zero but the number of degrees of freedom $2 n-2$ i.e. twice as large. The smaller number of degrees of freedom in the paired case is, however, often compensated for by the inclusion of the covariance.

### 38.10 Confidence Levels

In determining confidence levels or testing hypotheses using the $t$-distribution we define the quantity $t_{\alpha, n}$ from

$$
F\left(t_{\alpha, n}\right)=\int_{-\infty}^{t_{\alpha, n}} f(t ; n) d t=1-\alpha
$$

[^8]i.e. $\alpha$ is the probability that a variable distributed according to the $t$-distribution with $n$ degrees of freedom exceeds $t_{\alpha, n}$. Note that due to the symmetry about zero of the $t$-distribution $t_{\alpha, n}=-t_{1-\alpha, n}$.

In the case of one normal sample described above we may set a $1-\alpha$ confidence interval for $\mu$

$$
\bar{x}-\frac{s}{\sqrt{n}} t_{\alpha / 2, n-1} \leq \mu \leq \bar{x}+\frac{s}{\sqrt{n}} t_{\alpha / 2, n-1}
$$

Note that in the case where $\sigma^{2}$ is known we would not use the $t$-distribution. The appropriate distribution to use in order to set confidence levels in this case would be the normal distribution.

### 38.11 Testing Hypotheses

As indicated above we may use the $t$-statistics in order to test hypotheses regarding the means of populations from normal distributions.

In the case of one sample the null hypotheses would be $H_{0}: \mu=\mu_{0}$ and the alternative hypothesis $H_{1}: \mu \neq \mu_{0}$. We would then use $t=\frac{\bar{x}-\mu_{0}}{s / \sqrt{n}}$ as outlined above and reject $H_{0}$ at the $\alpha$ confidence level of significance if $|t|>\frac{s}{\sqrt{n}} t_{\alpha / 2, n-1}$. This test is two-tailed since we do not assume any a priori knowledge of in which direction an eventual difference would be. If the alternate hypothesis would be e.g. $H_{1}: \mu>\mu_{0}$ then a one-tailed test would be appropriate.

The probability to reject the hypothesis $H_{0}$ if it is indeed true is thus $\alpha$. This is a so called Type I error. However, we might also be interested in the probability of committing a Type II error implying that we would accept the hypothesis although it was wrong and the distribution instead had a mean $\mu_{1}$. In addressing this question the $t$-distribution could be modified yielding the non-central $t$-distribution. The probability content $\beta$ of this distribution in the confidence interval used would then be the probability of wrongly accepting the hypothesis. This calculation would depend on the choice of $\alpha$ as well as on the number of observations $n$. However, we do not describe details about this here.

In the two sample case we may want to test the null hypothesis $H_{0}: \mu_{x}=\mu_{y}$ as compared to $H_{1}: \mu_{x} \neq \mu_{y}$. Once again we would reject $H_{0}$ if the absolute value of the quantity $t=(\bar{x}-\bar{y}) / s \sqrt{\frac{1}{m}+\frac{1}{n}}$ would exceed $t_{\alpha / 2, n+m-2}$.

### 38.12 Calculation of Probability Content

In order to find confidence intervals or to test hypotheses we must be able to calculate integrals of the probability density function over certain regions. We recall the formula

$$
F\left(t_{\alpha, n}\right)=\int_{-\infty}^{t_{\alpha, n}} f(t ; n) d t=1-\alpha
$$

which defines the quantity $t_{\alpha, n}$ for a specified confidence level $\alpha$. The probability to get a value equal to $t_{\alpha, n}$ or higher is thus $\alpha$.

Classically all text-books in statistics are equipped with tables giving values of $t_{\alpha, n}$ for specific $\alpha$-values. This is sometimes useful and in table 8 on page 178 we show such a
table giving points where the distribution has a cumulative probability content of $1-\alpha$ for different number of degrees of freedom.

However, it is often preferable to calculate directly the exact probability that one would observe the actual $t$-value or worse. To calculate the integral on the left-hand side we differ between the case where the number of degrees of freedom is an odd or even integer. The equation above may either be adjusted such that a required $\alpha$ is obtained or we may replace $t_{\alpha, n}$ with the actual $t$-value found in order to calculate the probability for the present outcome of the experiment.

The algorithm proposed for calculating the probability content of the $t$-distribution is described in the following subsections.

### 38.12.1 Even number of degrees of freedom

For even $n$ we have putting $m=\frac{n}{2}$ and making the substitution $x=\frac{t}{\sqrt{n}}$

$$
1-\alpha=\int_{-\infty}^{t_{\alpha, n}} f(t ; n) d t=\frac{\Gamma\left(\frac{n+1}{2}\right)}{\sqrt{n \pi} \Gamma\left(\frac{n}{2}\right)} \int_{-\infty}^{t_{\alpha, n}} \frac{d t}{\left(1+\frac{t^{2}}{n}\right)^{\frac{n+1}{2}}}=\frac{\Gamma\left(m+\frac{1}{2}\right)}{\sqrt{\pi} \Gamma(m)} \int_{-\infty}^{t_{\alpha, n} / \sqrt{n}} \frac{d x}{\left(1+x^{2}\right)^{m+\frac{1}{2}}}
$$

For convenience (or maybe it is rather laziness) we make use of standard integral tables where we find the integral

$$
\int \frac{d x}{\left(a x^{2}+c\right)^{m+\frac{1}{2}}}=\frac{x}{\sqrt{a x^{2}+c}} \sum_{r=0}^{m-1} \frac{2^{2 m-2 r-1}(m-1)!m!(2 r)!}{(2 m)!(r!)^{2} c^{m-r}\left(a x^{2}+c\right)^{r}}
$$

where in our case $a=c=1$. Introducing $x_{\alpha}=t_{\alpha, n} / \sqrt{n}$ for convenience this gives

$$
1-\alpha=\frac{\Gamma\left(m+\frac{1}{2}\right)(m-1)!m!2^{2 m}}{\sqrt{\pi} \Gamma(m)(2 m)!} \cdot\left[\frac{x_{\alpha}}{2 \sqrt{1+x_{\alpha}^{2}}} \sum_{r=0}^{m-1} \frac{(2 r)!}{2^{2 r}(r!)^{2}\left(1+x_{\alpha}^{2}\right)^{r}}+\frac{1}{2}\right]
$$

The last term inside the brackets is the value of the integrand at $-\infty$ which is seen to equal $-\frac{1}{2}$. Looking at the factor outside the brackets using that $\Gamma(n)=(n-1)$ ! for $n$ a positive integer, $\Gamma\left(m+\frac{1}{2}\right)=\frac{(2 m-1)!!}{2^{m}} \sqrt{\pi}$, and rewriting $(2 m)!=(2 m)!!(2 m-1)!!=2^{m} m!(2 m-1)!!$ we find that it in fact is equal to one. We thus have

$$
1-\alpha=\frac{x_{\alpha}}{2 \sqrt{1+x_{\alpha}^{2}}} \sum_{r=0}^{m-1} \frac{(2 r)!}{2^{2 r}(r!)^{2}\left(1+x_{\alpha}^{2}\right)^{r}}+\frac{1}{2}
$$

In evaluating the sum it is useful to look at the individual terms. Denoting these by $u_{r}$ we find the recurrence relation

$$
u_{r}=u_{r-1} \cdot \frac{2 r(2 r-1)}{r^{2} 2^{2}\left(1+x_{\alpha}^{2}\right)}=u_{r-1} \cdot \frac{1-\frac{1}{2 r}}{1+x_{\alpha}^{2}}
$$

where we start with $u_{0}=1$.

To summarize: in order to determine the probability $\alpha$ to observe a value $t$ or bigger from a $t$-distribution with an even number of degrees of freedom $n$ we calculate

$$
1-\alpha=\frac{\frac{t}{\sqrt{n}}}{2 \sqrt{1+\frac{t^{2}}{n}}} \sum_{r=0}^{m-1} u_{r}+\frac{1}{2}
$$

where $u_{0}=1$ and $u_{r}=u_{r-1} \cdot \frac{1-\frac{1}{2 r}}{1+t^{2} / n}$.

### 38.12.2 Odd number of degrees of freedom

For odd $n$ we have putting $m=\frac{n-1}{2}$ and making the substitution $x=\frac{t}{\sqrt{n}}$

$$
1-\alpha=\int_{-\infty}^{t_{\alpha, n}} f(t ; n) d t=\frac{\Gamma\left(\frac{n+1}{2}\right)}{\sqrt{n \pi} \Gamma\left(\frac{n}{2}\right)} \int_{-\infty}^{t_{\alpha, n}} \frac{d t}{\left(1+\frac{t^{2}}{n}\right)^{\frac{n+1}{2}}}=\frac{\Gamma(m+1)}{\sqrt{\pi} \Gamma\left(m+\frac{1}{2}\right)} \int_{-\infty}^{x_{\alpha}} \frac{d x}{\left(1+x^{2}\right)^{m+1}}
$$

where we again have introduced $x_{\alpha}=t_{\alpha, n} / \sqrt{n}$. Once again we make use of standard integral tables where we find the integral

$$
\int \frac{d x}{\left(a+b x^{2}\right)^{m+1}}=\frac{(2 m)!}{(m!)^{2}}\left[\frac{x}{2 a} \sum_{r=1}^{m} \frac{r!(r-1)!}{(4 a)^{m-r}(2 r)!\left(a+b x^{2}\right)^{r}}+\frac{1}{(4 a)^{m}} \int \frac{d x}{a+b x^{2}}\right]
$$

where in our case $a=b=1$. We obtain

$$
1-\alpha=\frac{\Gamma(m+1)(2 m)!}{\sqrt{\pi} \Gamma\left(m+\frac{1}{2}\right) m!^{2} 4^{m}}\left[\frac{x_{\alpha}}{2} \sum_{r=1}^{m} \frac{4^{r} r!(r-1)!}{(2 r)!\left(1+x_{\alpha}^{2}\right)^{r}}+\arctan x_{\alpha}+\frac{\pi}{2}\right]
$$

where the last term inside the brackets is the value of the integrand at $-\infty$. The factor outside the brackets is equal to $\frac{1}{\pi}$ which is found using that $\Gamma(n)=(n-1)$ ! for $n$ a positive integer, $\Gamma\left(m+\frac{1}{2}\right)=\frac{(2 m-1)!!}{2^{m}} \sqrt{\pi}$, and $(2 m)!=(2 m)!!(2 m-1)!!=2^{m}(m)!(2 m-1)!!$. We get

$$
\begin{aligned}
1-\alpha & =\frac{1}{\pi}\left[\frac{x_{\alpha}}{2} \sum_{r=1}^{m} \frac{4^{r} r!(r-1)!}{(2 r)!\left(1+x_{\alpha}^{2}\right)^{r}}+\arctan x_{\alpha}+\frac{\pi}{2}\right]= \\
& =\frac{1}{\pi}\left[\frac{x_{\alpha}}{1+x_{\alpha}^{2}} \sum_{r=1}^{m} \frac{2^{2 r-1} r!(r-1)!}{(2 r)!\left(1+x_{\alpha}^{2}\right)^{r-1}}+\arctan x_{\alpha}+\frac{\pi}{2}\right]
\end{aligned}
$$

To compute the sum we denote the terms by $v_{r}$ and find the recurrence relation

$$
v_{r}=v_{r-1} \frac{4 r(r-1)}{2 r(2 r-1)\left(1+x_{\alpha}^{2}\right)}=v_{r-1} \frac{\left(1-\frac{1}{2 r-1}\right)}{\left(1+x_{\alpha}^{2}\right)}
$$

starting with $v_{1}=1$.
To summarize: in order to determine the probability $\alpha$ to observe a value $t$ or bigger from a $t$-distribution with an odd number of degrees of freedom $n$ we calculate

$$
1-\alpha=\frac{1}{\pi}\left[\frac{\frac{t}{\sqrt{n}}}{1+\frac{t^{2}}{n}} \sum_{r=1}^{\frac{n-1}{2}} v_{r}+\arctan \frac{t}{\sqrt{n}}\right]+\frac{1}{2}
$$

where $v_{1}=1$ and $v_{r}=v_{r-1} \cdot \frac{1-\frac{1}{2 r-1}}{1+t^{2} / n}$.

### 38.12.3 Final algorithm

The final algorithm to evaluate the probability content from $-\infty$ to $t$ for a $t$-distribution with $n$ degrees of freedom is

- Calculate $x=\frac{t}{\sqrt{n}}$
- For $n$ even:
- Put $m=\frac{n}{2}$
- Set $u_{0}=1, s=0$ and $i=0$.
- For $i=0,1,2, \ldots, m-1$ set $s=s+u_{i}, i=i+1$ and $u_{i}=u_{i-1} \frac{1-\frac{1}{2 i}}{1+x^{2}}$.
- $\alpha=\frac{1}{2}-\frac{1}{2} \cdot \frac{x}{\sqrt{1+x^{2}}} s$.
- For $n$ odd:
- Put $m=\frac{n-1}{2}$.
- Set $v_{1}=1, s=0$ and $i=1$.
- For $i=1,2, \ldots, m$ set $s=s+v_{i}, i=i+1$ and $v_{i}=v_{i-1} \cdot \frac{1-\frac{1}{2 i-1}}{1+x^{2}}$.
- $\alpha=\frac{1}{2}-\frac{1}{\pi}\left(\frac{x}{1+x^{2}} \cdot s+\arctan x\right)$.


### 38.13 Random Number Generation

Following the definition we may define a random number $t$ from a $t$-distribution, using random numbers from a normal and a chi-square distribution, as

$$
t=\frac{z}{\sqrt{y_{n} / n}}
$$

where $z$ is a standard normal and $y_{n}$ a chi-squared variable with $n$ degrees of freedom. To obtain random numbers from these distributions see the appropriate sections.

## 39 Triangular Distribution

### 39.1 Introduction

The triangular distribution is given by

$$
f(x ; \mu, \Gamma)=\frac{-|x-\mu|}{\Gamma^{2}}+\frac{1}{\Gamma}
$$

where the variable x is bounded to the interval $\mu-\Gamma \leq x \leq \mu+\Gamma$ and the location and scale parameters $\mu$ and $\Gamma(\Gamma>0)$ all are real numbers.

### 39.2 Moments

The expectation value of the distribution is $E(x)=\mu$. Due to the symmetry of the distribution odd central moments vanishes while even moments are given by

$$
\mu_{n}=\frac{2 \Gamma^{n}}{(n+1)(n+2)}
$$

for even values of $n$. In particular the variance $V(x)=\mu_{2}=\Gamma^{2} / 6$ and the fourth central moment $\mu_{4}=\Gamma^{4} / 15$. The coefficient of skewness is zero and the coefficient of kurtosis $\gamma_{2}=-0.6$.

### 39.3 Random Number Generation

The sum of two pseudorandom numbers uniformly distributed between $(\mu-\Gamma) / 2$ and $(\mu+\Gamma) / 2$ is distributed according to the triangular distribution. If $\xi_{1}$ and $\xi_{2}$ are uniformly distributed between zero and one then

$$
x=\mu+\left(\xi_{1}+\xi_{2}-1\right) \Gamma \quad \text { or } \quad x=\mu+\left(\xi_{1}-\xi_{2}\right) \Gamma
$$

follow the triangular distribution.
Note that this is a special case of a combination

$$
x=(a+b) \xi_{1}+(b-a) \xi_{2}-b
$$

with $b>a \geq 0$ which gives a random number from a symmetric trapezoidal distribution with vertices at $( \pm b, 0)$ and $\left( \pm a, \frac{1}{a+b}\right)$.

## 40 Uniform Distribution

### 40.1 Introduction

The uniform distribution is, of course, a very simple case with

$$
f(x ; a, b)=\frac{1}{b-a} \quad \text { for } \quad a \leq x \leq b
$$

The cumulative, distribution, function is thus given by

$$
F(x ; a, b)=\left\{\begin{array}{lll}
0 & \text { if } & x \leq a \\
\frac{x-a}{b-a} & \text { if } & a \leq x \leq b \\
1 & \text { if } & b \leq x
\end{array}\right.
$$

### 40.2 Moments

The uniform distribution has expectation value $E(x)=(a+b) / 2$, variance $V(x)=(b-$ $a)^{2} / 12, \mu_{3}=0, \mu_{4}=(b-a)^{4} / 80$, coefficient of skewness $\gamma_{1}=0$ and coefficient of kurtosis $\gamma_{2}=-1.2$. More generally all odd central moments vanish and for $n$ an even integer

$$
\mu_{n}=\frac{(b-a)^{n}}{2^{n}(n+1)}
$$

### 40.3 Random Number Generation

Since we assume the presence of a pseudorandom number generator giving random numbers $\xi$ between zero and one a random number from the uniform distribution is simply given by

$$
x=(b-a) \xi+a
$$

## 41 Weibull Distribution

### 41.1 Introduction

The Weibull distribution is given by

$$
f(x ; \eta, \sigma)=\frac{\eta}{\sigma}\left(\frac{x}{\sigma}\right)^{\eta-1} e^{-\left(\frac{x}{\sigma}\right)^{\eta}}
$$

where the variable $x$ and the parameters $\eta$ and $\sigma$ all are positive real numbers. The distribution is named after the swedish physicist Waloddi Weibull (1887-1979) a professor at the Technical Highschool in Stockholm 1924-1953.

The parameter $\sigma$ is simply a scale parameter and the variable $y=x / \sigma$ has the distribution

$$
g(y)=\eta y^{\eta-1} e^{-y^{\eta}}
$$

In figure 32 we show the distribution for a few values of $\eta$. For $\eta<1$ the distribution has its mode at $y=0$, at $\eta=1$ it is identical to the exponential distribution, and for $\eta>1$ the distribution has a mode at

$$
x=\left(\frac{\eta-1}{\eta}\right)^{\frac{1}{\eta}}
$$

which approaches $x=1$ as $\eta$ increases (at the same time the distribution gets more symmetric and narrow).


Figure 32: The Weibull distribution

### 41.2 Cumulative Distribution

The cumulative distribution is given ,by

$$
F(x)=\int_{0}^{x} f(u) d u=\int_{0}^{x} \frac{\eta}{\sigma}\left(\frac{u}{\sigma}\right)^{\eta-1} e^{-\left(\frac{u}{\sigma}\right)^{\eta}} d u=\int_{0}^{(x / \sigma)^{\eta}} e^{-y} d y=1-e^{-\left(\frac{x}{\sigma}\right)^{\eta}}
$$

where we have made the substitution $y=(u / \sigma)^{\eta}$ in order to simplify the integration.

### 41.3 Moments

Algebraic moments are given by

$$
E\left(x^{k}\right)=\int_{0}^{\infty} x^{k} f(x) d x=\sigma^{k} \int_{0}^{\infty} y^{\frac{k}{\eta}} e^{-y} d y=\sigma^{k} \Gamma\left(\frac{k}{\eta}+1\right)
$$

where we have made the same substitution as was used when evaluating the cumulative distribution above.

Especially the expectation value and the variance are given by

$$
E(x)=\sigma \Gamma\left(\frac{1}{\eta}+1\right) \quad \text { and } \quad V(x)=\sigma^{2}\left\{\Gamma\left(\frac{2}{\eta}+1\right)-\Gamma\left(\frac{1}{\eta}+1\right)^{2}\right\}
$$

### 41.4 Random Number Generation

To obtain random numbers from Weibull's distribution using $\xi$, a random number uniformly distributed from zero to one, we may solve the equation $F(x)=\xi$ to obtain a random number in $x$.

$$
F(x)=1-e^{-\left(\frac{x}{\sigma}\right)^{\eta}}=\xi \quad \Rightarrow \quad x=\sigma(-\ln \xi)^{\frac{1}{\eta}}
$$

## 42 Appendix A: The Gamma and Beta Functions

### 42.1 Introduction

In statistical calculations for standard statistical distributions such as the normal (or Gaussian) distribution, the Student's $t$-distribution, the chi-squared distribution, and the $F$ distribution one often encounters the so called Gamma and Beta functions. More specifically in calculating the probability content for these distributions the incomplete Gamma and Beta functions occur. In the following we briefly define these functions and give numerical methods on how to calculate them. Also connections to the different statistical distributions are given. The main references for this has been $[41,42,43]$ for the formalism and [10] for the numerical methods.

### 42.2 The Gamma Function

The Gamma function is normally defined as

$$
\Gamma(z)=\int_{0}^{\infty} t^{z-1} e^{-t} d t
$$

where $z$ is a complex variable with $\operatorname{Re}(z)>0$. This is the so called Euler's integral form for the Gamma function. There are, however, two other definitions worth mentioning. Firstly Euler's infinite limit form

$$
\Gamma(z)=\lim _{n \rightarrow \infty} \frac{1 \cdot 2 \cdot 3 \cdots n}{z(z+1)(z+2) \cdots(z+n)} n^{z} \quad z \neq 0,-1,-2, \ldots
$$

and secondly the infinite product form sometimes attributed to Euler and sometimes to Weierstrass

$$
\frac{1}{\Gamma(z)}=z^{\gamma z} \prod_{n=1}^{\infty}\left(1+\frac{z}{n}\right) e^{-\frac{z}{n}} \quad|z|<\infty
$$

where $\gamma \approx 0.5772156649$ is Euler's constant.
In figure 33 we show the Gamma function for real arguments from -5 to 5 . Note the singularities at $x=0,-1,-2, \ldots$.

For $z$ a positive real integer $n$ we have the well known relation to the factorial function

$$
n!=\Gamma(n+1)
$$

and, as the factorial function, the Gamma function satisfies the recurrence relation

$$
\Gamma(z+1)=z \Gamma(z)
$$

In the complex plane $\Gamma(z)$ has a pole at $z=0$ and at negative integer values of $z$. The reflection formula

$$
\Gamma(1-z)=\frac{\pi}{\Gamma(z) \sin (\pi z)}=\frac{\pi z}{\Gamma(z+1) \sin (\pi z)}
$$

may be used in order to get function values for $\operatorname{Re}(z)<1$ from values for $\operatorname{Re}(z)>1$.


Figure 33: The Gamma function

A well known approximation to the Gamma function is Stirling's formula

$$
\Gamma(z)=z^{z} e^{-z} \sqrt{\frac{2 \pi}{z}}\left(1+\frac{1}{12 z}+\frac{1}{288 z^{2}}-\frac{139}{51840 z^{3}}-\frac{571}{2488320 z^{4}}+\ldots\right)
$$

for $|\arg z|<\pi$ and $|z| \rightarrow \infty$ and where often only the first term (1) in the series expansion is kept in approximate calculations. For the faculty of a positive integer $n$ one often uses the approximation

$$
n!\approx \sqrt{2 \pi n} n^{n} e^{-n}
$$

which has the same origin and also is called Stirling's formula.

### 42.2.1 Numerical Calculation

There are several methods to calculate the Gamma function in terms of series expansions etc. For numerical calculations, however, the formula by Lanczos is very useful [10]

$$
\Gamma(z+1)=\left(z+\gamma+\frac{1}{2}\right)^{z+\frac{1}{2}} e^{-\left(z+\gamma+\frac{1}{2}\right)} \sqrt{2 \pi}\left[c_{0}+\frac{c_{1}}{z+1}+\frac{c_{2}}{z+2}+\cdots+\frac{c_{n}}{z+n}+\epsilon\right]
$$

for $z>0$ and an optimal choice of the parameters $\gamma, n$, and $c_{0}$ to $c_{n}$. For $\gamma=5$, $n=6$ and a certain set of $c$ 's the error is smaller than $|\epsilon|<2 \cdot 10^{-10}$. This bound is true for all complex $z$ in the half complex plane $\operatorname{Re}(z)>0$. The coefficients normally used are $c_{0}=1, c_{1}=76.18009173, c_{2}=-86.50532033, c_{3}=24.01409822, c_{4}=-1.231739516$, $c_{5}=0.00120858003$, and $c_{6}=-0.00000536382$. Use the reflection formula given above to obtain results for $\operatorname{Re}(z)<1 \mathrm{e} . g$. for negative real arguments. Beware, however, to avoid the singularities. While implementing routines for the Gamma function it is recommendable to evaluate the natural logarithm in order to avoid numerical overflow.

An alternative way of evaluating $\ln \Gamma(z)$ is given in references [44,45], giving formulæ which also are used in order to evaluate the Digamma function below. The expressions used for $\ln \Gamma(z)$ are

$$
\ln \Gamma(z)= \begin{cases}\left(z-\frac{1}{2}\right) \ln z-z+\frac{1}{2} \ln 2 \pi+z \sum_{k=1}^{K} \frac{B_{2 k}}{2 k(2 k-1)} z^{-2 k}+R_{K}(z) & \text { for } 0<x_{0} \leq x \\ \ln \Gamma(z+n)-\ln \prod_{k=0}^{n-1}(z+k) & \text { for } 0 \leq x<x_{0} \\ \ln \pi+\ln \Gamma(1-z)-\ln \sin \pi z & \text { for } x<0\end{cases}
$$

Here $n=\left[x_{0}\right]-[x]$ (the difference of integer parts, where $x$ is the real part of $z=x+\imath y$ ) and e.g. $K=10$ and $x_{0}=7.0$ gives excellent accuracy i.e. small $R_{K}$. Note that Kölbig [45] gives the wrong sign on the (third) constant term in the first case above.

### 42.2.2 Formulæ

Below we list some useful relations concerning the Gamma function, faculties and semifaculties (denoted by two exclamation marks here). For a more complete list consult e.g. [42].

$$
\begin{aligned}
\Gamma(z) & =\int_{0}^{1}\left(\ln \frac{1}{t}\right)^{z-1} d t \\
\Gamma(z+1) & =z \Gamma(z)=z! \\
\Gamma(z) & =\alpha^{z} \int_{0}^{\infty} t^{z-1} e^{-\alpha t} d t \quad \text { for } \quad \operatorname{Re}(z)>0, \quad \operatorname{Re}(\alpha)>0 \\
\Gamma(k) & =(k-1)!\quad \text { for } \quad k \geq 1 \quad(\text { integer }, 0!=1) \\
z! & =\Gamma(z+1)=\int_{0}^{\infty} e^{-t} t^{z} d t \quad \text { for } \quad \operatorname{Re}(z)>-1 \\
\Gamma\left(\frac{1}{2}\right) & =\sqrt{\pi} \\
\Gamma\left(n+\frac{1}{2}\right) & =\frac{(2 n-1)!!}{2^{n}} \sqrt{\pi} \\
\Gamma(z) \Gamma(1-z) & =\frac{\pi}{\sin \pi z} \\
z!(-z)! & =\frac{\pi z}{\sin \pi z} \\
(2 m)!! & =2 \cdot 4 \cdot 6 \cdots 2 m=2^{m} m! \\
(2 m-1)!! & =1 \cdot 3 \cdot 5 \cdots(2 m-1) \\
(2 m)! & =(2 m)!(2 m-1)!!=2^{m} m!(2 m-1)!!
\end{aligned}
$$

### 42.3 Digamma Function

It is often convenient to work with the logarithm of the Gamma function in order to avoid numerical overflow in the calculations. The first derivatives of this function

$$
\psi(z)=\frac{d}{d z} \ln \Gamma(z)=\frac{1}{\Gamma(z)} \frac{d \Gamma(z)}{d z}
$$

is known as the Digamma, or Psi, function. A series expansion of this function is given by

$$
\psi(z+1)=-\gamma-\sum_{n=1}^{\infty}\left(\frac{1}{z+n}-\frac{1}{n}\right) \quad \text { for } \quad z \neq 0,-1,-2,-3, \ldots
$$

where $\gamma \approx 0.5772156649$ is Euler's constant which is seen to be equal to $-\psi(1)$. If the derivative of the Gamma function itself is required we may thus simply use $d \Gamma(z) / d z=$ $\Gamma(z) \cdot \psi(z)$. Note that some authors write $\psi(z)=\frac{d}{d z} \ln \Gamma(z+1)=\frac{d}{d z}$ ! for the Digamma function, and similarly for the polygamma functions below, thus shifting the argument by one unit.

In figure 34 we show the Digamma function for real arguments from -5 to 5 . Note the singularities at $x=0,-1,-2, \ldots$..


Figure 34: The Digamma, or Psi, function
For integer values of $z$ we may write

$$
\psi(n)=-\gamma+\sum_{m=1}^{n-1} \frac{1}{m}
$$

which is efficient enough for numerical calculations for not too large values of $n$. Similarly for half-integer values we have

$$
\psi\left(n+\frac{1}{2}\right)=-\gamma-2 \ln 2+2 \sum_{m=1}^{n} \frac{1}{2 m-1}
$$

However, for arbitrary arguments the series expansion above is unusable. Following the recipe given in an article by K. S. Kölbig [45] we use

$$
\psi(z)= \begin{cases}\ln z-\frac{1}{2 z}-\sum_{k=1}^{K} \frac{B_{2 k}}{2 k} z^{-2 k}+R_{K}(z) & \text { for } 0<x_{0} \leq x \\ \psi(z+n)-\sum_{k=0}^{n-1} \frac{1}{z+k} & \text { for } 0 \leq x<x_{0} \\ \psi(-z)+\frac{1}{z}+\pi \cot \pi z & \text { for } x<0\end{cases}
$$

Here $n=\left[x_{0}\right]-[x]$ (the difference of integer parts, where $x$ is the real part of $z=x+\imath y$ ) and we have chosen $K=10$ and $x_{0}=7.0$ which gives a very good accuracy (i.e. small $R_{K}$, typically less than $10^{-15}$ ) for double precision calculations. The main interest in statistical calculations is normally function values for $\psi(x)$ for real positive arguments but the formulæ above are valid for any complex argument except for the singularities along the real axis at $z=0,-1,-2,-3, \ldots$. The $B_{2 k}$ are Bernoulli numbers given by $B_{0}=1, B_{1}=-\frac{1}{2}, B_{2}=\frac{1}{6}, B_{4}=-\frac{1}{30}, B_{6}=\frac{1}{42}, B_{8}=-\frac{1}{30}, B_{10}=\frac{5}{66}, B_{12}=-\frac{691}{2730}, B_{14}=$ $\frac{7}{6}, B_{16}=-\frac{3617}{510}, B_{18}=\frac{43867}{798}, B_{20}=-\frac{174611}{330} \ldots$

### 42.4 Polygamma Function

Higher order derivatives of $\ln \Gamma(z)$ are called Polygamma functions ${ }^{11}$

$$
\psi^{(n)}(z)=\frac{d^{n}}{d z^{n}} \psi(z)=\frac{d^{n+1}}{d z^{n+1}} \ln \Gamma(z) \quad \text { for } \quad n=1,2,3, \ldots
$$

Here a series expansion is given by

$$
\psi^{(n)}(z)=(-1)^{n+1} n!\sum_{k=0}^{\infty} \frac{1}{(z+k)^{n+1}} \quad \text { for } \quad z \neq 0,-1,-2, \ldots
$$

For numerical calculations we have adopted a technique similar to what was used to evaluate $\ln \Gamma(z)$ and $\psi(z)$.

$$
\psi^{(n)}(z)= \begin{cases}(-1)^{n-1}\left[t_{1}+\frac{n!}{2 z^{n+1}}+\sum_{k=1}^{K} B_{2 k} \frac{(2 k+n-1)!}{(2 k)!z^{2 k+n}}+R_{K}(z)\right] & \text { for } 0<x_{0} \leq x \\ \psi^{(n)}(z+m)-(-1)^{n} n!\sum_{k=0}^{m-1} \frac{1}{(z+k)^{n+1}} & \text { for } 0 \leq x<x_{0}\end{cases}
$$

where $t_{1}=-\ln z$ for $n=0$ and $t_{1}=(n-1)!/ z^{n}$ for $n>0$. Here $m=\left[x_{0}\right]-[x]$ i.e. the difference of integer parts, where $x$ is the real part of $z=x+\imath y$. We treat primarily the case for real positive arguments $x$ and if complex arguments are required one ought to add a third reflection formula as was done in the previous cases. Without any special optimization we have chosen $K=14$ and $x_{0}=7.0$ which gives a very good accuracy, i.e. small $R_{K}$, typically less than $10^{-15}$, even for double precision calculations except for higher orders and low values of $x$ where the function value itself gets large. ${ }^{12}$

For more relations on the Polygamma (and the Digamma) functions see e.g. [42]. Two useful relations used in this document in finding cumulants for some distributions are

$$
\begin{aligned}
\psi^{(n)}(1) & =(-1)^{n+1} n!\zeta(n+1) \\
\psi^{(n)}\left(\frac{1}{2}\right) & =(-1)^{n+1} n!\left(2^{n+1}-1\right) \zeta(n+1)=\left(2^{n+1}-1\right) \psi^{(n)}(1)
\end{aligned}
$$

where $\zeta$ is Riemann's zeta function (see page 59 and [31]).

[^9]
### 42.5 The Incomplete Gamma Function

For the incomplete Gamma function there seem to be several definitions in the literature. Defining the two integrals

$$
\gamma(a, x)=\int_{0}^{x} t^{a-1} e^{-t} d t \quad \text { and } \quad \Gamma(a, x)=\int_{x}^{\infty} t^{a-1} e^{-t} d t
$$

with $\operatorname{Re}(a)>0$ the incomplete Gamma function is normally defined as

$$
P(a, x)=\frac{\gamma(a, x)}{\Gamma(a)}
$$

but sometimes also $\gamma(a, x)$ and $\Gamma(a, x)$ is referred to under the same name as well as the complement to $P(a, x)$

$$
Q(a, x)=1-P(a, x)=\frac{\Gamma(a, x)}{\Gamma(a)}
$$

Note that, by definition, $\gamma(a, x)+\Gamma(a, x)=\Gamma(a)$.
In figure 35 the incomplete Gamma function $P(a, x)$ is shown for a few $a$-values ( 0.5 , 1, 5 and 10).


Figure 35: The incomplete Gamma function

### 42.5.1 Numerical Calculation

For numerical evaluations of $P$ two formulæ are useful [10]. For values $x<a+1$ the series

$$
\gamma(a, x)=e^{-x} x^{a} \sum_{n=0}^{\infty} \frac{\Gamma(a)}{\Gamma(a+n+1)} x^{n}
$$

converges rapidly while for $x \geq a+1$ the continued fraction

$$
\Gamma(a, x)=e^{-x} x^{a}\left(\frac{1}{x+} \frac{1-a}{1+} \frac{1}{x+} \frac{2-a}{1+} \frac{2}{x+} \cdots\right)
$$

is a better choice.

### 42.5.2 Formulæ

Below we list some relations concerning the incomplete Gamma function. For a more complete list consult e.g. [42].

$$
\begin{aligned}
\Gamma(a) & =\gamma(a, x)+\Gamma(a, x) \\
\gamma(a, x) & =\int_{0}^{x} e^{-t} t^{a-1} d t \quad \text { for } \quad \operatorname{Re}(a)>0 \\
\gamma(a+1, x) & =a \gamma(a, x)-x^{a} e^{-x} \\
\gamma(n, x) & =(n-1)!\left[1-e^{-x} \sum_{r=0}^{n-1} \frac{x^{r}}{r!}\right] \\
\Gamma(a, x) & =\int_{x}^{\infty} e^{-t} t^{a-1} d t \\
\Gamma(a+1, x) & =a \Gamma(a, x)-x^{a} e^{-x} \\
\Gamma(n, x) & =(n-1)!e^{-x} \sum_{r=0}^{n-1} \frac{x^{r}}{r!} \quad n=1,2, \ldots
\end{aligned}
$$

### 42.5.3 Special Cases

The usage of the incomplete Gamma function $P(a, x)$ in calculations made in this document often involves integer or half-integer values for $a$. These cases may be solved by the following formulæ

$$
\begin{aligned}
P(n, x) & =1-e^{-x} \sum_{k=0}^{n-1} \frac{x^{k}}{k!} \\
P\left(\frac{1}{2}, x\right) & =\operatorname{erf} \sqrt{x} \\
P(a+1, x) & =P(a, x)-\frac{x^{a} e^{-x}}{\Gamma(a+1)}=P(a, x)-\frac{x^{a} e^{-x}}{a \Gamma(a)} \\
P\left(\frac{n}{2}, x\right) & =\operatorname{erf} \sqrt{x}-\sum_{k=1}^{\frac{n-1}{2}} \frac{x^{\frac{2 k-1}{2}} e^{-x}}{\Gamma\left(\frac{2 k+1}{2}\right)}=\operatorname{erf} \sqrt{x}-2 e^{-x} \sqrt{\frac{x}{\pi}} \sum_{k=1}^{\frac{n-1}{2}} \frac{(2 x)^{k-1}}{(2 k-1)!!}
\end{aligned}
$$

the last formula for odd values of $n$.

### 42.6 The Beta Function

The Beta function is defined through the integral formula

$$
B(a, b)=B(b, a)=\int_{0}^{1} t^{a-1}(1-t)^{b-1} d t
$$

and is related to the Gamma function by

$$
B(a, b)=\frac{\Gamma(a) \Gamma(b)}{\Gamma(a+b)}
$$

The most straightforward way to calculate the Beta function is by using this last expression and a well optimized routine for the Gamma function. In table 9 on page 179 expressions for the Beta function for low integer and half-integer arguments are given.

Another integral, obtained by the substitution $x=t /(1-t)$, yielding the Beta function is

$$
B(a, b)=\int_{0}^{\infty} \frac{x^{a-1}}{(1+x)^{a+b}} d x
$$

### 42.7 The Incomplete Beta Function

The incomplete Beta function is defined as

$$
I_{x}(a, b)=\frac{B_{x}(a, b)}{B(a, b)}=\frac{1}{B(a, b)} \int_{0}^{x} t^{a-1}(1-t)^{b-1} d t
$$

for $a, b>0$ and $0 \leq x \leq 1$.
The function $B_{x}(a, b)$, often also called the incomplete Beta function, satisfies the following formula

$$
\begin{aligned}
B_{x}(a, b)= & \int_{0}^{\frac{x}{1-x}} \frac{u^{a-1}}{(1+u)^{a+b}} d u=B_{1}(b, a)-B_{1-x}(b, a)= \\
= & x^{a}\left[\frac{1}{a}+\frac{1-b}{a+1} x+\frac{(1-b)(2-b)}{2!(a+2)} x^{2}+\right. \\
& \left.\cdots+\frac{(1-b)(2-b) \cdots(n-b)}{n!(a+n)} x^{n}+\cdots\right]
\end{aligned}
$$

In figure 36 the incomplete Beta function is shown for a few $(a, b)$-values. Note that by symmetry the $(1,5)$ and $(5,1)$ curves are reflected around the diagonal. For large values of $a$ and $b$ the curve rises sharply from near zero to near one around $x=a /(a+b)$.

### 42.7.1 Numerical Calculation

In order to obtain $I_{x}(a, b)$ the series expansion

$$
I_{x}(a, b)=\frac{x^{a}(1-x)^{b}}{a B(a, b)}\left[1+\sum_{n=0}^{\infty} \frac{B(a+1, n+1)}{B(a+b, n+1)} x^{n+1}\right]
$$

is not the most useful formula for computations. The continued fraction formula

$$
I_{x}(a, b)=\frac{x^{a}(1-x)^{b}}{a B(a, b)}\left[\frac{1}{1+} \frac{d_{1}}{1+} \frac{d_{2}}{1+} \cdots\right]
$$



Figure 36: The incomplete Beta function
turns out to be a better choice [10]. Here

$$
d_{2 m+1}=-\frac{(a+m)(a+b+m) x}{(a+2 m)(a+2 m+1)} \quad \text { and } \quad d_{2 m}=\frac{m(b-m) x}{(a+2 m-1)(a+2 m)}
$$

and the formula converges rapidly for $x<(a+1) /(a+b+1)$. For other $x$-values the same formula may be used after applying the symmetry relation

$$
I_{x}(a, b)=1-I_{1-x}(b, a)
$$

### 42.7.2 Approximation

For higher values of $a$ and $b$, well already from $a+b>6$, the incomplete Beta function may be approximated by

- For $(a+b+1)(1-x) \leq 0.8$ using an approximation to the chi-square distribution in the variable $\chi^{2}=(a+b-1)(1-x)(3-x)-(1-x)(b-1)$ with $n=2 b$ degrees of freedom.
- For $(a+b+1)(1-x) \geq 0.8$ using an approximation to the standard normal distribution in the variable

$$
z=\frac{3\left[w_{1}\left(1-\frac{1}{9 b}\right)-w_{2}\left(1-\frac{1}{9 a}\right)\right]}{\sqrt{\frac{w_{1}^{2}}{b}+\frac{w_{2}^{2}}{a}}}
$$

where $w_{1}=\sqrt[3]{b x}$ and $w_{2}=\sqrt[3]{a(1-x)}$
In both cases the maximum difference to the true cumulative distribution is below 0.005 all way down to the limit where $a+b=6$ [26].

### 42.8 Relations to Probability Density Functions

The incomplete Gamma and Beta functions, $P(a, x)$ and $I_{x}(a, b)$ are related to many standard probability density functions or rather to their cumulative (distribution) functions. We give very brief examples here. For more details on these distributions consult any book in statistics.

### 42.8.1 The Beta Distribution

The cumulative distribution for the Beta distribution with parameters $p$ and $q$ is given by

$$
F(x)=\frac{1}{B(p, q)} \int_{0}^{x} t^{p-1}(1-t)^{q-1} d t=\frac{B_{x}(p, q)}{B(p, q)}=I_{x}(p, q)
$$

i.e. simply the incomplete Beta function.

### 42.8.2 The Binomial Distribution

For the binomial distribution with parameters $n$ and $p$

$$
\sum_{j=k}^{n}\binom{n}{j} p^{j}(1-p)^{n-j}=I_{p}(k, n-k+1)
$$

i.e. the cumulative distribution may be obtained by

$$
P(k)=\sum_{i=0}^{k}\binom{n}{i} p^{i}(1-p)^{n-i}=I_{1-p}(n-k, k+1)
$$

However, be careful to evaluate $P(n)$, which obviously is unity, using the incomplete Beta function since this is not defined for arguments which are less or equal to zero.

### 42.8.3 The Chi-squared Distribution

The cumulative chi-squared distribution for $n$ degrees of freedom is given by

$$
\begin{aligned}
F(x) & =\frac{1}{2 \Gamma\left(\frac{n}{2}\right)} \int_{0}^{x}\left(\frac{x}{2}\right)^{\frac{n}{2}-1} e^{-\frac{x}{2}} d x=\frac{1}{2 \Gamma\left(\frac{n}{2}\right)} \int_{0}^{\frac{x}{2}} y^{\frac{n}{2}-1} e^{-y} 2 d y= \\
& =\frac{\gamma\left(\frac{n}{2}, \frac{x}{2}\right)}{\Gamma\left(\frac{n}{2}\right)}=P\left(\frac{n}{2}, \frac{x}{2}\right)
\end{aligned}
$$

where $x$ is the chi-squared value sometimes denoted $\chi^{2}$. In this calculation we made the simple substitution $y=x / 2$ in simplifying the integral.

### 42.8.4 The $F$-distribution

The cumulative $F$-distribution with $m$ and $n$ degrees of freedom is given by

$$
\begin{aligned}
F(x) & =\frac{1}{B\left(\frac{m}{2}, \frac{n}{2}\right)} \int_{0}^{x} \frac{m^{\frac{m}{2}} n^{\frac{n}{2}} F^{\frac{m}{2}-1}}{(m F+n)^{\frac{m+n}{2}}} d F=\frac{1}{B\left(\frac{m}{2}, \frac{n}{2}\right)} \int_{0}^{x}\left(\frac{m F}{m F+n}\right)^{\frac{m}{2}}\left(\frac{n}{m F+n}\right)^{\frac{n}{2}} \frac{d F}{F}= \\
& =\frac{1}{B\left(\frac{m}{2}, \frac{n}{2}\right)} \int_{0}^{\frac{m x}{m x+n}} y^{\frac{m}{2}}(1-y)^{\frac{n}{2}} \frac{d y}{y(1-y)}=\frac{1}{B\left(\frac{m}{2}, \frac{n}{2}\right)} \int_{0}^{\frac{m x}{m x+n}} y^{\frac{m}{2}-1}(1-y)^{\frac{n}{2}-1} d y= \\
& =\frac{B_{z}\left(\frac{m}{2}, \frac{n}{2}\right)}{B\left(\frac{m}{2}, \frac{n}{2}\right)}=I_{z}\left(\frac{m}{2}, \frac{n}{2}\right)
\end{aligned}
$$

with $z=m x /(n+m x)$. Here we have made the substitution $y=m F /(m F+n)$, leading to $d F / F=d y / y(1-y)$, in simplifying the integral.

### 42.8.5 The Gamma Distribution

Not surprisingly the cumulative distribution for the Gamma distribution with parameters $a$ and $b$ is given by an incomplete Gamma function.

$$
\begin{aligned}
F(x) & =\int_{0}^{x} f(x) d x=\frac{a^{b}}{\Gamma(b)} \int_{0}^{x} u^{b-1} e^{-a u} d u=\frac{a^{b}}{\Gamma(b)} \int_{0}^{a x}\left(\frac{v}{a}\right)^{b-1} e^{-v} \frac{d v}{a}= \\
& =\frac{1}{\Gamma(b)} \int_{0}^{a x} v^{b-1} e^{-v} d v=\frac{\gamma(b, a x)}{\Gamma(b)}=P(b, a x)
\end{aligned}
$$

### 42.8.6 The Negative Binomial Distribution

The negative binomial distribution with parameters $n$ and $p$ is related to the incomplete Beta function via the relation

$$
\sum_{s=a}^{n}\binom{n+s-1}{s} p^{n}(1-p)^{s}=I_{1-p}(a, n)
$$

Also the geometric distribution, a special case of the negative binomial distribution, is connected to the incomplete Beta function, see summary below.

### 42.8.7 The Normal Distribution

The cumulative normal, or Gaussian, distribution is given by ${ }^{13}$

$$
F(x)= \begin{cases}\frac{1}{2}+\frac{1}{2} P\left(\frac{1}{2}, \frac{x^{2}}{2}\right) & \text { if } x \geq 0 \\ \frac{1}{2}-\frac{1}{2} P\left(\frac{1}{2}, \frac{x^{2}}{2}\right) & \text { if } x<0\end{cases}
$$

[^10]where $P\left(\frac{1}{2}, \frac{x^{2}}{2}\right)$ is the incomplete Gamma function occurring as twice the integral of the standard normal curve from 0 to $x$ since
\[

$$
\begin{aligned}
\frac{1}{\sqrt{2 \pi}} \int_{0}^{x} e^{-\frac{t^{2}}{2}} d t & =\frac{1}{2 \sqrt{\pi}} \int_{0}^{\frac{x^{2}}{2}} \frac{e^{-u}}{\sqrt{u}} d u=\frac{1}{2 \Gamma\left(\frac{1}{2}\right)} \int_{0}^{\frac{x^{2}}{2}} u^{-\frac{1}{2}} e^{-u} d u= \\
& =\frac{\gamma\left(\frac{1}{2}, \frac{x^{2}}{2}\right)}{2 \Gamma\left(\frac{1}{2}\right)}=\frac{1}{2} P\left(\frac{1}{2}, \frac{x^{2}}{2}\right)
\end{aligned}
$$
\]

The so called error function may be expressed in terms of the incomplete Gamma function

$$
\operatorname{erf} x=\frac{2}{\sqrt{\pi}} \int_{0}^{x} e^{-t^{2}} d t=P\left(\frac{1}{2}, x^{2}\right)
$$

as is the case for the complementary error function

$$
\operatorname{erfc} x=1-\operatorname{erf} x=\frac{2}{\sqrt{\pi}} \int_{x}^{\infty} e^{-t^{2}} d t=1-P\left(\frac{1}{2}, x^{2}\right)
$$

defined for $x \geq 0$, for $x<0$ use $\operatorname{erf}(-x)=-\operatorname{erf}(x)$ and $\operatorname{erfc}(-x)=1+\operatorname{erf}(x)$. See also section 13.

There are also other series expansions for erf $x$ like

$$
\begin{aligned}
\operatorname{erf} x & =\frac{2}{\sqrt{\pi}}\left[x-\frac{x^{3}}{3 \cdot 1!}+\frac{x^{5}}{5 \cdot 2!}-\frac{x^{7}}{7 \cdot 3!}+\ldots\right]= \\
& =1-\frac{e^{-x^{2}}}{\sqrt{\pi} x}\left[1-\frac{1}{2 x^{2}}+\frac{1 \cdot 3}{\left(2 x^{2}\right)^{2}}-\frac{1 \cdot 3 \cdot 5}{\left(2 x^{2}\right)^{3}}+\ldots\right]
\end{aligned}
$$

### 42.8.8 The Poisson Distribution

Although the Poisson distribution is a probability density function in a discrete variable the cumulative distribution may be expressed in terms of the incomplete Gamma function. The probability for outcomes from zero to $k-1$ inclusive for a Poisson distribution with parameter (mean) $\mu$ is

$$
P_{\mu}(<k)=\sum_{n=0}^{k-1} \frac{\mu^{n} e^{-\mu}}{n!}=1-P(k, \mu) \quad \text { for } \quad k=1,2, \ldots
$$

### 42.8.9 Student's $t$-distribution

The symmetric integral of the $t$-distribution with $n$ degrees of freedom, often denoted $A(t \mid n)$, is given by

$$
A(t \mid n)=\frac{1}{\sqrt{n} B\left(\frac{1}{2}, \frac{n}{2}\right)} \int_{-t}^{t}\left(1+\frac{x^{2}}{n}\right)^{-\frac{n+1}{2}} d x=\frac{2}{\sqrt{n} B\left(\frac{1}{2}, \frac{n}{2}\right)} \int_{0}^{t}\left(\frac{n}{n+x^{2}}\right)^{\frac{n+1}{2}} d x=
$$

$$
\begin{aligned}
& =\frac{2}{\sqrt{n} B\left(\frac{1}{2}, \frac{n}{2}\right)} \int_{0}^{\frac{t^{2}}{n+t^{2}}}(1-y)^{\frac{n+1}{2}} \frac{n}{2}\left(\frac{1}{1-y}\right)^{2} \sqrt{\frac{1-y}{n y}} d y= \\
& =\frac{1}{B\left(\frac{1}{2}, \frac{n}{2}\right)} \int_{0}^{\frac{t^{2}}{n+t^{2}}} y^{-\frac{1}{2}}(1-y)^{\frac{n}{2}-1} d y=\frac{B_{z}\left(\frac{1}{2}, \frac{n}{2}\right)}{B\left(\frac{1}{2}, \frac{n}{2}\right)}=I_{z}\left(\frac{1}{2}, \frac{n}{2}\right)
\end{aligned}
$$

with $z=t^{2} /\left(n+t^{2}\right)$.

### 42.8.10 Summary

The following table summarizes the relations between the cumulative, distribution, functions of some standard probability density functions and the incomplete Gamma and Beta functions.

|  |  |  |  |
| :--- | :--- | :--- | :--- |
| Distribution | Parameters | Cumulative distribution | Range |
| Beta | $p, q$ | $F(x)=I_{x}(p, q)$ | $0 \leq x \leq 1$ |
| Binomial | $n, p$ | $P(k)=I_{1-p}(n-k, k+1)$ | $k=0,1, \ldots, n$ |
| Chi-squared | $n$ | $F(x)=P\left(\frac{n}{2}, \frac{x}{2}\right)$ | $x \geq 0$ |
| F | $m, n$ | $F(x)=I_{\frac{m x}{n+m x}}\left(\frac{m}{2}, \frac{n}{2}\right)$ | $x \geq 0$ |
| Gamma | $a, b$ | $F(x)=P(b, a x)$ | $x \geq 0$ |
| Geometric | $p$ | $P(k)=I_{p}(1, k)$ | $k=1,2, \ldots$ |
| Negative binomial | $n, p$ | $P(k)=I_{p}(n, k+1)$ | $k=0,1, \ldots$ |
| Standard normal |  | $F(x)=\frac{1}{2}-\frac{1}{2} P\left(\frac{1}{2}, \frac{x^{2}}{2}\right)$ | $-\infty<x<0$ |
| Poisson | $F$ | $F(x)=\frac{1}{2}+\frac{1}{2} P\left(\frac{1}{2}, \frac{x^{2}}{2}\right)$ | $0 \leq x<\infty$ |
| Student | $P(k)=1-P(k+1, \mu)$ | $k=0,1, \ldots$ |  |
|  | $F(x)=\frac{1}{2}-\frac{1}{2} I \frac{x^{2}}{n+x^{2}}\left(\frac{1}{2}, \frac{n}{2}\right)$ | $-\infty<x<0$ |  |
|  |  | $F(x)=\frac{1}{2}+\frac{1}{2} I \frac{x^{2}}{n+x^{2}}\left(\frac{1}{2}, \frac{n}{2}\right)$ | $0 \leq x<\infty$ |

## 43 Appendix B: Hypergeometric Functions

### 43.1 Introduction

The hypergeometric and the confluent hypergeometric functions has a central role inasmuch as many standard functions may be expressed in terms of them. This appendix is based on information from $[41,46,47]$ in which much more detailed information on the hypergeometric and confluent hypergeometric function may be found.

### 43.2 Hypergeometric Function

The hypergeometric function, sometimes called Gauss's differential equation, is given by [41,46]

$$
x(1-x) \frac{\partial^{2} f(x)}{\partial x^{2}}+[c-(a+b+1) x] \frac{\partial f(x)}{\partial x}-a b f(x)=0
$$

One solution is

$$
f(x)={ }_{2} F_{1}(a, b, c ; x)=1+\frac{a b}{c} \frac{x}{1!}+\frac{a(a+1) b(b+1)}{c(c+1)} \frac{x^{2}}{2!}+\cdots \quad c \neq 0,-1,-2,-3, \ldots
$$

The range of convergence is $|x|<1$ and $x=1$, for $c>a+b$, and $x=-1$, for $c>a+b-1$. Using the so called Pochhammer symbol

$$
(a)_{n}=a(a+1)(a+2) \cdots(a+n-1)=\frac{(a+n-1)!}{(a-1)!}=\frac{\Gamma(a+n)}{\Gamma(a)}
$$

with $(a)_{0}=1$ this solution may be written ${ }^{14}$ as

$$
{ }_{2} F_{1}(a, b, c ; x)=\sum_{n=0}^{\infty} \frac{(a)_{n}(b)_{n}}{(c)_{n}} \frac{x^{n}}{n!}=\frac{\Gamma(c)}{\Gamma(a) \Gamma(b)} \sum_{n=0}^{\infty} \frac{\Gamma(a+n) \Gamma(b+n)}{\Gamma(c+n)} \frac{x^{n}}{n!}
$$

By symmetry ${ }_{2} F_{1}(a, b, c ; x)={ }_{2} F_{1}(b, a, c ; x)$ and sometimes the indices are dropped and when the risk for confusion is negligible one simply writes $F(a, b, c ; x)$.

Another independent solution to the hypergeometric equation is

$$
f(x)=x^{1-c}{ }_{2} F_{1}(a+1-c, b+1-c, 2-c ; x) \quad c \neq 2,3,4, \ldots
$$

The $n$ :th derivative of the hypergeometric function is given by

$$
\frac{d^{n}}{d x^{n}}{ }_{2} F_{1}(a, b, c ; x)=\frac{(a)_{n}(b)_{n}}{(c)_{n}}{ }_{2} F_{1}(a+n, b+n, c+n ; x)
$$

and

$$
{ }_{2} F_{1}(a, b, c ; x)=(1-x)^{c-a-b}{ }_{2} F_{1}(c-a, c-b, c ; x)
$$

Several common mathematical function may be expressed in terms of the hypergeometric function such as, the incomplete Beta function $B_{x}(a, b)$, the complete elliptical integrals

[^11]$K$ and $E$, the Gegenbauer functions $T_{n}^{\beta}(x)$, the Legendre functions $P_{n}(x), P_{n}^{m}(x)$ and $Q_{\nu}(x)$ (second kind), and the Chebyshev functions $T_{n}(x), U_{n}(x)$ and $V_{n}(x)$
\[

$$
\begin{aligned}
& (1-z)^{-a}={ }_{2} F_{1}(a, b, b ; z) \\
& \ln (1+z)=x \cdot{ }_{2} F_{1}(1,1,2 ;-z) \\
& \arctan z=z \cdot{ }_{2} F_{1}\left(\frac{1}{2}, 1, \frac{3}{2} ;-z^{2}\right) \\
& \arcsin z=z \cdot{ }_{2} F_{1}\left(\frac{1}{2}, \frac{1}{2}, \frac{3}{2} ; z^{2}\right)=z \sqrt{1-z^{2}}{ }_{2} F_{1}\left(1,1, \frac{3}{2} ; z^{2}\right) \\
& B_{x}(a, b)=\frac{x^{a}}{a}{ }_{2} F_{1}(a, 1-b, a+1 ; x) \\
& K=\int_{0}^{\frac{\pi}{2}}\left(1-k^{2} \sin ^{2} \theta\right)^{-\frac{1}{2}} d \theta=\frac{\pi}{2}{ }_{2} F_{1}\left(\frac{1}{2}, \frac{1}{2}, 1 ; k^{2}\right) \\
& E=\int_{0}^{\frac{\pi}{2}}\left(1-k^{2} \sin ^{2} \theta\right)^{\frac{1}{2}} d \theta=\frac{\pi}{2}{ }_{2} F_{1}\left(\frac{1}{2},-\frac{1}{2}, 1 ; k^{2}\right) \\
& T_{n}^{\beta}(x)=\frac{(n+2 \beta)!}{2^{\beta} n!\beta!}{ }_{2} F_{1}\left(-n, n+2 \beta+1,1+\beta ; \frac{1-x}{2}\right) \\
& P_{n}(x)={ }_{2} F_{1}\left(-n, n+1,1 ; \frac{1-x}{2}\right) \\
& P_{n}^{m}(x)=\frac{(n+m)!}{(n-m)!} \frac{\left(1-x^{2}\right)^{\frac{m}{2}}}{2^{m} m!}{ }_{2} F_{1}\left(m-n, m+n+1, m+1 ; \frac{1-x}{2}\right) \\
& P_{2 n}(x)=(-1)^{n} \frac{(2 n-1)!!}{(2 n)!!}{ }_{2} F_{1}\left(-n, n+\frac{1}{2}, \frac{1}{2} ; x^{2}\right) \\
& P_{2 n+1}(x)=(-1)^{n} \frac{(2 n+1)!!}{(2 n)!!} x \cdot{ }_{2} F_{1}\left(-n, n+\frac{3}{2}, \frac{3}{2} ; x^{2}\right) \\
& Q_{\nu}(x)=\frac{\sqrt{\pi} \nu!}{\left(\nu+\frac{1}{2}\right)!(2 x)^{\nu+1}}{ }_{2} F_{1}\left(\frac{\nu+1}{2}, \frac{\nu}{2}+1, \frac{\nu+3}{2} ; \frac{1}{x^{2}}\right) \\
& T_{n}(x)={ }_{2} F_{1}\left(-n, n, \frac{1}{2} ; \frac{1-x}{2}\right) \\
& U_{n}(x)=(n+1) \cdot{ }_{2} F_{1}\left(-n, n+2, \frac{3}{2} ; \frac{1-x}{2}\right) \\
& V_{n}(x)=n \sqrt{1-x^{2}}{ }_{2} F_{1}\left(-n+1, n+1, \frac{3}{2} ; \frac{1-x}{2}\right)
\end{aligned}
$$
\]

for $Q_{\nu}(x)$ the conditions are $|x|>1,|\arg x|<\pi$, and $\nu \neq-1,-2,-3, \ldots$. See [46] for many more similar and additional formulæ.

### 43.3 Confluent Hypergeometric Function

The confluent hypergeometric equation, or Kummer's equation as it is often called, is given by $[41,47]$

$$
x \frac{\partial^{2} f(x)}{\partial x^{2}}+(c-x) \frac{\partial f(x)}{\partial x}-a f(x)=0
$$

One solution to this equation is

$$
f(x)={ }_{1} F_{1}(a, c ; x)=M(a, c ; x)=\sum_{n=0}^{\infty} \frac{(a)_{n}}{(c)_{n}} \frac{x^{n}}{n!} \quad c \neq 0,-1,-2, \ldots
$$

This solution is convergent for all finite real $x$ (or complex $z$ ). Another solution is given by

$$
f(x)=x^{1-c} M(a+1-c, 2-c ; x) \quad c \neq 2,3,4, \ldots
$$

Often a linear combination of the first and second solution is used

$$
U(a, c ; x)=\frac{\pi}{\sin \pi c}\left[\frac{M(a, c ; x)}{(a-c)!(c-1)!}-\frac{x^{1-c} M(a+1-c, 2-c ; x)}{(a-1)!(1-c)!}\right]
$$

The confluent hypergeometric functions $M$ and $U$ may be expressed in integral form as

$$
\begin{aligned}
M(a, c ; x) & =\frac{\Gamma(c)}{\Gamma(a) \Gamma(c-a)} \int_{0}^{1} e^{x t} t^{a-1}(1-t)^{c-a-1} d t \quad \text { Re } c>0, \text { Re } a>0 \\
U(a, c ; x) & =\frac{1}{\Gamma(a)} \int_{0}^{\infty} e^{-x t} t^{a-1}(1+t)^{c-a-1} d t \quad \text { Re } x>0, \text { Re } a>0
\end{aligned}
$$

Useful formulæ are the Kummer transformations

$$
\begin{aligned}
M(a, c ; x) & =e^{x} M(c-a, c ;-x) \\
U(a, c ; x) & =x^{1-c} U(a-c+1,2-c ; x)
\end{aligned}
$$

The $n$ :th derivatives of the confluent hypergeometric functions are given by

$$
\begin{aligned}
\frac{d^{n}}{d z^{n}} M(a, b ; z) & =\frac{(a)_{n}}{(b)_{n}} M(a+n, b+n ; z) \\
\frac{d^{n}}{d z^{n}} U(a, b ; z) & =(-1)^{n}(a)_{n} U(a+n, b+n ; z)
\end{aligned}
$$

Several common mathematical function may be expressed in terms of the hypergeometric function such as the error function, the incomplete Gamma function $\gamma(a, x)$, Bessel functions $J_{\nu}(x)$, modified Bessel functions of the first kind $I_{\nu}(x)$, Hermite functions $H_{n}(x)$, Laguerre functions $L_{n}(x)$, associated Laguerre functions $L_{n}^{m}(x)$, Whittaker functions $M_{k \mu}(x)$ and $W_{k \mu}(x)$, Fresnel integrals $C(x)$ and $S(x)$, modified Bessel function of the second kind $K_{\nu}(x)$

$$
\begin{aligned}
e^{z} & =M(a, a ; z) \\
\operatorname{erf}(x) & =\frac{2}{\sqrt{\pi}} x M\left(\frac{1}{2}, \frac{3}{2} ;-x^{2}\right)=\frac{2}{\sqrt{\pi}} x e^{-x^{2}} M\left(1, \frac{3}{2} ; x^{2}\right) \\
\gamma(a, x) & =\frac{x^{a}}{a} M(a, a+1 ;-x) \quad \text { Re } a>0 \\
J_{\nu}(x) & =\frac{e^{-\imath x}}{\nu!}\left(\frac{x}{2}\right)^{\nu} M\left(\nu+\frac{1}{2}, 2 \nu+1 ; 2 \imath x\right) \\
I_{\nu}(x) & =\frac{e^{-x}}{\nu!}\left(\frac{x}{2}\right)^{\nu} M\left(\nu+\frac{1}{2}, 2 \nu+1 ; 2 x\right) \\
H_{2 n}(x) & =(-1)^{n} \frac{(2 n)!}{n!} M\left(-n, \frac{1}{2} ; x^{2}\right) \\
H_{2 n+1}(x) & =(-1)^{n} \frac{2(2 n+1)!}{n!} x M\left(-n, \frac{3}{2} ; x^{2}\right)
\end{aligned}
$$

$$
\begin{aligned}
L_{n}(x) & =M(-n, 1 ; x) \\
L_{n}^{m}(x) & =(-1)^{m} \frac{\partial^{m}}{\partial x^{m}} L_{n+m}(x)=\frac{(n+m)!}{n!m!} M(-n, m+1 ; x) \\
M_{k \mu}(x) & =e^{-\frac{x}{2}} x^{\mu+\frac{1}{2}} M\left(\mu-k+\frac{1}{2}, 2 \mu+1 ; x\right) \\
W_{k \mu}(x) & =e^{-\frac{x}{2}} x^{\mu+\frac{1}{2}} U\left(\mu-k+\frac{1}{2}, 2 \mu+1 ; x\right) \\
C(x)+\imath S(x) & =x M\left(\frac{1}{x}, \frac{3}{2} ; \frac{\imath \pi x^{2}}{2}\right) \\
K_{\nu}(x) & =\sqrt{\pi} e^{-x}(2 x)^{\nu} U\left(\nu+\frac{1}{2}, 2 \nu+1 ; 2 x\right)
\end{aligned}
$$

See [47] for many more similar and additional formulæ.

Table 1: Percentage points of the chi-square distribution

|  | $1-\alpha$ |  |  |  |  |  |  |  |
| ---: | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |
|  | 0.5000 | 0.8000 | 0.9000 | 0.9500 | 0.9750 | 0.9900 | 0.9950 | 0.9990 |
| 1 | 0.4549 | 1.6424 | 2.7055 | 3.8415 | 5.0239 | 6.6349 | 7.8794 | 10.828 |
| 2 | 1.3863 | 3.2189 | 4.6052 | 5.9915 | 7.3778 | 9.2103 | 10.597 | 13.816 |
| 3 | 2.3660 | 4.6416 | 6.2514 | 7.8147 | 9.3484 | 11.345 | 12.838 | 16.266 |
| 4 | 3.3567 | 5.9886 | 7.7794 | 9.4877 | 11.143 | 13.277 | 14.860 | 18.467 |
| 5 | 4.3515 | 7.2893 | 9.2364 | 11.070 | 12.833 | 15.086 | 16.750 | 20.515 |
| 6 | 5.3481 | 8.5581 | 10.645 | 12.592 | 14.449 | 16.812 | 18.548 | 22.458 |
| 7 | 6.3458 | 9.8032 | 12.017 | 14.067 | 16.013 | 18.475 | 20.278 | 24.322 |
| 8 | 7.3441 | 11.030 | 13.362 | 15.507 | 17.535 | 20.090 | 21.955 | 26.124 |
| 9 | 8.3428 | 12.242 | 14.684 | 16.919 | 19.023 | 21.666 | 23.589 | 27.877 |
| 10 | 9.3418 | 13.442 | 15.987 | 18.307 | 20.483 | 23.209 | 25.188 | 29.588 |
| 11 | 10.341 | 14.631 | 17.275 | 19.675 | 21.920 | 24.725 | 26.757 | 31.264 |
| 12 | 11.340 | 15.812 | 18.549 | 21.026 | 23.337 | 26.217 | 28.300 | 32.909 |
| 13 | 12.340 | 16.985 | 19.812 | 22.362 | 24.736 | 27.688 | 29.819 | 34.528 |
| 14 | 13.339 | 18.151 | 21.064 | 23.685 | 26.119 | 29.141 | 31.319 | 36.123 |
| 15 | 14.339 | 19.311 | 22.307 | 24.996 | 27.488 | 30.578 | 32.801 | 37.697 |
| 16 | 15.338 | 20.465 | 23.542 | 26.296 | 28.845 | 32.000 | 34.267 | 39.252 |
| 17 | 16.338 | 21.615 | 24.769 | 27.587 | 30.191 | 33.409 | 35.718 | 40.790 |
| 18 | 17.338 | 22.760 | 25.989 | 28.869 | 31.526 | 34.805 | 37.156 | 42.312 |
| 19 | 18.338 | 23.900 | 27.204 | 30.144 | 32.852 | 36.191 | 38.582 | 43.820 |
| 20 | 19.337 | 25.038 | 28.412 | 31.410 | 34.170 | 37.566 | 39.997 | 45.315 |
| 21 | 20.337 | 26.171 | 29.615 | 32.671 | 35.479 | 38.932 | 41.401 | 46.797 |
| 22 | 21.337 | 27.301 | 30.813 | 33.924 | 36.781 | 40.289 | 42.796 | 48.268 |
| 23 | 22.337 | 28.429 | 32.007 | 35.172 | 38.076 | 41.638 | 44.181 | 49.728 |
| 24 | 23.337 | 29.553 | 33.196 | 36.415 | 39.364 | 42.980 | 45.559 | 51.179 |
| 25 | 24.337 | 30.675 | 34.382 | 37.652 | 40.646 | 44.314 | 46.928 | 52.620 |
| 26 | 25.336 | 31.795 | 35.563 | 38.885 | 41.923 | 45.642 | 48.290 | 54.052 |
| 27 | 26.336 | 32.912 | 36.741 | 40.113 | 43.195 | 46.963 | 49.645 | 55.476 |
| 28 | 27.336 | 34.027 | 37.916 | 41.337 | 44.461 | 48.278 | 50.993 | 56.892 |
| 29 | 28.336 | 35.139 | 39.087 | 42.557 | 45.722 | 49.588 | 52.336 | 58.301 |
| 30 | 29.336 | 36.250 | 40.256 | 43.773 | 46.979 | 50.892 | 53.672 | 59.703 |
| 40 | 39.335 | 47.269 | 51.805 | 55.758 | 59.342 | 63.691 | 66.766 | 73.402 |
| 50 | 49.335 | 58.164 | 63.167 | 67.505 | 71.420 | 76.154 | 79.490 | 86.661 |
| 60 | 59.335 | 68.972 | 74.397 | 79.082 | 83.298 | 88.379 | 91.952 | 99.607 |
| 70 | 69.334 | 79.715 | 85.527 | 90.531 | 95.023 | 100.43 | 104.21 | 112.32 |
| 80 | 79.334 | 90.405 | 96.578 | 101.88 | 106.63 | 112.33 | 116.32 | 124.84 |
| 90 | 89.334 | 101.05 | 107.57 | 113.15 | 118.14 | 124.12 | 128.30 | 137.21 |
| 100 | 99.334 | 111.67 | 118.50 | 124.34 | 129.56 | 135.81 | 140.17 | 149.45 |
|  |  |  |  |  |  |  |  |  |

Table 2: Extreme confidence levels for the chi-square distribution

|  | Chi-square Confidence Levels (as $\chi^{2}$ values) |  |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| d.f. | 0.1 | 0.01 | $10^{-3}$ | $10^{-4}$ | $10^{-5}$ | $10^{-6}$ | $10^{-7}$ | $10^{-8}$ | $10^{-9}$ | $10^{-10}$ | $10^{-11}$ | $10^{-12}$ |
| 1 | 2.71 | 6.63 | 10.8 | 15.1 | 19.5 | 23.9 | 28.4 | 32.8 | 37.3 | 41.8 | 46.3 | 50.8 |
| 2 | 4.61 | 9.21 | 13.8 | 18.4 | 23.0 | 27.6 | 32.2 | 36.8 | 41.4 | 46.1 | 50.7 | 55.3 |
| 3 | 6.25 | 11.3 | 16.3 | 21.1 | 25.9 | 30.7 | 35.4 | 40.1 | 44.8 | 49.5 | 54.2 | 58.9 |
| 4 | 7.78 | 13.3 | 18.5 | 23.5 | 28.5 | 33.4 | 38.2 | 43.1 | 47.9 | 52.7 | 57.4 | 62.2 |
| 5 | 9.24 | 15.1 | 20.5 | 25.7 | 30.9 | 35.9 | 40.9 | 45.8 | 50.7 | 55.6 | 60.4 | 65.2 |
| 6 | 10.6 | 16.8 | 22.5 | 27.9 | 33.1 | 38.3 | 43.3 | 48.4 | 53.3 | 58.3 | 63.2 | 68.1 |
| 7 | 12.0 | 18.5 | 24.3 | 29.9 | 35.3 | 40.5 | 45.7 | 50.8 | 55.9 | 60.9 | 65.9 | 70.8 |
| 8 | 13.4 | 20.1 | 26.1 | 31.8 | 37.3 | 42.7 | 48.0 | 53.2 | 58.3 | 63.4 | 68.4 | 73.5 |
| 9 | 14.7 | 21.7 | 27.9 | 33.7 | 39.3 | 44.8 | 50.2 | 55.4 | 60.7 | 65.8 | 70.9 | 76.0 |
| 10 | 16.0 | 23.2 | 29.6 | 35.6 | 41.3 | 46.9 | 52.3 | 57.7 | 62.9 | 68.2 | 73.3 | 78.5 |
| 11 | 17.3 | 24.7 | 31.3 | 37.4 | 43.2 | 48.9 | 54.4 | 59.8 | 65.2 | 70.5 | 75.7 | 80.9 |
| 12 | 18.5 | 26.2 | 32.9 | 39.1 | 45.1 | 50.8 | 56.4 | 61.9 | 67.3 | 72.7 | 78.0 | 83.2 |
| 13 | 19.8 | 27.7 | 34.5 | 40.9 | 46.9 | 52.7 | 58.4 | 64.0 | 69.5 | 74.9 | 80.2 | 85.5 |
| 14 | 21.1 | 29.1 | 36.1 | 42.6 | 48.7 | 54.6 | 60.4 | 66.0 | 71.6 | 77.0 | 82.4 | 87.8 |
| 15 | 22.3 | 30.6 | 37.7 | 44.3 | 50.5 | 56.5 | 62.3 | 68.0 | 73.6 | 79.1 | 84.6 | 90.0 |
| 16 | 23.5 | 32.0 | 39.3 | 45.9 | 52.2 | 58.3 | 64.2 | 70.0 | 75.7 | 81.2 | 86.7 | 92.2 |
| 17 | 24.8 | 33.4 | 40.8 | 47.6 | 54.0 | 60.1 | 66.1 | 71.9 | 77.6 | 83.3 | 88.8 | 94.3 |
| 18 | 26.0 | 34.8 | 42.3 | 49.2 | 55.7 | 61.9 | 68.0 | 73.8 | 79.6 | 85.3 | 90.9 | 96.4 |
| 19 | 27.2 | 36.2 | 43.8 | 50.8 | 57.4 | 63.7 | 69.8 | 75.7 | 81.6 | 87.3 | 92.9 | 98.5 |
| 20 | 28.4 | 37.6 | 45.3 | 52.4 | 59.0 | 65.4 | 71.6 | 77.6 | 83.5 | 89.3 | 94.9 | 101 |
| 25 | 34.4 | 44.3 | 52.6 | 60.1 | 67.2 | 73.9 | 80.4 | 86.6 | 92.8 | 98.8 | 105 | 111 |
| 30 | 40.3 | 50.9 | 59.7 | 67.6 | 75.0 | 82.0 | 88.8 | 95.3 | 102 | 108 | 114 | 120 |
| 35 | 46.1 | 57.3 | 66.6 | 74.9 | 82.6 | 89.9 | 97.0 | 104 | 110 | 117 | 123 | 129 |
| 40 | 51.8 | 63.7 | 73.4 | 82.1 | 90.1 | 97.7 | 105 | 112 | 119 | 125 | 132 | 138 |
| 45 | 57.5 | 70.0 | 80.1 | 89.1 | 97.4 | 105 | 113 | 120 | 127 | 134 | 140 | 147 |
| 50 | 63.2 | 76.2 | 86.7 | 96.0 | 105 | 113 | 120 | 128 | 135 | 142 | 149 | 155 |
| 60 | 74.4 | 88.4 | 99.6 | 110 | 119 | 127 | 135 | 143 | 150 | 158 | 165 | 172 |
| 70 | 85.5 | 100 | 112 | 123 | 132 | 141 | 150 | 158 | 166 | 173 | 181 | 188 |
| 80 | 96.6 | 112 | 125 | 136 | 146 | 155 | 164 | 172 | 180 | 188 | 196 | 204 |
| 90 | 108 | 124 | 137 | 149 | 159 | 169 | 178 | 187 | 195 | 203 | 211 | 219 |
| 100 | 118 | 136 | 149 | 161 | 172 | 182 | 192 | 201 | 209 | 218 | 226 | 234 |
| 120 | 140 | 159 | 174 | 186 | 198 | 209 | 219 | 228 | 237 | 246 | 255 | 263 |
| 150 | 173 | 193 | 209 | 223 | 236 | 247 | 258 | 268 | 278 | 288 | 297 | 306 |
| 200 | 226 | 249 | 268 | 283 | 297 | 310 | 322 | 333 | 344 | 355 | 365 | 374 |
| 300 | 332 | 360 | 381 | 400 | 416 | 431 | 445 | 458 | 471 | 483 | 495 | 506 |
| 400 | 437 | 469 | 493 | 514 | 532 | 549 | 565 | 580 | 594 | 607 | 620 | 632 |
| 500 | 541 | 576 | 603 | 626 | 646 | 665 | 682 | 698 | 714 | 728 | 742 | 756 |
| 600 | 645 | 684 | 713 | 737 | 759 | 779 | 798 | 815 | 832 | 847 | 862 | 877 |
| 800 | 852 | 896 | 929 | 957 | 982 | 1005 | 1026 | 1045 | 1064 | 1081 | 1098 | 1114 |
| 1000 | 1058 | 1107 | 1144 | 1175 | 1202 | 1227 | 1250 | 1272 | 1292 | 1311 | 1330 | 1348 |

Table 3: Extreme confidence levels for the chi-square distribution (as $\chi^{2} /$ d.f. values)

|  | Chi-square Confidence Levels (as $\chi^{2} /$ d.f. values) |  |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| d.f. | 0.1 | 0.01 | $10^{-3}$ | $10^{-4}$ | $10^{-5}$ | $10^{-6}$ | $10^{-7}$ | $10^{-8}$ | $10^{-9}$ | $10^{-10}$ | $10^{-11}$ | $10^{-12}$ |
| 1 | 2.71 | 6.63 | 10.83 | 15.14 | 19.51 | 23.93 | 28.37 | 32.84 | 37.32 | 41.82 | 46.33 | 50.84 |
| 2 | 2.30 | 4.61 | 6.91 | 9.21 | 11.51 | 13.82 | 16.12 | 18.42 | 20.72 | 23.03 | 25.33 | 27.63 |
| 3 | 2.08 | 3.78 | 5.42 | 7.04 | 8.63 | 10.22 | 11.80 | 13.38 | 14.95 | 16.51 | 18.08 | 19.64 |
| 4 | 1.94 | 3.32 | 4.62 | 5.88 | 7.12 | 8.34 | 9.56 | 10.77 | 11.97 | 13.17 | 14.36 | 15.55 |
| 5 | 1.85 | 3.02 | 4.10 | 5.15 | 6.17 | 7.18 | 8.17 | 9.16 | 10.14 | 11.11 | 12.08 | 13.05 |
| 6 | 1.77 | 2.80 | 3.74 | 4.64 | 5.52 | 6.38 | 7.22 | 8.06 | 8.89 | 9.72 | 10.54 | 11.35 |
| 7 | 1.72 | 2.64 | 3.47 | 4.27 | 5.04 | 5.79 | 6.53 | 7.26 | 7.98 | 8.70 | 9.41 | 10.12 |
| 8 | 1.67 | 2.51 | 3.27 | 3.98 | 4.67 | 5.34 | 6.00 | 6.65 | 7.29 | 7.92 | 8.56 | 9.18 |
| 9 | 1.63 | 2.41 | 3.10 | 3.75 | 4.37 | 4.98 | 5.57 | 6.16 | 6.74 | 7.31 | 7.88 | 8.45 |
| 10 | 1.60 | 2.32 | 2.96 | 3.56 | 4.13 | 4.69 | 5.23 | 5.77 | 6.29 | 6.82 | 7.33 | 7.85 |
| 11 | 1.57 | 2.25 | 2.84 | 3.40 | 3.93 | 4.44 | 4.94 | 5.44 | 5.92 | 6.41 | 6.88 | 7.35 |
| 12 | 1.55 | 2.18 | 2.74 | 3.26 | 3.76 | 4.24 | 4.70 | 5.16 | 5.61 | 6.06 | 6.50 | 6.93 |
| 13 | 1.52 | 2.13 | 2.66 | 3.14 | 3.61 | 4.06 | 4.49 | 4.92 | 5.34 | 5.76 | 6.17 | 6.58 |
| 14 | 1.50 | 2.08 | 2.58 | 3.04 | 3.48 | 3.90 | 4.31 | 4.72 | 5.11 | 5.50 | 5.89 | 6.27 |
| 15 | 1.49 | 2.04 | 2.51 | 2.95 | 3.37 | 3.77 | 4.16 | 4.54 | 4.91 | 5.28 | 5.64 | 6.00 |
| 16 | 1.47 | 2.00 | 2.45 | 2.87 | 3.27 | 3.65 | 4.01 | 4.37 | 4.73 | 5.08 | 5.42 | 5.76 |
| 17 | 1.46 | 1.97 | 2.40 | 2.80 | 3.17 | 3.54 | 3.89 | 4.23 | 4.57 | 4.90 | 5.22 | 5.55 |
| 18 | 1.44 | 1.93 | 2.35 | 2.73 | 3.09 | 3.44 | 3.78 | 4.10 | 4.42 | 4.74 | 5.05 | 5.36 |
| 19 | 1.43 | 1.90 | 2.31 | 2.67 | 3.02 | 3.35 | 3.67 | 3.99 | 4.29 | 4.59 | 4.89 | 5.18 |
| 20 | 1.42 | 1.88 | 2.27 | 2.62 | 2.95 | 3.27 | 3.58 | 3.88 | 4.17 | 4.46 | 4.75 | 5.03 |
| 25 | 1.38 | 1.77 | 2.10 | 2.41 | 2.69 | 2.96 | 3.21 | 3.47 | 3.71 | 3.95 | 4.19 | 4.42 |
| 30 | 1.34 | 1.70 | 1.99 | 2.25 | 2.50 | 2.73 | 2.96 | 3.18 | 3.39 | 3.60 | 3.80 | 4.00 |
| 35 | 1.32 | 1.64 | 1.90 | 2.14 | 2.36 | 2.57 | 2.77 | 2.96 | 3.15 | 3.34 | 3.52 | 3.69 |
| 40 | 1.30 | 1.59 | 1.84 | 2.05 | 2.25 | 2.44 | 2.62 | 2.80 | 2.97 | 3.13 | 3.29 | 3.45 |
| 45 | 1.28 | 1.55 | 1.78 | 1.98 | 2.16 | 2.34 | 2.50 | 2.66 | 2.82 | 2.97 | 3.12 | 3.26 |
| 50 | 1.26 | 1.52 | 1.73 | 1.92 | 2.09 | 2.25 | 2.41 | 2.55 | 2.70 | 2.84 | 2.97 | 3.11 |
| 60 | 1.24 | 1.47 | 1.66 | 1.83 | 1.98 | 2.12 | 2.25 | 2.38 | 2.51 | 2.63 | 2.75 | 2.86 |
| 70 | 1.22 | 1.43 | 1.60 | 1.75 | 1.89 | 2.02 | 2.14 | 2.25 | 2.37 | 2.48 | 2.58 | 2.68 |
| 80 | 1.21 | 1.40 | 1.56 | 1.70 | 1.82 | 1.94 | 2.05 | 2.15 | 2.26 | 2.35 | 2.45 | 2.54 |
| 90 | 1.20 | 1.38 | 1.52 | 1.65 | 1.77 | 1.87 | 1.98 | 2.07 | 2.17 | 2.26 | 2.35 | 2.43 |
| 100 | 1.18 | 1.36 | 1.49 | 1.61 | 1.72 | 1.82 | 1.92 | 2.01 | 2.09 | 2.18 | 2.26 | 2.34 |
| 120 | 1.17 | 1.32 | 1.45 | 1.55 | 1.65 | 1.74 | 1.82 | 1.90 | 1.98 | 2.05 | 2.12 | 2.19 |
| 150 | 1.15 | 1.29 | 1.40 | 1.49 | 1.57 | 1.65 | 1.72 | 1.79 | 1.85 | 1.92 | 1.98 | 2.04 |
| 200 | 1.13 | 1.25 | 1.34 | 1.42 | 1.48 | 1.55 | 1.61 | 1.67 | 1.72 | 1.77 | 1.82 | 1.87 |
| 300 | 1.11 | 1.20 | 1.27 | 1.33 | 1.39 | 1.44 | 1.48 | 1.53 | 1.57 | 1.61 | 1.65 | 1.69 |
| 400 | 1.09 | 1.17 | 1.23 | 1.28 | 1.33 | 1.37 | 1.41 | 1.45 | 1.48 | 1.52 | 1.55 | 1.58 |
| 500 | 1.08 | 1.15 | 1.21 | 1.25 | 1.29 | 1.33 | 1.36 | 1.40 | 1.43 | 1.46 | 1.48 | 1.51 |
| 600 | 1.07 | 1.14 | 1.19 | 1.23 | 1.27 | 1.30 | 1.33 | 1.36 | 1.39 | 1.41 | 1.44 | 1.46 |
| 800 | 1.06 | 1.12 | 1.16 | 1.20 | 1.23 | 1.26 | 1.28 | 1.31 | 1.33 | 1.35 | 1.37 | 1.39 |
| 1000 | 1.06 | 1.11 | 1.14 | 1.17 | 1.20 | 1.23 | 1.25 | 1.27 | 1.29 | 1.31 | 1.33 | 1.35 |

Table 4: Exact and approximate values for the Bernoulli numbers


Table 5: Percentage points of the $F$-distribution

| $\alpha=0.10$ |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $m$ | $n$ |  |  |  |  |  |  |  |  |  |
|  | 1 | 2 | 3 | 4 | 5 | 10 | 20 | 50 | 100 | $\infty$ |
| 1 | 39.86 | 49.50 | 53.59 | 55.83 | 57.24 | 60.19 | 61.74 | 62.69 | 63.01 | 63.33 |
| 2 | 8.526 | 9.000 | 9.162 | 9.243 | 9.293 | 9.392 | 9.441 | 9.471 | 9.481 | 9.491 |
| 3 | 5.538 | 5.462 | 5.391 | 5.343 | 5.309 | 5.230 | 5.184 | 5.155 | 5.144 | 5.134 |
| 4 | 4.545 | 4.325 | 4.191 | 4.107 | 4.051 | 3.920 | 3.844 | 3.795 | 3.778 | 3.761 |
| 5 | 4.060 | 3.780 | 3.619 | 3.520 | 3.453 | 3.297 | 3.207 | 3.147 | 3.126 | 3.105 |
| 10 | 3.285 | 2.924 | 2.728 | 2.605 | 2.522 | 2.323 | 2.201 | 2.117 | 2.087 | 2.055 |
| 20 | 2.975 | 2.589 | 2.380 | 2.249 | 2.158 | 1.937 | 1.794 | 1.690 | 1.650 | 1.607 |
| 50 | 2.809 | 2.412 | 2.197 | 2.061 | 1.966 | 1.729 | 1.568 | 1.441 | 1.388 | 1.327 |
| 100 | 2.756 | 2.356 | 2.139 | 2.002 | 1.906 | 1.663 | 1.494 | 1.355 | 1.293 | 1.214 |
| $\infty$ | 2.706 | 2.303 | 2.084 | 1.945 | 1.847 | 1.599 | 1.421 | 1.263 | 1.185 | 1.000 |


| $\alpha=0.05$ |  | $n$ |  |  |  |  |  |  |  |  |  |  |  |
| ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: | :---: | :---: |
|  |  |  |  |  |  |  |  |  |  |  |  |  |  |
|  | 1 | 2 | 3 | 4 | 5 | 10 | 20 | 50 | 100 | $\infty$ |  |  |  |
| 1 | 161.4 | 199.5 | 215.7 | 224.6 | 230.2 | 241.9 | 248.0 | 251.8 | 253.0 | 254.3 |  |  |  |
| 2 | 18.51 | 19.00 | 19.16 | 19.25 | 19.30 | 19.40 | 19.45 | 19.48 | 19.49 | 19.50 |  |  |  |
| 3 | 10.13 | 9.552 | 9.277 | 9.117 | 9.013 | 8.786 | 8.660 | 8.581 | 8.554 | 8.526 |  |  |  |
| 4 | 7.709 | 6.944 | 6.591 | 6.388 | 6.256 | 5.964 | 5.803 | 5.699 | 5.664 | 5.628 |  |  |  |
| 5 | 6.608 | 5.786 | 5.409 | 5.192 | 5.050 | 4.735 | 4.558 | 4.444 | 4.405 | 4.365 |  |  |  |
| 10 | 4.965 | 4.103 | 3.708 | 3.478 | 3.326 | 2.978 | 2.774 | 2.637 | 2.588 | 2.538 |  |  |  |
| 20 | 4.351 | 3.493 | 3.098 | 2.866 | 2.711 | 2.348 | 2.124 | 1.966 | 1.907 | 1.843 |  |  |  |
| 50 | 4.034 | 3.183 | 2.790 | 2.557 | 2.400 | 2.026 | 1.784 | 1.599 | 1.525 | 1.438 |  |  |  |
| 100 | 3.936 | 3.087 | 2.696 | 2.463 | 2.305 | 1.927 | 1.676 | 1.477 | 1.392 | 1.283 |  |  |  |
| $\infty$ | 3.841 | 2.996 | 2.605 | 2.372 | 2.214 | 1.831 | 1.571 | 1.350 | 1.243 | 1.000 |  |  |  |


| $\alpha=0.01$ |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $m$ | $n$ |  |  |  |  |  |  |  |  |  |
|  | 1 | 2 | 3 | 4 | 5 | 10 | 20 | 50 | 100 | $\infty$ |
| 1 | 4052 | 5000 | 5403 | 5625 | 5764 | 6056 | 6209 | 6303 | 6334 | 6366 |
| 2 | 98.50 | 99.00 | 99.17 | 99.25 | 99.30 | 99.40 | 99.45 | 99.48 | 99.49 | 99.50 |
| 3 | 34.12 | 30.82 | 29.46 | 28.71 | 28.24 | 27.23 | 26.69 | 26.35 | 26.24 | 26.13 |
| 4 | 21.20 | 18.00 | 16.69 | 15.98 | 15.52 | 14.55 | 14.02 | 13.69 | 13.58 | 13.46 |
| 5 | 16.26 | 13.27 | 12.06 | 11.39 | 10.97 | 10.05 | 9.553 | 9.238 | 9.130 | 9.020 |
| 10 | 10.04 | 7.559 | 6.552 | 5.994 | 5.636 | 4.849 | 4.405 | 4.115 | 4.014 | 3.909 |
| 20 | 8.096 | 5.849 | 4.938 | 4.431 | 4.103 | 3.368 | 2.938 | 2.643 | 2.535 | 2.421 |
| 50 | 7.171 | 5.057 | 4.199 | 3.720 | 3.408 | 2.698 | 2.265 | 1.949 | 1.825 | 1.683 |
| 100 | 6.895 | 4.824 | 3.984 | 3.513 | 3.206 | 2.503 | 2.067 | 1.735 | 1.598 | 1.427 |
| $\infty$ | 6.635 | 4.605 | 3.782 | 3.319 | 3.017 | 2.321 | 1.878 | 1.523 | 1.358 | 1.000 |

Table 6: Probability content from $-z$ to $z$ of Gauss distribution in $\%$

|  | 0.00 | 0.01 | 0.02 | 0.03 | 0.04 | 0.05 | 0.06 | 0.07 | 0.08 | 0.09 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 0.0 |  | 0.80 | 1.60 | 2.39 | 19 | 3.99 | 4.78 | 5.58 |  | 7.17 |
| 0.1 | 7.97 | 8.76 | 9.55 | 10.34 | 11.13 | 11.92 | 12 | 13.50 | 28 | 15.07 |
| 0.2 | 15.85 | 16.63 | 17.41 | 18.19 | 18.97 | 19.74 | 20. | 21.28 | 5 | 22.82 |
| 0. | 23.5 | 24. | 25.10 | 25.86 | 26.61 | 27.37 | 28.12 | 28 | 29.61 | 30.35 |
| 0. | 31.08 | 31.82 | 32.55 | 33.28 | 34.01 | 34.73 | 35.45 | 36.16 | 36.88 | 59 |
| 0.5 | 38 | 38. | 39.69 | 40.39 | 08 | 41. | 42 | 43 | 43.81 | 44.48 |
| 0.6 | 45. | 45.81 | 46.47 | 47.13 | 47.78 | 48.43 | 49.0 | 49.7 | 50.35 | 50.98 |
| 0.7 | 51.61 | 52.23 | 52.85 | 53.46 | 54.07 | 54.67 | 55.27 | 55.8 | 56.46 | 57.05 |
| 0.8 | 57.63 | 58.21 | 58.78 | 59.35 | 59.91 | 60.47 | 61.02 | 61.57 | 62.11 | 62.65 |
| 0.9 | 63.19 | 63.72 | 64.24 | 64.76 | 65.28 | 65.79 | 66.29 | 66.80 | 67.29 | 67.78 |
| 1.0 | 68.27 | 68.75 | 69.23 | 69.70 | 70.17 | 70.63 | 71.09 | 71.54 | 71.99 | 72.43 |
| 1.1 | 72.87 | 73.30 | 73.73 | 74.15 | 74.57 | 74.99 | 75.40 | 75.80 | 76.20 | 76.60 |
| 1.2 | 76.99 | 77.37 | 77.75 | 78.13 | 78.50 | 78.87 | 79.23 | 79.59 | 79.95 | 29 |
| 1. | 80.64 | 80.98 | 81.32 | 81.65 | 81.98 | 82.30 | 82.6 | 82.93 | 83.24 | 55 |
| 1.4 | 83. | 84. | 84. | 84 | 85 | 85 | 85 | 85.84 | 86.11 | 86.38 |
| 1.5 | 86 | 86. | 87 | 87 | 87 | 87 | 88.12 | 88 | 88 |  |
| 1.6 | 89.0 | 89.26 | 89.48 | 89.69 | 89.90 | 90.1 | 90.3 | 90.51 | 90.70 | 90.90 |
| 1.7 | 91.09 | 91.27 | 91.46 | 91.64 | 91.81 | 91.99 | 92.16 | 92.33 | 92.49 | 92.65 |
| 1. | 92.81 | 92.97 | 93.12 | 93.27 | 93.42 | 93.57 | 93.71 | 93.8 | 93.99 | 94.12 |
| 1.9 | 94.26 | 94.39 | 94.51 | 94.64 | 94.76 | 94.88 | 95.00 | 95.12 | 95.23 | 95.34 |
| 2.0 | 95.45 | 95.56 | 95.66 | 95.76 | 95.86 | 95.96 | 96.0 | 96.1 | 96.2 | 96.34 |
| 2. | 96.4 | 96.5 | 96.60 | 96.68 | 96.76 | 96.8 | 96.9 | 97.0 | 97.0 | 97.15 |
| 2.2 | 97.22 | 97.29 | 97.36 | 97.4 | 97.49 | 97.5 | 97.6 | 97. | 97. | 97.80 |
| 2. | 97.8 | 97.9 | 97.9 | 98.02 | 98.0 | 98. | 98 | 98 | 98.2 | 98.32 |
| 2. | 98. | 98. | 98 | 98 | 98.53 | 98 | 98. | 98. | 98.69 | 98.72 |
| 2. | 98.7 | 98.79 | 98.8 | 98.86 | 98.89 | 98 | 98 | 98 | 99. | 99 |
| 2. | 99.07 | 99.09 | 99.12 | 99.15 | 99.17 | 99.20 | 99.22 | 99.2 | 99.26 | 99.29 |
| 2.7 | 99.31 | 99.33 | 99.35 | 99.37 | 99.39 | 99.40 | 99.42 | 99.4 | 99.46 | 99.47 |
| 2.8 | 99.49 | 99.50 | 99.52 | 99.53 | 99.55 | 99.56 | 99.58 | 99.59 | 99.60 | 99.61 |
| 2.9 | 99.63 | 99.64 | 99.65 | 99.66 | 99.67 | 99.68 | 99.69 | 99.70 | 99.71 | 99.72 |
| 3. | 99.73 | 99.74 | 99.75 | 99.76 | 99.76 | 99.77 | 99.78 | 99.7 | 99.7 | 99.80 |
| 3.1 | 99.8 | 99.8 | 99.82 | 99.83 | 99.83 | 99.8 | 99.8 | 99.8 | 99. | 99.86 |
| 3. | 99.8 | 99.8 | 99.8 | 99.88 | 99.8 | 99.8 | 99. | 99. | 99. | 99.90 |
| 3. | 99.9 | 99.9 | 99.91 | 99.91 | 99.92 | 99.92 | 99.9 | 99.92 | 99.93 | 99.93 |
| 3. | 99.93 | 99.94 | 99.94 | 99.94 | 99.94 | 99.9 | 99.95 | 99.95 | 99.95 | 99. |
| 3.5 | 99.95 | 99.96 | 99.96 | 99.96 | 99.96 | 99.96 | 99.96 | 99.96 | 99.97 | 99.97 |
| 3.6 | 99.97 | 99.97 | 99.97 | 99.97 | 99.97 | 99.97 | 99.97 | 99.98 | 99.98 | 99.98 |
| 3.7 | 99.98 | 99.98 | 99.98 | 99.98 | 99.98 | 99.98 | 99.98 | 99.98 | 99.98 | 99.98 |
| 3.8 | 99.99 | 99.99 | 99.99 | 99.99 | 99.99 | 99.99 | 99.99 | 99.99 | 99.99 | 99.99 |
| 3.9 | 99.99 | 99.99 | 99.99 | 99.99 | 99.99 | 99.99 | 99.99 | 99.99 | 99.99 | 99.99 |

Table 7: Standard normal distribution $z$-values for a specific probability content from $-z$ to $z$. Read column-wise and add marginal column and row $z$. Read column-wise and add marginal column and row figures to find probabilities.

| Prob. | 0.00 | 0.10 | 0.20 | 0.30 | 0.40 | 0.50 | 0.60 | 0.70 | 0.80 | 0.90 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 0.000 | 0.000 | 0.125 | 0.253 | 0.385 | 0.524 | 0.674 | 0.841 | 1.036 | 1.282 | 1.645 |
| 0.002 | 0.002 | 0.128 | 0.256 | 0.388 | 0.527 | 0.677 | 0.845 | 1.041 | 1.287 | 1.655 |
| 0.004 | 0.005 | 0.130 | 0.258 | 0.390 | 0.530 | 0.681 | 0.849 | 1.045 | 1.293 | 1.665 |
| 0.006 | 0.007 | 0.133 | 0.261 | 0.393 | 0.533 | 0.684 | 0.852 | 1.049 | 1.299 | 1.675 |
| 0.008 | 0.010 | 0.135 | 0.263 | 0.396 | 0.536 | 0.687 | 0.856 | 1.054 | 1.305 | 1.685 |
| 0.010 | 0.012 | 0.138 | 0.266 | 0.398 | 0.538 | 0.690 | 0.859 | 1.058 | 1.311 | 1.696 |
| 0.012 | 0.015 | 0.141 | 0.268 | 0.401 | 0.541 | 0.693 | 0.863 | 1.063 | 1.317 | 1.706 |
| 0.014 | 0.017 | 0.143 | 0.271 | 0.404 | 0.544 | 0.696 | 0.867 | 1.067 | 1.323 | 1.717 |
| 0.016 | 0.020 | 0.146 | 0.274 | 0.407 | 0.547 | 0.700 | 0.870 | 1.071 | 1.329 | 1.728 |
| 0.018 | 0.022 | 0.148 | 0.276 | 0.409 | 0.550 | 0.703 | 0.874 | 1.076 | 1.335 | 1.740 |
| 0.020 | 0.025 | 0.151 | 0.279 | 0.412 | 0.553 | 0.706 | 0.878 | 1.080 | 1.341 | 1.751 |
| 0.022 | 0.027 | 0.153 | 0.281 | 0.415 | 0.556 | 0.709 | 0.881 | 1.085 | 1.347 | 1.763 |
| 0.024 | 0.030 | 0.156 | 0.284 | 0.417 | 0.559 | 0.712 | 0.885 | 1.089 | 1.353 | 1.775 |
| 0.026 | 0.033 | 0.158 | 0.287 | 0.420 | 0.562 | 0.716 | 0.889 | 1.094 | 1.360 | 1.787 |
| 0.028 | 0.035 | 0.161 | 0.289 | 0.423 | 0.565 | 0.719 | 0.893 | 1.099 | 1.366 | 1.800 |
| 0.030 | 0.038 | 0.163 | 0.292 | 0.426 | 0.568 | 0.722 | 0.896 | 1.103 | 1.372 | 1.812 |
| 0.032 | 0.040 | 0.166 | 0.295 | 0.428 | 0.571 | 0.725 | 0.900 | 1.108 | 1.379 | 1.825 |
| 0.034 | 0.043 | 0.168 | 0.297 | 0.431 | 0.574 | 0.729 | 0.904 | 1.112 | 1.385 | 1.839 |
| 0.036 | 0.045 | 0.171 | 0.300 | 0.434 | 0.577 | 0.732 | 0.908 | 1.117 | 1.392 | 1.853 |
| 0.038 | 0.048 | 0.173 | 0.302 | 0.437 | 0.580 | 0.735 | 0.911 | 1.122 | 1.399 | 1.867 |
| 0.040 | 0.050 | 0.176 | 0.305 | 0.439 | 0.582 | 0.739 | 0.915 | 1.126 | 1.405 | 1.881 |
| 0.042 | 0.053 | 0.179 | 0.308 | 0.442 | 0.585 | 0.742 | 0.919 | 1.131 | 1.412 | 1.896 |
| 0.044 | 0.055 | 0.181 | 0.310 | 0.445 | 0.588 | 0.745 | 0.923 | 1.136 | 1.419 | 1.911 |
| 0.046 | 0.058 | 0.184 | 0.313 | 0.448 | 0.591 | 0.749 | 0.927 | 1.141 | 1.426 | 1.927 |
| 0.048 | 0.060 | 0.186 | 0.316 | 0.451 | 0.594 | 0.752 | 0.931 | 1.146 | 1.433 | 1.944 |
| 0.050 | 0.063 | 0.189 | 0.318 | 0.453 | 0.597 | 0.755 | 0.935 | 1.150 | 1.440 | 1.960 |
| 0.052 | 0.065 | 0.191 | 0.321 | 0.456 | 0.600 | 0.759 | 0.938 | 1.155 | 1.447 | 1.978 |
| 0.054 | 0.068 | 0.194 | 0.323 | 0.459 | 0.603 | 0.762 | 0.942 | 1.160 | 1.454 | 1.996 |
| 0.056 | 0.070 | 0.196 | 0.326 | 0.462 | 0.606 | 0.765 | 0.946 | 1.165 | 1.461 | 2.015 |
| 0.058 | 0.073 | 0.199 | 0.329 | 0.464 | 0.609 | 0.769 | 0.950 | 1.170 | 1.469 | 2.034 |
| 0.060 | 0.075 | 0.202 | 0.331 | 0.467 | 0.612 | 0.772 | 0.954 | 1.175 | 1.476 | 2.054 |
| 0.062 | 0.078 | 0.204 | 0.334 | 0.470 | 0.615 | 0.775 | 0.958 | 1.180 | 1.484 | 2.075 |
| 0.064 | 0.080 | 0.207 | 0.337 | 0.473 | 0.619 | 0.779 | 0.962 | 1.185 | 1.491 | 2.097 |
| 0.066 | 0.083 | 0.209 | 0.339 | 0.476 | 0.622 | 0.782 | 0.966 | 1.190 | 1.499 | 2.121 |
| 0.068 | 0.085 | 0.212 | 0.342 | 0.478 | 0.625 | 0.786 | 0.970 | 1.195 | 1.507 | 2.145 |
| 0.070 | 0.088 | 0.214 | 0.345 | 0.481 | 0.628 | 0.789 | 0.974 | 1.200 | 1.514 | 2.171 |
| 0.072 | 0.090 | 0.217 | 0.347 | 0.484 | 0.631 | 0.792 | 0.978 | 1.206 | 1.522 | 2.198 |
| 0.074 | 0.093 | 0.219 | 0.350 | 0.487 | 0.634 | 0.796 | 0.982 | 1.211 | 1.530 | 2.227 |
| 0.076 | 0.095 | 0.222 | 0.353 | 0.490 | 0.637 | 0.799 | 0.986 | 1.216 | 1.539 | 2.258 |
| 0.078 | 0.098 | 0.225 | 0.355 | 0.493 | 0.640 | 0.803 | 0.990 | 1.221 | 1.547 | 2.291 |
| 0.080 | 0.100 | 0.227 | 0.358 | 0.495 | 0.643 | 0.806 | 0.994 | 1.227 | 1.555 | 2.327 |
| 0.082 | 0.103 | 0.230 | 0.361 | 0.498 | 0.646 | 0.810 | 0.999 | 1.232 | 1.564 | 2.366 |
| 0.084 | 0.105 | 0.232 | 0.363 | 0.501 | 0.649 | 0.813 | 1.003 | 1.237 | 1.572 | 2.409 |
| 0.086 | 0.108 | 0.235 | 0.366 | 0.504 | 0.652 | 0.817 | 1.007 | 1.243 | 1.581 | 2.458 |
| 0.088 | 0.110 | 0.237 | 0.369 | 0.507 | 0.655 | 0.820 | 1.011 | 1.248 | 1.590 | 2.513 |
| 0.090 | 0.113 | 0.240 | 0.371 | 0.510 | 0.659 | 0.824 | 1.015 | 1.254 | 1.599 | 2.576 |
| 0.092 | 0.115 | 0.243 | 0.374 | 0.513 | 0.662 | 0.827 | 1.019 | 1.259 | 1.608 | 2.652 |
| 0.094 | 0.118 | 0.245 | 0.377 | 0.515 | 0.665 | 0.831 | 1.024 | 1.265 | 1.617 | 2.748 |
| 0.096 | 0.120 | 0.248 | 0.379 | 0.518 | 0.668 | 0.834 | 1.028 | 1.270 | 1.626 | 2.879 |
| 0.098 | 0.123 | 0.250 | 0.382 | 0.521 | 0.671 | 0.838 | 1.032 | 1.276 | 1.636 | 3.091 |

Table 8: Percentage points of the $t$-distribution

|  | $1-\alpha$ |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $n$ | 0.60 | 0.70 | 0.80 | 0.90 | 0.95 | 0.975 | 0.990 | 0.995 | 0.999 | 0.9995 |
| 1 | 0.325 | 0.727 | 1.376 | 3.078 | 6.314 | 12.71 | 31.82 | 63.66 | 318.3 | 636.6 |
| 2 | 0.289 | 0.617 | 1.061 | 1.886 | 2.920 | 4.303 | 6.965 | 9.925 | 22.33 | 31.60 |
| 3 | 0.277 | 0.584 | 0.978 | 1.638 | 2.353 | 3.182 | 4.541 | 5.841 | 10.21 | 12.92 |
| 4 | 0.271 | 0.569 | 0.941 | 1.533 | 2.132 | 2.776 | 3.747 | 4.604 | 7.173 | 8.610 |
| 5 | 0.267 | 0.559 | 0.920 | 1.476 | 2.015 | 2.571 | 3.365 | 4.032 | 5.893 | 6.869 |
| 6 | 0.265 | 0.553 | 0.906 | 1.440 | 1.943 | 2.447 | 3.143 | 3.707 | 5.208 | 5.959 |
| 7 | 0.263 | 0.549 | 0.896 | 1.415 | 1.895 | 2.365 | 2.998 | 3.499 | 4.785 | 5.408 |
| 8 | 0.262 | 0.546 | 0.889 | 1.397 | 1.860 | 2.306 | 2.896 | 3.355 | 4.501 | 5.041 |
| 9 | 0.261 | 0.543 | 0.883 | 1.383 | 1.833 | 2.262 | 2.821 | 3.250 | 4.297 | 4.781 |
| 10 | 0.260 | 0.542 | 0.879 | 1.372 | 1.812 | 2.228 | 2.764 | 3.169 | 4.144 | 4.587 |
| 11 | 0.260 | 0.540 | 0.876 | 1.363 | 1.796 | 2.201 | 2.718 | 3.106 | 4.025 | 4.437 |
| 12 | 0.259 | 0.539 | 0.873 | 1.356 | 1.782 | 2.179 | 2.681 | 3.055 | 3.930 | 4.318 |
| 13 | 0.259 | 0.538 | 0.870 | 1.350 | 1.771 | 2.160 | 2.650 | 3.012 | 3.852 | 4.221 |
| 14 | 0.258 | 0.537 | 0.868 | 1.345 | 1.761 | 2.145 | 2.624 | 2.977 | 3.787 | 4.140 |
| 15 | 0.258 | 0.536 | 0.866 | 1.341 | 1.753 | 2.131 | 2.602 | 2.947 | 3.733 | 4.073 |
| 16 | 0.258 | 0.535 | 0.865 | 1.337 | 1.746 | 2.120 | 2.583 | 2.921 | 3.686 | 4.015 |
| 17 | 0.257 | 0.534 | 0.863 | 1.333 | 1.740 | 2.110 | 2.567 | 2.898 | 3.646 | 3.965 |
| 18 | 0.257 | 0.534 | 0.862 | 1.330 | 1.734 | 2.101 | 2.552 | 2.878 | 3.610 | 3.922 |
| 19 | 0.257 | 0.533 | 0.861 | 1.328 | 1.729 | 2.093 | 2.539 | 2.861 | 3.579 | 3.883 |
| 20 | 0.257 | 0.533 | 0.860 | 1.325 | 1.725 | 2.086 | 2.528 | 2.845 | 3.552 | 3.850 |
| 21 | 0.257 | 0.532 | 0.859 | 1.323 | 1.721 | 2.080 | 2.518 | 2.831 | 3.527 | 3.819 |
| 22 | 0.256 | 0.532 | 0.858 | 1.321 | 1.717 | 2.074 | 2.508 | 2.819 | 3.505 | 3.792 |
| 23 | 0.256 | 0.532 | 0.858 | 1.319 | 1.714 | 2.069 | 2.500 | 2.807 | 3.485 | 3.768 |
| 24 | 0.256 | 0.531 | 0.857 | 1.318 | 1.711 | 2.064 | 2.492 | 2.797 | 3.467 | 3.745 |
| 25 | 0.256 | 0.531 | 0.856 | 1.316 | 1.708 | 2.060 | 2.485 | 2.787 | 3.450 | 3.725 |
| 26 | 0.256 | 0.531 | 0.856 | 1.315 | 1.706 | 2.056 | 2.479 | 2.779 | 3.435 | 3.707 |
| 27 | 0.256 | 0.531 | 0.855 | 1.314 | 1.703 | 2.052 | 2.473 | 2.771 | 3.421 | 3.690 |
| 28 | 0.256 | 0.530 | 0.855 | 1.313 | 1.701 | 2.048 | 2.467 | 2.763 | 3.408 | 3.674 |
| 29 | 0.256 | 0.530 | 0.854 | 1.311 | 1.699 | 2.045 | 2.462 | 2.756 | 3.396 | 3.659 |
| 30 | 0.256 | 0.530 | 0.854 | 1.310 | 1.697 | 2.042 | 2.457 | 2.750 | 3.385 | 3.646 |
| 40 | 0.255 | 0.529 | 0.851 | 1.303 | 1.684 | 2.021 | 2.423 | 2.704 | 3.307 | 3.551 |
| 50 | 0.255 | 0.528 | 0.849 | 1.299 | 1.676 | 2.009 | 2.403 | 2.678 | 3.261 | 3.496 |
| 60 | 0.254 | 0.527 | 0.848 | 1.296 | 1.671 | 2.000 | 2.390 | 2.660 | 3.232 | 3.460 |
| 70 | 0.254 | 0.527 | 0.847 | 1.294 | 1.667 | 1.994 | 2.381 | 2.648 | 3.211 | 3.435 |
| 80 | 0.254 | 0.526 | 0.846 | 1.292 | 1.664 | 1.990 | 2.374 | 2.639 | 3.195 | 3.416 |
| 90 | 0.254 | 0.526 | 0.846 | 1.291 | 1.662 | 1.987 | 2.368 | 2.632 | 3.183 | 3.402 |
| 100 | 0.254 | 0.526 | 0.845 | 1.290 | 1.660 | 1.984 | 2.364 | 2.626 | 3.174 | 3.390 |
| 110 | 0.254 | 0.526 | 0.845 | 1.289 | 1.659 | 1.982 | 2.361 | 2.621 | 3.166 | 3.381 |
| 120 | 0.254 | 0.526 | 0.845 | 1.289 | 1.658 | 1.980 | 2.358 | 2.617 | 3.160 | 3.373 |
| $\infty$ | 0.253 | 0.524 | 0.842 | 1.282 | 1.645 | 1.960 | 2.326 | 2.576 | 3.090 | 3.291 |

Table 9: Expressions for the Beta function $B(m, n)$ for integer and half-integer arguments

| $\begin{gathered} n \rightarrow \\ m \downarrow \end{gathered}$ | ${ }^{\frac{1}{2}}$ | 1 | $\frac{3}{2}$ | 2 | $\frac{5}{2}$ | 3 | $\frac{7}{2}$ | 4 | $\frac{9}{2}$ | 5 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\frac{1}{2}$ | $\pi$ |  |  |  |  |  |  |  |  |  |
| 1 | 2 | 1 |  |  |  |  |  |  |  |  |
| $\frac{3}{2}$ | $\frac{1}{2} \pi$ | $\frac{2}{3}$ | $\frac{1}{8} \pi$ |  |  |  |  |  |  |  |
| 2 | ${ }^{\frac{4}{3}}$ | $\frac{1}{2}$ | $\frac{4}{15}$ | $\frac{1}{6}$ |  |  |  |  |  |  |
| $\frac{5}{2}$ | $\frac{3}{8} \pi$ | $\frac{2}{5}$ | $\frac{1}{16} \pi$ | $\frac{4}{35}$ | $\frac{3}{128} \pi$ |  |  |  |  |  |
| 3 | ${ }^{\frac{16}{15}}$ | $\frac{1}{3}$ | $\frac{16}{105}$ | $\frac{1}{12}$ | $\frac{16}{315}$ | $\frac{1}{30}$ |  |  |  |  |
| ${ }^{\frac{7}{2}}$ | $\frac{5}{16} \pi$ | $\frac{2}{7}$ | $\frac{5}{128} \pi$ | ${ }^{\frac{4}{63}}$ | $\frac{3}{256} \pi$ | ${ }^{\frac{16}{693}}$ | $\frac{5}{1024} \pi$ |  |  |  |
| 4 | $\frac{32}{35}$ | $\frac{1}{4}$ | $\frac{32}{315}$ | $\frac{1}{20}$ | $\frac{32}{1155}$ | $\frac{1}{60}$ | $\frac{32}{3003}$ | $\frac{1}{140}$ |  |  |
| $\frac{9}{2}$ | $\frac{35}{128} \pi$ | $\frac{2}{9}$ | $\frac{7}{256} \pi$ | ${ }^{\frac{4}{99}}$ | $\frac{7}{1024} \pi$ | $\frac{16}{1287}$ | $\frac{5}{2048} \pi$ | $\frac{32}{6435}$ | $\frac{35}{32768} \pi$ |  |
| 5 | ${ }^{256}$ | $\frac{1}{5}$ | ${ }^{2566}$ | $\frac{1}{30}$ | $\frac{256}{15015}$ | $\frac{1}{105}$ | $\frac{256}{45045}$ | $\frac{1}{280}$ | ${ }_{\text {256 }}^{103955}$ | $\frac{1}{630}$ |
| $\frac{11}{2}$ | $\frac{63}{256} \pi$ | $\frac{2}{11}$ | $\frac{21}{1024} \pi$ | $\frac{4}{143}$ | $\frac{9}{2048} \pi$ | $\frac{16}{2145}$ | $\frac{45}{32768} \pi$ | $\frac{32}{12155}$ | $\frac{35}{65536} \pi$ | $\frac{256}{230945}$ |
| 6 | $\frac{512}{693}$ | $\frac{1}{6}$ | $\frac{512}{9009}$ | $\frac{1}{42}$ | $\frac{512}{45045}$ | $\frac{1}{168}$ | ${ }^{\frac{512}{153153}}$ | $\frac{1}{504}$ | $\frac{512}{415701}$ | $\frac{1}{1260}$ |
| $\frac{13}{2}$ | ${ }^{231} 1024$ | $\frac{2}{13}$ | $\frac{33}{2048} \pi$ | $\frac{4}{195}$ | $\frac{996}{32768}$ | $\frac{16}{3315}$ | ${ }_{6}^{55536} \pi$ | $\frac{32}{20995}$ | $\frac{77}{262144} \pi$ | $\frac{256}{408895}$ |
| 7 | ${ }^{2048} 3003$ | $\frac{1}{7}$ | $\frac{2048}{45045}$ | $\frac{1}{56}$ | ${ }^{2048} \mathbf{2 5 2 5 5}$ | $\frac{1}{252}$ | ${ }^{20498969}$ | $\frac{1}{840}$ | ${ }^{\frac{2048}{290907}}$ | $\frac{1}{2310}$ |
| $\frac{15}{2}$ | $\frac{429}{2048} \pi$ | $\frac{2}{15}$ | $\frac{429}{32768} \pi$ | $\frac{4}{255}$ | $\frac{143}{65536}$ | $\frac{16}{4845}$ | $\frac{143}{262144} \pi$ | $\frac{32}{33915}$ | $\frac{91}{524288} \pi$ | $\frac{256}{78045}$ |
| $\begin{aligned} & n \rightarrow \\ & m \downarrow \\ & \hline \end{aligned}$ | $\frac{11}{2}$ |  | 6 | $\frac{13}{2}$ | 7 |  | $\frac{15}{2}$ |  |  |  |
| $\frac{11}{2}$ | $\frac{63}{262144} \pi$ |  |  |  |  |  |  |  |  |  |
| 6 | $\frac{512}{969969}$ |  | $\frac{1}{2772}$ |  |  |  |  |  |  |  |
| $\frac{13}{2}$ | $\frac{63}{524288} \pi$ |  | $\frac{512}{2028117}$ | $\frac{231}{4193304} \pi$ |  |  |  |  |  |  |
| 7 | $\frac{2048}{7436429}$ |  | $\frac{1}{5544}$ | $\frac{2048}{1690975}$ | $\frac{1}{12012}$ |  |  |  |  |  |
| $\frac{15}{2}$ | $\frac{273}{4194304} \pi$ |  | $\frac{512}{390225}$ | $\frac{231}{8388600} \pi$ | $\frac{2048}{35102025}$ |  | ${ }^{335554432}$ \% |  |  |  |

## Mathematical Constants

## Introduction

It is handy to have available the values of different mathematical constants appearing in many expressions in statistical calculations. In this section we list, with high precision, many of those which may be needed. In some cases we give, after the tables, basic expressions which may be nice to recall. Note, however, that this is not full explanations and consult the main text or other sources for details.

## Some Basic Constants

|  | exact |
| :--- | :---: |
| $\pi$ | $3.14159 \quad 265358979323846$ |
| e | 2.71828182845904523536 |
| $\gamma$ | 0.57721566490153286061 |
| $\sqrt{\pi}$ | 1.77245385090551602730 |
| $1 / \sqrt{2 \pi}$ | 0.39894228040143267794 |

$$
\begin{aligned}
& e=\sum_{n=0}^{\infty} \frac{1}{n!}=\lim _{n \rightarrow \infty}\left(1+\frac{1}{n}\right)^{n} \\
& \gamma=\lim _{n \rightarrow \infty}\left(\sum_{k=1}^{n} \frac{1}{k}-\ln n\right)
\end{aligned}
$$

## Gamma Function

|  | exact | approx. |
| :--- | :--- | ---: |
| $\Gamma\left(\frac{1}{2}\right)$ | $\sqrt{\pi}$ | 1.77245385090551602730 |
| $\Gamma\left(\frac{3}{2}\right)$ | $\frac{1}{2} \sqrt{\pi}$ | 0.88622692545275801365 |
| $\Gamma\left(\frac{5}{2}\right)$ | $\frac{3}{4} \sqrt{\pi}$ | 1.32934038817913702047 |
| $\Gamma\left(\frac{7}{2}\right)$ | $\frac{15}{8} \sqrt{\pi}$ | 3.32335097044784255118 |
| $\Gamma\left(\frac{9}{2}\right)$ | $\frac{105}{16} \sqrt{\pi}$ | 11.63172839656744892914 |

$$
\begin{aligned}
\Gamma(z) & =\int_{0}^{\infty} t^{z-1} e^{-t} d t \\
n! & =\Gamma(n+1)=n \Gamma(n) \\
\Gamma\left(n+\frac{1}{2}\right) & =\frac{(2 n-1)!!}{2^{n}} \Gamma\left(\frac{1}{2}\right)= \\
& =\frac{(2 n)!}{2^{2 n} n!} \sqrt{\pi}
\end{aligned}
$$

See further section 42.2 on page 154 and reference [42] for more details.

## Beta Function

For exact expressions for the Beta function for halfinteger and integer values see table 9 on page 179.

$$
\begin{aligned}
B(a, b) & =\frac{\Gamma(a) \Gamma(b)}{\Gamma(a+b)}= \\
& =\int_{0}^{1} x^{a-1}(1-x)^{b-1} d x= \\
& =\int_{0}^{\infty} \frac{x^{a-1}}{(1+x)^{a+b}} d x
\end{aligned}
$$

See further section 42.6 on page 160 .

## Digamma Function

|  | exact | approx. |  |
| :--- | :--- | ---: | :--- |
| $\psi\left(\frac{1}{2}\right)$ | $-\gamma-2 \ln 2$ | -1.96351002602142347944 |  |
| $\psi\left(\frac{3}{2}\right)$ | $\psi\left(\frac{1}{2}\right)+2$ | 0.03648997397857652056 |  |
| $\psi\left(\frac{5}{2}\right)$ | $\psi\left(\frac{1}{2}\right)+\frac{8}{3}$ | 0.70315664064524318723 |  |
| $\psi\left(\frac{7}{2}\right)$ | $\psi\left(\frac{1}{2}\right)+\frac{46}{15}$ | 1.10315664064524318723 |  |
| $\psi\left(\frac{9}{2}\right)$ | $\psi\left(\frac{1}{2}\right)+\frac{352}{105}$ | 1.38887092635952890151 |  |
| $\psi\left(\frac{11}{2}\right)$ | $\psi\left(\frac{9}{2}\right)+\frac{2}{9}$ | 1.61109314858175112373 |  |
| $\psi\left(\frac{13}{2}\right)$ | $\psi\left(\frac{11}{2}\right)+\frac{2}{11}$ | 1.79291133039993294192 |  |
| $\psi\left(\frac{15}{2}\right)$ | $\psi\left(\frac{13}{2}\right)+\frac{2}{13}$ | 1.94675748424608678807 |  |
| $\psi\left(\frac{17}{2}\right)$ | $\psi\left(\frac{15}{2}\right)+\frac{2}{15}$ | 2.08009081759420121402 |  |
| $\psi\left(\frac{19}{2}\right)$ | $\psi\left(\frac{17}{2}\right)+\frac{2}{17}$ | 2.19773787640294953317 |  |
|  |  |  |  |
| $\psi(1)$ | $-\gamma$ | -0.57721566490153286061 |  |
| $\psi(2)$ | $1-\gamma$ | 0.42278433509846713939 |  |
| $\psi(3)$ | $\frac{3}{2}-\gamma$ | 0.92278433509846713939 |  |
| $\psi(4)$ | $\frac{11}{6}-\gamma$ | 1.25611766843180047273 |  |
| $\psi(5)$ | $\frac{25}{12}-\gamma$ | 1.50611766843180047273 |  |
| $\psi(6)$ | $\frac{137}{60}-\gamma$ | 1.70611766843180047273 |  |
| $\psi(7)$ | $\frac{49}{20}-\gamma$ | 1.87278433509846713939 |  |
| $\psi(8)$ | $\frac{363}{140}-\gamma$ | 2.01564147795560999654 |  |
| $\psi(9)$ | $\frac{761}{280}-\gamma$ | 2.14064147795560999654 |  |
| $\psi(10)$ | $\frac{7129}{2520}-\gamma$ | 2.25175258906672110765 |  |

$$
\begin{aligned}
\psi(z) & =\frac{d}{d z} \ln \Gamma(z)=\frac{1}{\Gamma(z)} \frac{d \Gamma(z)}{d z} \\
\psi(z+1) & =\psi(z)+\frac{1}{z} \\
\psi(n) & =-\gamma+\sum_{m=1}^{n-1} \frac{1}{m} \\
\psi\left(n+\frac{1}{2}\right) & =-\gamma-2 \ln 2+2 \sum_{m=1}^{n} \frac{1}{2 m-1}
\end{aligned}
$$

See further section 42.3 on page 156 .

## Polygamma Function

|  | exact | approx. |
| :---: | :---: | ---: |
| $\psi^{(1)}\left(\frac{1}{2}\right)$ | $\pi^{2} / 2$ | 4.93480220054467930942 |
| $\psi^{(2)}\left(\frac{1}{2}\right)$ | $-14 \zeta_{3}$ | -16.82879664423431999560 |
| $\psi^{(3)}\left(\frac{1}{2}\right)$ | $\pi^{4}$ | 97.40909103400243723644 |
| $\psi^{(4)}\left(\frac{1}{2}\right)$ | $-744 \zeta_{5}$ | -771.47424982666722519054 |
| $\psi^{(5)}\left(\frac{1}{2}\right)$ | $8 \pi^{6}$ | 7691.11354860243549624176 |
| $\psi^{(1)}(1)$ | $\zeta_{2}$ | 1.64493406684822643647 |
| $\psi^{(2)}(1)$ | $-2 \zeta_{3}$ | -2.40411380631918857080 |
| $\psi^{(3)}(1)$ | $6 \zeta_{4}$ | 6.49393940226682914910 |
| $\psi^{(4)}(1)$ | $-24 \zeta_{5}$ | -24.88626612344087823195 |
| $\psi^{(5)}(1)$ | $120 \zeta_{6}$ | 122.08116743813389676574 |

$$
\begin{aligned}
\psi^{(n)}(z)= & \frac{d^{n}}{d z^{n}} \psi(z)=\frac{d^{n+1}}{d z^{n+1}} \ln \Gamma(z) \\
\psi^{(n)}(z)= & (-1)^{n+1} n!\sum_{k=0}^{\infty} \frac{1}{(z+k)^{n+1}} \\
\psi^{(n)}(1)= & (-1)^{n+1} n!\zeta_{n+1} \\
\psi^{(n)}\left(\frac{1}{2}\right)= & \left(2^{n+1}-1\right) \psi^{(n)}(1) \\
\psi^{(m)}(n+1)= & (-1)^{m} m!\left[-\zeta_{m+1}+1+\right. \\
& \left.+\frac{1}{2^{m+1}}+\ldots+\frac{1}{n^{m+1}}\right]= \\
= & \psi^{(m)}(n)+(-1)^{m} m!\frac{1}{n^{m+1}}
\end{aligned}
$$

See further section 42.4 on page 158.

## Bernoulli Numbers

See table 4 on page 174 .

## Riemann's Zeta-function

|  | exact | approx. |  |  |
| :--- | :--- | :--- | :--- | :---: |
| $\zeta_{0}$ | -1 |  |  |  |
| $\zeta_{1}$ | $\infty$ |  |  |  |
| $\zeta_{2}$ | $\pi^{2} / 6$ | 1.64493406684822643647 |  |  |
| $\zeta_{3}$ |  | 1.20205690315959428540 |  |  |
| $\zeta_{4}$ | $\pi^{4} / 90$ | 1.08232323371113819152 |  |  |
| $\zeta_{5}$ |  | 1.03692775514336992633 |  |  |
| $\zeta_{6}$ | $\pi^{6} / 945$ | 1.01734306198444913971 |  |  |
| $\zeta_{7}$ |  | 1.00834927738192282684 |  |  |
| $\zeta_{8}$ | $\pi^{8} / 9450$ | 1.00407735619794433938 |  |  |
| $\zeta_{9}$ |  | 1.00200839282608221442 |  |  |
| $\zeta_{10}$ | $\pi^{10} / 93555$ | 1.00099457512781808534 |  |  |

$$
\begin{aligned}
\zeta_{n} & =\sum_{k=1}^{\infty} \frac{1}{k^{n}} \\
\zeta_{2 n} & =\frac{2^{2 n-1} \pi^{2 n}\left|B_{2 n}\right|}{(2 n)!}
\end{aligned}
$$

See also page 59 and for details reference [31].

## Sum of Powers

In many calculations involving discrete distributions sums of powers are needed. A general formula for this is given by

$$
\sum_{k=1}^{n} k^{i}=\sum_{j=0}^{i}(-1)^{j} B_{j}\binom{i}{j} \frac{n^{i-j+1}}{i-j+1}
$$

where $B_{j}$ denotes the Bernoulli numbers (see page 174). More specifically

$$
\sum_{\substack{k=1 \\ n}}^{n}=n(n+1) / 2
$$

$$
\sum_{k=1}^{n} k^{2}=n(n+1)(2 n+1) / 6
$$

$$
\sum_{k=1}^{n} k^{3}=n^{2}(n+1)^{2} / 4=\left(\sum_{k=1}^{n} k\right)^{2}
$$

$$
\sum_{k=1}^{n} k^{4}=n(n+1)(2 n+1)\left(3 n^{2}+3 n-1\right) / 30
$$

$$
\sum_{k=1}^{n} k^{5}=n^{2}(n+1)^{2}\left(2 n^{2}+2 n-1\right) / 12
$$

$\sum_{k=1}^{n} k^{6}=n(n+1)(2 n+1)\left(3 n^{4}+6 n^{3}-3 n+1\right) / 42$

## ERRATA et ADDENDA

Errors in this report are corrected as they are found but for those who have printed an early version of the hand-book we list here errata. These are thus already obsolete in this copy. Minor errors in language etc are not listed. Note, however, that a few additions (subsections and tables) have been made to the original report (see below).

- Contents part now having roman page numbers thus shifting arabic page numbers for the main text.
- A new section 6.2 on conditional probability density for binormal distribution has been added after the first edition
- Section 42.6, formula, line $2, \nu$ changed into $\lambda$ giving

$$
f(x ; \mu, \lambda)=\frac{\lambda}{2} e^{-\lambda|x-\mu|}
$$

- Section 10.3, formula 2, line 4 has been corrected

$$
\phi_{x}(t)=E\left(e^{\imath t x}\right)=e^{\imath t \mu} E\left(e^{\imath t(x-\mu)}\right)=e^{\imath t \mu} \phi_{x-\mu}(t)=e^{\imath t \mu} \frac{\lambda^{2}}{\lambda^{2}+t^{2}}
$$

- Section 14.4, formula, line 2 changed to

$$
\phi(t)=E\left(e^{\imath t x}\right)=\frac{1}{\alpha} \int_{0}^{\infty} e^{\left(\imath t-\frac{1}{\alpha}\right) x} d x=\frac{1}{1-\imath t \alpha}
$$

- Section 18.1, figure 14 was erroneous in early editions and should look as is now shown in figure 73.
- Section 27.2, line 12: change $\nu r_{i}$ to $\nu p_{i}$.
- Section 27.6 on significance levels for the multinomial distribution has been added after the first edition.
- Section 27.7 on the case with equal group probabilities for a multinomial distribution has been added after the first edition.
- A small paragraph added to section 28.1 introducing the multinormal distribution.
- A new section 28.2 on conditional probability density for the multinormal distribution has been added after the first edition.
- Section 36.4, first formula, line 5, should read:

$$
P(r)=\sum_{k=0}^{r} \frac{\mu^{k} e^{-\mu}}{k!}=1-P(r+1, \mu)
$$

- Section 36.4, second formula, line 9, should read:

$$
P(r)=\sum_{k=0}^{r} \frac{\mu^{k} e^{-\mu}}{k!}=1-\int_{0}^{2 \mu} f(x ; \nu=2 r+2) d x
$$

- and in the next line it should read $f(x ; \nu=2 r+2)$.
- Section 42.5.2, formula 3, line 6, should read:

$$
\Gamma(z)=\alpha^{z} \int_{0}^{\infty} t^{z-1} e^{-\alpha t} d t \quad \text { for } \quad \operatorname{Re}(z)>0, \quad \operatorname{Re}(\alpha)>0
$$

- Section 42.6, line 6: a reference to table 9 has been added (cf below).
- Table 9 on page 179, on the Beta function $B(m, n)$ for integer and half-integer arguments, has been added after the first version of the paper.

These were, mostly minor, changes up to the 18th of March 1998 in order of apperance. In October 1998 the first somewhat larger revision was made:

- Some text concerning the coefficient of kurtosis added in section 2.2.
- Figure 6 for the chi-square distribution corrected for a normalization error for the $n=10$ curve.
- Added figure 8 for the chi distribution on page 44.
- Added section 11 for the doubly non-central $F$-distribution and section 12 for the doubly non-central $t$-distribution.
- Added figure 12 for the $F$-distribution on page 61.
- Added section 30 on the non-central Beta-distribution on page 108.
- For the non-central chi-square distribution we have added figure 22 and subsections 31.4 and 31.6 for the cumulative distribution and random number generation, respectively.
- For the non-central $F$-distribution figure 23 has been added on page 113. Errors in the formulæ for $f\left(F^{\prime} ; m, n, \lambda\right)$ in the introduction and $z_{1}$ in the section on approximations have been corrected. Subsections 32.2 on moments, 32.3 for the cumulative distribution, and 32.5 for random number generation have been added.
- For the non-central $t$-distribution figure 24 has been added on page 116, some text altered in the first subsection, and an error corrected in the denominator of the approximation formula in subsection 33.5. Subsections 33.2 on the derivation of the distribution, 33.3 on its moments, 33.4 on the cumulative distribution, and 33.6 on random number generation have been added.
- A new subsection 34.8.9 has been added on yet another method, using a ratio between two uniform deviates, to achieve standard normal random numbers. With this change three new references $[38,39,40]$ were introduced.
- A comparison of the efficiency for different algorithms to obtain standard normal random numbers have been introduced as subsection 34.8.10.
- Added a comment on factorial moments and cumulants for a Poisson distribution in section 36.2.
- This list of "Errata et Addenda" for past versions of the hand-book has been added on page 182 and onwards.
- Table 2 on page 172 and table 3 on page 173 for extreme significance levels of the chi-square distribution have been added thus shifting the numbers of several other tables. This also slightly affected the text in section 8.10.
- The Bernoulli numbers used in section 15.4 now follow the same convention used e.g. in section 42.3. This change also affected the formula for $\kappa_{2 n}$ in section 23.4. Table 4 on page 174 on Bernoulli numbers was introduced at the same time shifting the numbers of several other tables.
- A list of some mathematical constants which are useful in statistical calculations have been introduced on page 180 .

Minor changes afterwards include:

- Added a "proof" for the formula for algebraic moments of the log-normal distribution in section 24.2 and added a section for the cumulative distribution as section 24.3.
- Added formula also for $c<0$ for $F(x)$ of a Generalized Gamma distribution in section 18.2.
- Corrected bug in first formula in section 6.6.
- Replaced table for multinormal confidence levels on page 100 with a more precise one based on an analytical formula.
- New section on sums of powers on page 181.
- The illustration for the log-normal distribution in Figure 16 in section 24 was wrong and has been replaced.


## References

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[^0]:    ${ }^{1}$ From greek kyrtosis $=$ curvature from kyrt(ós) $=$ curved, arched, round, swelling, bulging. Sometimes, especially in older literature, $\gamma_{2}$ is called the coefficient of excess.

[^1]:    ${ }^{2}$ This is possible only if we require random numbers from one and the same binomial distribution with fixed values of $N$ and $p$.

[^2]:    ${ }^{3}$ If only $b$ is an integer use the relation $I_{x}(a, b)=1-I_{1-x}(b, a)$.

[^3]:    ${ }^{4}$ When integrated from zero to $x$ the incomplete gamma function is often denoted by $\gamma(a, x)$ while for the complement, integrated from $x$ to infinity, it is denoted $\Gamma(a, x)$. Sometimes the ratio $P(a, x)=\gamma(a, x) / \Gamma(a)$ is called the incomplete Gamma function.

[^4]:    ${ }^{5}$ This limit is at $N=k=2$ for the $50 \%$-curve, 3.92659 for $90 \%, 6.47061$ for $99 \%$ and 8.93077 for $99.9 \%$

[^5]:    ${ }^{6}$ Generating random numbers from a Poisson distribution with mean $\mu=-k \ln p$ branching to a Logarithmic distribution with parameter $p$ will give a Negative Binomial distribution with parameters $k$ and p.
    ${ }^{7}$ Taking a Poisson distribution with a mean distributed according to a Gamma distribution with parameters $a=k / \bar{n}$ and $b=k$.

[^6]:    ${ }^{8}$ The timing was done on a Digital Personal Workstation 433 au workstation running Unix version 4.0D and all methods were programmed in standard Fortran as functions giving one random number at each call.

[^7]:    ${ }^{9}$ See e.g. the discussion in the description of the moments for the Cauchy distribution which is the special case where $m=n=1$.

[^8]:    ${ }^{10}$ If $y$ is a normal theory estimate of $\sigma^{2}$ with $k$ degrees of freedom then $k y / \sigma^{2}$ is distributed according to the chi-square distribution with $k$ degrees of freedom.

[^9]:    ${ }^{11}$ Sometimes the more specific notation tri-, tetra-, penta- and hexagamma functions are used for $\psi^{\prime}$, $\psi^{\prime \prime}, \psi^{(3)}$ and $\psi^{(4)}$, respectively.
    ${ }^{12}$ For this calculation we need a few more Bernoulli numbers not given on page 158 above namely $B_{22}=\frac{854513}{138}, B_{24}=-\frac{236364091}{2730}, B_{26}=\frac{8553103}{6}$, and $B_{28}=-\frac{23749461029}{870}$

[^10]:    ${ }^{13}$ Without loss of generality it is enough to regard the standard normal density.

[^11]:    ${ }^{14}$ The notation ${ }_{2} F_{1}$ indicates the presence of two Pochhammer symbols in the numerator and one in the denominator.

